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Obituary

PERO M. DJURKOVIĆ

(1908–1981)

On 5 January 1981 died in Belgrade Pero M. Djurković, the retired research counsellor of the Belgrade Observatory and its former director.

Pero Djurković was born in Srpska Trnova (SR Bosna and Herzegovina) in 1908. He graduated from high school in Bijeljina and received in 1932 his degrees from the Belgrade Faculty of Philosophy – Department of Mathematics.



Pero Djurković joined the Belgrade Observatory late in 1932. From 1932 to 1936, jointly with his colleagues, he was occupied with the computing of various ephemerides, and sharing duties connected with the Time Service. In the period until the war he devoted himself to the minor planet identification, their observation, calculation of their orbits, as well as to the observation of comets – a speciality he mastered during his nine months sojourn at the Uccle Observatory during 1935/36.

In 1941 he was taken prisoner and remained in captivity in Germany until the liberation by the Allied armies in 1945.

Back in Belgrade in 1945 Djurković joined very actively in the restoring of the Observatory, extensively damaged in the liberation fightings in 1944. Not until 1947 was he, therefore, able to resume the astrono-

mical work proper. In that year, namely, Djurković, co-njointly with B. Ševarlić and Z. Brkić, organized the Latitude Service of the Belgrade Observatory. At the same time (1945–1950) he took an active part in the observations with both Transit Instrument and the Zenith-telescope.

But his main merit for the Belgrade Observatory is connected with the organization, during 1950–52, and the management until his retirement in 1971, of the Double Star Division. The instrument this Division is footing upon, is Zeiss equatorial, with an aperture of 65 cm. and the focal length 1055 cm. Djurković made with equal devotion the observations and the treatment of purely theoretical problems related to the double stars.

Yet, even in this period of his engagement in his main subject – double stars – he found time for other activities of a wider astronomical interest, demonstrating here again his high qualities of a scientific organizer. It was under his direction that regular observations of the solar phenomena were carried out under the IGY project from 1957 to 1959. The Publication No 11 of the Belgrade Observatory with over three hundred pages, containing observational data systematically collected by him during this memorable international undertaking, is the testimony of both Djurković's professional and working capacity.

Let us make a summing-up his fruitful work on the double star visual observations. In a period of 20 years he made over 3200 measurements. We owe him also the discovery of 10 pairs. His pair DJU 3 = 13 Vulpeculae is a real curiosity, its distance being 0.8 and the components' magnitudes 4.6 and 7.8 respectively. The pair met with interest among the observers and is presently systematically measured. Djurković's preference was measuring the most difficult pairs, conscious of their inherent interest and of their being most in demand. This is why he is ranked among the prominent double star observers.

Djurković is the author of over 40 scientific and professional works. He was a member of the IAU Commission 26 (Double Stars).

He took an active part in the Yugoslav National Astronomical Meetings and Conferences. His participation is also to be mentioned in the international meetings: in the Astro-Soviet consultation on the solar research in 1956 in Tbilisi (SSSR); 5th Session of the MGY Committee in 1958 in Moscow, where he headed the Yugoslav MGY national delegation, and several IAU General Assemblies.

Djurković considered it his moral obligation to contribute as much as possible to the popularization of astronomy in Yugoslavia. In this respect he may be regarded as one of the co-founders of the Belgrade's People's Observatory, dedicated to the observations by the amateur astronomers. Likewise he was one of those who

lent an active support to the initiative of building a sea-side observatory in the Adriatic island Hvar. He served several terms as member of the Yugoslav National Committee. He was also noted for his socio-political occupation. From 1965 to 1971 he was entrusted with the directorship of the observatory.

Remarkable abilities of organization of the research work as well as worthy results achieved in all the fields of his work and interest, in particular great enthusiasm for observation of Pero Djurković will serve a long time as an example to his younger colleagues, with whom he heartily cooperated. With his passing away the Belgrade Observatory lost a collaborator who devotedly served its interests and development for over 40 years.

BULLETIN

DE

L'OBSERVATOIRE ASTRONOMIQUE DE BELGRADE

N^o 132

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PRELIMINARY RESULTS OF THE FLEXURE INVESTIGATION OF THE BELGRADE LARGE VERTICAL CIRCLE IN THE PERIOD 1976–1979

M. Mijatov and DJ. Bozhichkovich

(Received May 3, 1981)

SUMMARY: An account is given of the preliminary results of the flexure determinations of the Large Vertical Circle of the Belgrade Observatory, carried out regularly under the observing program of bright stars in the period 1976–1979. A satisfactory accuracy of the measurements is stated, except for the year 1976, namely $\pm 0.^{\circ}20$. Of particular importance is the finding that the flexure of our LVC keeps its properties over longer time intervals. More intense investigations of this parameter is recommended.

1. Upon completing, late in 1976, the reconstruction of the Large Vertical Circle of the Belgrade Observatory, carried out under the direction of D.S. Usanov (Usanov et al., 1978), we took up, early in 1976, the observation by absolute method of a program of bright stars within the zone $+65^{\circ}$ to $+90^{\circ}$ declination (Teleki, Mijatov, 1976). Alongside with these observations, regular flexure determinations of the instrument's tube were made as required by absolute method. An account is herein presented of the preliminary results of the flexure determinations executed in the period 1976–1979.

2. The flexure of the LVC was investigated according to Bessel method using a pair of collimators (80/1000 mm.) installed horizontally in the east-west direction. The measurements of the horizontal flexure component b were accomplished either by two observers, whereby one of them had first to secure the alignment

of the collimators, after which he made the settings of the movable wire of the main instrument on the cross-wires of the collimators, or, failing two, one single observer carried out all the requisite operations.

Before starting the measurements the indoor and outdoor temperatures, as well as humidity, were read off. Besides, note was taken on the atmospheric conditions (clearness, wind).

The horizontal flexure component was determined on almost each observing night. In addition, several nights were entirely dedicated to the flexure determination in order to establish the variation of this parameter with the time. Thus there were in the period from March 1976 to October 1979 altogether 229 series of flexure determination.

3. Table 1 presents the number of flexure determinations according to observers.

Table I

Year	Observers						Total
	MM, DJB	DJB, MM	MM, BK	MM, MD	DJT, BK	BK	
1976	-	-	7	16	9	3	2 37
1977	82	-	1	-	-	-	83
1978	25	30	-	-	-	-	55
1979	28	24	2	-	-	-	54
Total	135	54	10	16	9	3	2 229

Observers: MM-M.Mijatov, DJB-Dj.Božičković,
BK-B.Kubiceša, MD-M.Dačić i
DJT-Dj.Teleki

As apparent from the above Table, the year 1976 is characteristic by the fewness of observations, carried out by a great number of observers. This fact is reflected in lower accuracy of the results from this period.

Table 2 gives the number of flexure determinations according to seasons.

Table II

Year	I-III	IV-VI	VII-IX	X-XII	Total
1976	-	24	13	-	37
1977	21	34	26	2	83
1978	3	5	36	11	55
1979	13	24	17	-	54
Total	37	87	92	13	229

The greatest number of determinations was, quite naturally, effected in the spring-sommer period.

4. The accidental error of a single determination of ϵ_b the horizontal flexure component was derived in two ways: from the difference of two consecutively determined values in the course of a series of measurements of $(b_{ew} - b_{we})$ and from several successive series, executed in the course of those nights which were dedicated solely to the flexure investigation. Table 3 presents the mean systematic difference $b_{ew} - b_{we}$ and ϵ_b .

ϵ_b was calculated according to formula:

$$\epsilon_b = 0.625 \frac{\sum |\Delta|}{n} \quad (1)$$

where: Δ – the values $(b_{ew} - b_{we})$ freed from the systematic difference $(b_{ew} - b_{we})$ from Table 3, and n the number of differences Δ (equal to the number of series of measurements)

Table III

Year	$b_{ew} - b_{we}$	ϵ_b	n
1976	+0.084	± 0.387	37
1977	+0.147	± 0.245	83
1978	+0.096	± 0.200	55
1979	+0.094	± 0.163	54

As evident from Table 3, the measurements performed in 1976 are of an inadequate accuracy, as already stated in Section 3. As for the remaining years, the accuracy was on the level characteristic of the flexure determination by Bessel method with this type of instruments. It was increasing from year to year due, probably, to the growing experience of observers. For the whole period of measurements we have: $\epsilon_b = \pm 0.^{\circ}24$, but if the the year 1976 is omitted then $\epsilon_b = \pm 0.^{\circ}20$.

The accuracy of the successive series, executed in the course of a single night is $\beta_b = \pm 0.^{\circ}20$, which is in harmony with the result obtained by the formula (1)

It has been found from earlier determinations (Mijatov, 1971–1972) that the principal systematic influences on the flecure determination by Bessel method had their origin in the unsteadiness of the collimators and in the local refraction. The systematic $(b_{ew} - b_{we})$ in Table 3 of the order $0.^{\circ}1$ indicate that there must have been some shifting of the collimators. It has, in addition, been stated that the effects of the local refraction have at times been very strong.

5. Preliminary data analysis has already that b was dependent on the temperature and, possibly, on the humidity. In order to scrutinize these effects we used the linear relation:

$$b = b_0 + a(t - 15^{\circ}\text{C}) + c(v - 70\%) \quad (2)$$

The results obtained hence by the least square method are given in Table 4.

Table IV

Year	b_0	a	c	ϵ_b	K	n
1976	+0.710 ± 137	+0.003 ± 32	+0.002 ± 7	+0.40 ± 0.47	0.08	23
1977	+0.726 ± 97	+0.089 ± 8	+0.002 ± 4	+0.47 ± 0.47	0.79	83
1978	+0.851 ± 206	+0.018 ± 9	-0.011 ± 5	+0.33 ± 0.33	0.49	47
1979	+0.719 ± 71	+0.042 ± 5	-0.005 ± 4	+0.28 ± 0.28	0.77	54

The coefficients of a and c in (2) for the year 1976 are uncertain, and the correlation coefficient k is too small. Thus, one cannot judge on the existence of the dependence of b on the temperature and the humidity on the basis of data from 1976. This is a consequence as earlier stated, of an insufficient accuracy of measurements carried out in that year. In the remaining years a distinct influence, that of the temperature in particular and even with a high correlation coefficient, must be stated.

The residuals of the relation (3) do not exhibit any systematical character, hence it can be claimed that no separate seasonal effect, other than the one dependent on temperature, is in existence.

It is of particular importance to point out that b_0 has remained virtually unchanged during the whole period of investigations, which indicates that the flexure of LVC, under the same temperature conditions, does not undergo any significant changes over longer time intervals. This has a particular weight with the absolute declination determinations.

An appreciable variation of b , of the order $0^{\circ}4$, during the observing night, has been stated. However, this variation has been determined from the measurements at the beginning and of the observing tour. Bearing in mind that the measurements at the beginning of the observing tour (early in the evening) are rather uncertain, there is no sufficient foundation to treat this variation as an established fact. This view is supported by

the circumstance that in the course of several nights, which were dedicated exclusively to the flexure determination, no such variation was perceived. But these measurements were never started early in the evening.

6. It follows from the preliminary flexure investigations of LVC that its horizontal component is, in more recent time, being determined with a satisfactory accuracy. To maintain this precision it is necessary that the measurements are executed by experienced observers. It is also necessary to investigate more closely the eventual systematic effects on the determination of b , on those accomplished in the evening in particular. Considerable temperature effects on the values of b have been stated, but in order to determine it definitely it is necessary to have the measurements more evenly distributed over the year.

Applied in the reductions, the values of b improve the accuracy of the declinations observed. We are confident that the flexure determinations of LVC will contribute to obtaining the most accurate possible declinations of the stars comprised by our observing program.

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RESULTS OF COMPARISON OF THE BELGRADE AND THE BORDEAUX CATALOGS OF THE NORTHERN PZT STARS

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and B. Ševarlić

(Received may 13, 1981)

SUMMARY: Results are presented of comparison the NPZT star catalogs compiled in Belgrade and Bordeaux, systematic differences between them derived and analysed with respect to the right ascension, declination, apparent magnitudes and spectral classes.

INTRODUCTION

The proposal of H. Jasuda (1971) of the Tokyo Observatory to make observation of all stars of the NPZT program with the meridian instruments situated in the northern hemisphere, was adopted by the Commission 8 at the 14th IAU General Assembly in Brighton in 1970 and recommended to all the observatories in possession of such instruments. The undertaking was aimed at bringing into accord catalogs, obtained independently by 14 photographic zenith tubes, or more

precisely, to reduce them to the FK4 system. The list, proposed by Yasuda, comprehended 1719 stars, distributed over declinations from $+10^\circ$ to $+60^\circ$, with magnitudes ranging from 5.6^m to 10.2^m . The results of the observations, once completed, should appear in the form of a „Derivation Catalog of the Northern PZT Stars.“

In the implementation of this program took part, among others, the observatories in Belgrade (BGD, $\varphi = 44^\circ 48'$; $\lambda = -1^\circ 22^m$) and in Bordeaux (BRD, $\varphi =$

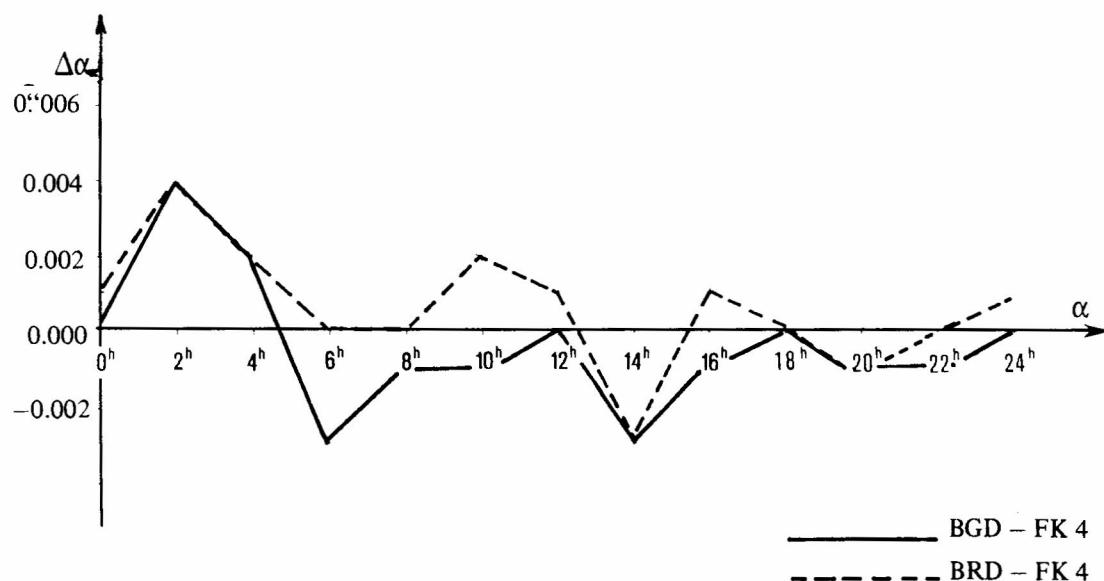


Fig. 1

$= 44^{\circ}50'$, $\lambda = -0^{\text{h}}02^{\text{m}}$). Both observatories compiled catalogs, listing stars positions (α, δ) determined by relative method in the FK4 system. Each one of the stars, appearing in these catalogs, was observed 6 times on the average. These are, for the time being, the only published catalogs of this series. At the Belgrade observatory the observations were carried out visually in the classical way, using classical technics, whereas in Bordeaux the „tracking“ photoelectric micrometer was applied in observing right ascensions, but the declinations were observed visually in the standard manner. The accuracy of a single observation of the right ascension amounts to $\epsilon_{\alpha} \cos \delta = \pm 0^{\circ}032$ in Belgrade and $\epsilon_{\alpha} \cos \delta = \pm 0.0086$ in Bordeaux. For declinations the respective values are $\epsilon_{\delta} = \pm 0^{\circ}26$ and $\epsilon_{\delta} = \pm 0^{\circ}203$. Mean epoch of observation in Belgrade is 1977.2 for the right ascensions and 1976.8 for the declinations (Sadžakov et al., 1981). The corresponding values related to the Bordeaux observations are 1973.7 and 1973.8 (Mazurier et al. 1977).

Prior to our entering into the work of forming the differences of coordinates as given by the two catalogues, and investigating their characteristics and origin, we derived the differences in coordinates of the reference stars from the BGD and the FK4 catalogs, as well as differences of the BRD and the FK4 catalogs. The differences in right ascension are illustrated in Fig. 1. The two curves in Fig. 1. display a good accordance by their general course. It is evident also that the amount of the differences in α are not negligible. In our view

these differences conceal the variations of the systematic errors of the fundamental system itself. No such differences in declination were stated.

Mean differences of coordinates in the two catalogues and the determination of systematic effects in them.

By making use of the proper motions from the catalogue of 1719 stars (Yasuda, 1971), star positions were reduced to the epoch of the Belgrade catalogue. Thereupon the differences were formed

$$\Delta\alpha = \alpha_{\text{BGD}} - \alpha_{\text{BRD}}$$

$$\Delta\delta = \delta_{\text{BGD}} - \delta_{\text{BRD}}$$

The differences exceeding 3ϵ which make about 3%, were discarded. (Tables 1 and 2, where ϵ_0 is the error in the derived differences and n the number of stars used for deriving this error).

As is well known, various systematic effects are comprised by these differences. These influences may be represented in the form:

$$\Delta\alpha = \Delta\alpha_0 + \Delta\alpha_{\delta} + \Delta\alpha_{\alpha} + \Delta\alpha_m + \Delta\alpha_{sp}$$

$$\Delta\delta = \Delta\delta_0 + \Delta\delta_{\delta} + \Delta\delta_{\alpha} + \Delta\delta_m + \Delta\delta_{sp}$$

in 0.0001 units

Table I

$\alpha \backslash \delta$	$10^{\circ}-20^{\circ}$	$35^{\circ}-38^{\circ}$	$39^{\circ}-45^{\circ}$	$47^{\circ}-52^{\circ}$	$53^{\circ}-59^{\circ}$	$10^{\circ}-59^{\circ}$
	$\Delta \epsilon_0$	n	$\Delta \epsilon_0$	n	$\Delta \epsilon_0$	n
$0^{\text{h}} - 2^{\text{h}}$	-65	71 17	-100	45 22	20	39 23
$2^{\text{h}} - 4^{\text{h}}$	-41	38 25	22	54 23	-14	24 20
$4^{\text{h}} - 6^{\text{h}}$	21	44 22	-101	43 19	-30	39 22
$6^{\text{h}} - 8^{\text{h}}$	-66	48 23	-37	48 23	-74	48 27
$8^{\text{h}} - 10^{\text{h}}$	-28	57 20	-82	43 19	-11	46 23
$10^{\text{h}} - 12^{\text{h}}$	28	47 18	-12	47 18	-48	35 19
$12^{\text{h}} - 14^{\text{h}}$	-47	39 19	-47	34 20	-21	42 20
$14^{\text{h}} - 16^{\text{h}}$	-76	45 24	38	39 18	119	36 19
$16^{\text{h}} - 18^{\text{h}}$	3	42 18	181	44 19	109	35 26
$18^{\text{h}} - 20^{\text{h}}$	-90	48 29	181	41 20	158	31 26
$20^{\text{h}} - 22^{\text{h}}$	8	23 29	101	46 23	197	39 29
$22^{\text{h}} - 24^{\text{h}}$	-74	35 19	68	43 20	56	50 27
$\Delta\alpha_0 + \Delta\alpha_{\delta}$	-37	13 263	17	14 244	44	13 281
					-84	9 388
					-61	9 390
					-31	5 1566

in $0^{\text{s}}001$ units

Table II

$\delta \backslash \alpha$	10° - 25°	35° - 38°	39° - 45°	47° - 52°	53° - 59°	10° - 59°				
	$\Delta \epsilon_0$	n	$\Delta \epsilon_0$	n	$\Delta \epsilon_0$	n	$\Delta \delta_\alpha$	ϵ_0	n	
0° - 2°	-74 +86	17	85 +73	20	-115 +81	20	-78 +56	38	19 +79	29
2° - 4°	242 46	21	-11 66	20	271 75	18	25 69	33	86 73	31
4° - 6°	183 63	21	63 67	16	-21 89	22	-50 69	28	44 58	30
6° - 8°	128 75	22	101 59	21	100 59	30	-26 63	31	132 51	28
8° - 10°	130 57	22	-99 56	18	197 68	24	-28 67	32	66 71	29
10° - 12°	-56 59	20	11 83	17	-44 55	19	-115 67	30	-31 66	29
12° - 14°	75 62	19	-7 68	21	1 71	20	72 67	23	-64 58	29
14° - 16°	-22 89	20	51 106	17	-36 78	19	-1 58	28	107 60	30
16° - 18°	-9 77	22	18 54	23	6 64	24	-60 68	28	168 65	31
18° - 20°	89 60	30	-10 75	22	86 71	23	-6 49	33	13 43	32
20° - 22°	126 59	27	-72 73	23	107 63	29	114 55	30	-42 49	32
22° - 24°	166 79	17	165 62	23	157 70	27	-61 54	33	-12 44	46
$\Delta \delta_0 + \Delta \delta_\alpha$	85 20	258	25 20	241	70 21	275	-21 18	367	38 17	376
									35 9	1517

where the subscripts denote the argument upon which the individual part of the systematic influence is dependent. At first we separated in the customary way the systematic influences $(\Delta \alpha_0 + \Delta \alpha_\delta)$ and $(\Delta \delta_0 + \Delta \delta_\delta)$, found in the last row of Tables 1. and 2. Next followed the calculation of $\Delta \alpha_\alpha$ and $\Delta \delta_\alpha$, given in the last column of the same Tables, as well as in Figs. 2 and 3.

As illustrated in Fig. 2, representing the values $\Delta \alpha_\alpha$, the inflection points of the curve fall within the intermediary seasons while its extremes take place during the summer (maximum) and the winter (minimum). As here the differences are in question of the seasonal

effects in Belgrade and Bordeaux, it is quite possible that they have their origin in the sharp seasonal changes at the former station. Concerning the systematic effects of the $\Delta \delta_\alpha$ type, their variation are confined within the $+0.^{\circ}1$ and $-0.^{\circ}1$ limits for the southern stars (the zone $+10^\circ$ to $+25^\circ$ declination) and between the $+0.^{\circ}05$ to $-0.^{\circ}05$ limits for the northern stars (the zone $+53^\circ$ to $+59^\circ$ declination). These values for other zones are negligible.

The systematic effects of the type $\Delta \alpha_\alpha$ and $\Delta \delta_\alpha$ have been submitted to the analysis by means of the Fourier polynomial of the form:

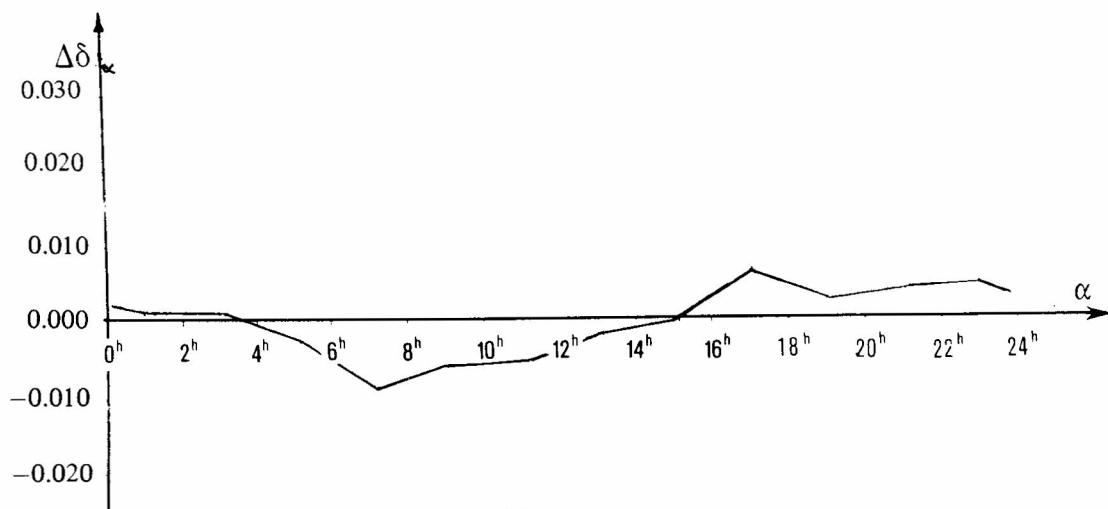


Fig. 2

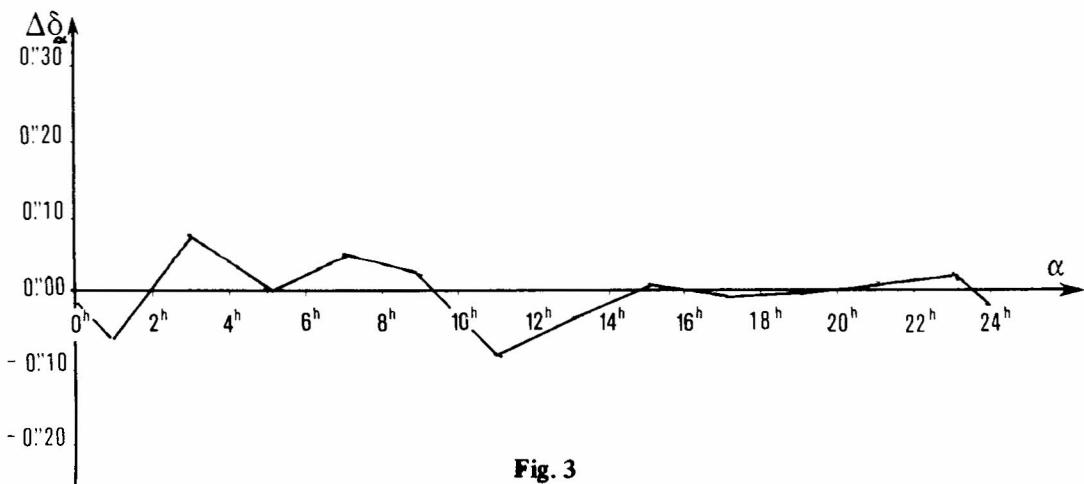


Fig. 3

$$\Delta\alpha_\alpha = A_0 + A_1 \sin\alpha + B_1 \cos\alpha + A_2 \sin 2\alpha + B_2 \cos 2\alpha$$

$$\Delta\delta_\alpha = B_0 + A'_1 \sin\alpha + B'_1 \cos\alpha + A'_2 \sin 2\alpha + B'_2 \cos 2\alpha$$

The coefficients $A_i, B_i; A'_i, B'_i$ were determined by the least squares method and are given in the Tables 3 and 4.

The Fourier polynomials, illustrated in Figs. 4 and 5, indicate the existence of both periodical and the seasonal character of the two systematic effects dealt with here.

Some interest must be attached to the existence of a certain correlation between the systematic differences $\Delta\alpha_\delta$ and $\Delta\delta_\delta$ with the derived Belgrade instrumental system (Daćić, 1980), whose parallel values are given in Table 5.

in 0.0001 units

Table III

δ	A_0	A_1	B_1	A_2	B_2	ϵ_0	n
10°-25°	-37 ± 13	14 ± 18	-12 ± 19	-9 ± 18	-9 ± 19	± 21	263
35°-38°	16 ± 13	-114 ± 18	2 ± 18	-8 ± 18	-51 ± 19	20	244
37°-45°	37 ± 12	-103 ± 16	36 ± 17	-18 ± 17	-18 ± 17	20	281
47°-52°	-85 ± 9	-6 ± 13	5 ± 13	4 ± 13	13 ± 13	18	388
53°-59°	-63 ± 9	-24 ± 13	24 ± 13	19 ± 13	42 ± 13	18	390
10°-59°	-33 ± 5	-4 ± 7	27 ± 7	8 ± 7	0 ± 7	± 20	1566

$$\Delta\alpha = A_0 + A_1 \sin\alpha + B_1 \cos\alpha + A_2 \sin 2\alpha + B_2 \cos 2\alpha$$

in 0.001 units

Table IV

δ	A'_0	A'_1	B'_1	A'_2	B'_2	ϵ'_0	n
10°-25°	81 ± 20	58 ± 27	59 ± 30	-18 ± 28	-41 ± 29	± 32	258
35°-38°	25 ± 20	6 ± 29	36 ± 29	15 ± 29	20 ± 29	32	241
37°-45°	57 ± 21	22 ± 29	23 ± 31	-26 ± 30	-4 ± 30	35	275
47°-52°	-18 ± 18	-22 ± 26	-3 ± 26	-6 ± 26	-15 ± 26	34	367
53°-59°	43 ± 17	10 ± 24	-8 ± 24	48 ± 24	-64 ± 24	33	376
10°-59°	35 ± 9	11 ± 12	16 ± 12	4 ± 12	-27 ± 12	33	1517

$$\Delta\delta = A'_0 + A'_1 \sin\alpha + B'_1 \cos\alpha + A'_2 \sin 2\alpha + B'_2 \cos 2\alpha$$

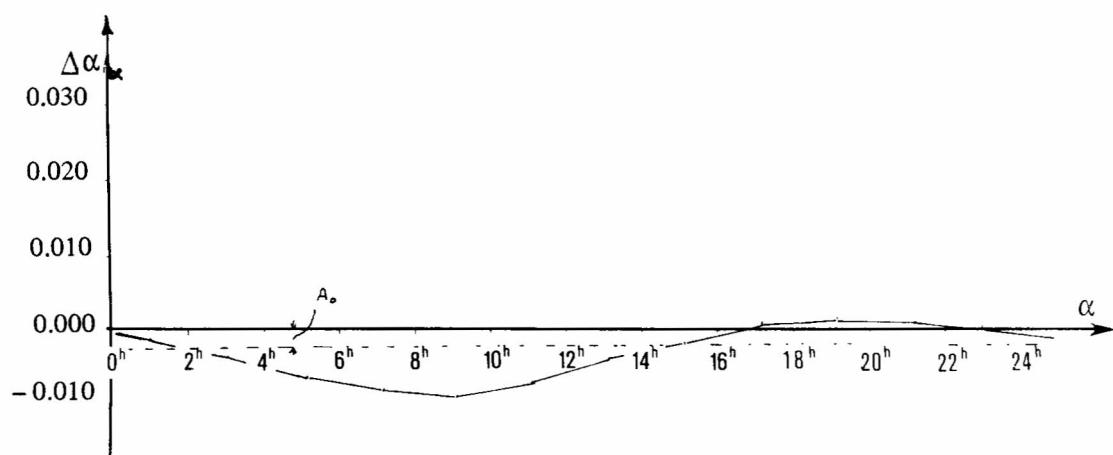


Fig. 4

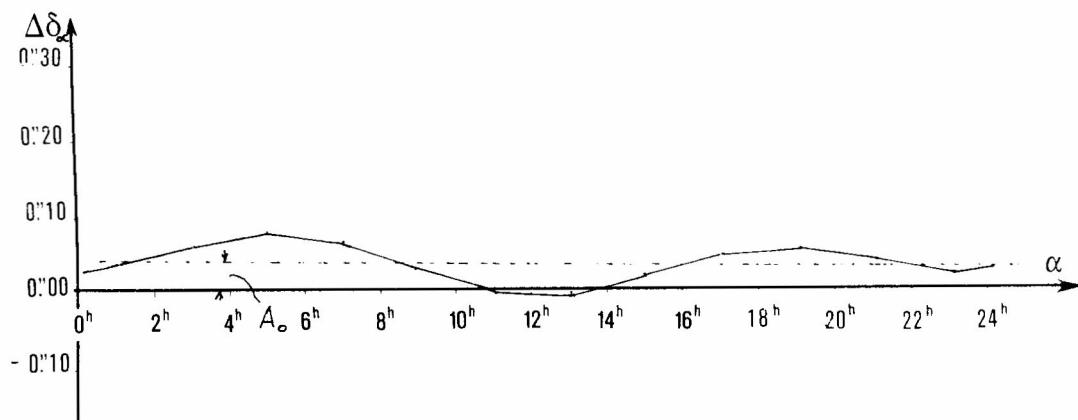


Fig. 5

Table V

Zone	$\Delta\alpha_{\delta}$	System of LMI in right ascension	$\Delta\delta_{\delta}$	System of LMI in declination
10° - 25°	-6	+45	+50	+45
35° - 38°	+48	+47	-10	-24
35° - 45°	+75	+40	+35	-61
47° - 52°	-53	-12	-56	-71
53° - 59°	-30	-20	+ 3	+ 5

The values in the second and the third columns are given in $0^{\circ} 0.001$ units, while those in the forth and the fifth columns in $0^{\circ} 0.001$ units. Had they been accounted for in the Belgrade catalogue the systematic effects of the $\Delta\alpha_{\delta}$ and $\Delta\delta_{\delta}$ types would evidently have been considerably smaller.

Systematic influences depending on the apparent magnitude and the spectral class.

Upon eliminating the systematic influences of type $(\Delta\alpha_0 + \Delta\alpha_{\delta})$, $(\Delta\delta_0 + \Delta\delta_{\delta})$ and $\Delta\alpha_{\alpha}$ and $\Delta\delta_{\alpha}$ from

the total values $\Delta\alpha$ and $\Delta\delta$ we formed Tables 6 and 7 giving residuals of the stated differences, arranged according to their apparent magnitudes (rows) and the spectral class (columns).

In the last row systematic differences are listed which are dependent on the apparent magnitude ($\Delta\alpha_m$, $\Delta\delta_m$), along with the corresponding values ϵ_0 and n , while the last column gives the systematic differences dependent on the spectral class ($\Delta\alpha_{sp}$, $\Delta\delta_{sp}$), with the corresponding values ϵ_0 and n . The values $\Delta\alpha_{sp}$, $\Delta\delta_{sp}$ are derived step by step after eliminating the quantities $\Delta\alpha_m$, $\Delta\delta_m$ from the original residuals.

The numerical values of the systematic influences $\Delta\alpha_m$, $\Delta\alpha_{sp}$, $\Delta\delta_m$, $\Delta\delta_{sp}$ appear convincing for $\Delta\delta_{sp}$ only, a pronounced variation with the spectral class from O through M being evident. It is hard to say whether the amounts of $\Delta\alpha_m$, $\Delta\delta_m$, $\Delta\alpha_{sp}$ are unambiguously dependent on the app. magnitude or the spectral class. This conclusion is confirmed by the Abbe's criterion. An earlier analysis of the systematic influences $\Delta\delta_{sp}$ (Sadža-

in $0.^s0001$ units

Table VI

$\Delta \alpha_{sp}$	$m \leq 6,5$	$6,5 - 7,5$	$7,5 - 8,5$	$8,5 < m$	$\Delta \alpha_{sp}$
	$\Delta \epsilon_o$	$\Delta \epsilon_o$	$\Delta \epsilon_o$	$\Delta \epsilon_o$	ϵ_o
0.8	-18 +39 24	51 +28 38	64 +26 46	21 +54 17	38 +17 125
A	29 28 46	10 18 103	-19 15 192	-14 22 82	-5 10 423
F	-45 45 24	-21 22 84	-36 17 138	25 24 65	-19 11 311
G	33 35 25	7 24 49	7 16 135	-5 20 71	9 11 280
K,M	-15 24 75	36 17 149	-18 15 177	12 37 26	3 10 427
$\Delta \alpha_m$	-3 14 194	16 9 423	-12 8 688	3 12 261	

in $0.^m001$ units

Table VII

$\Delta \alpha_{sp}$	$m \leq 6,5$	$6,5 - 7,5$	$7,5 - 8,5$	$8,5 < m$	$\Delta \alpha_{sp}$
	$\Delta \epsilon_o$	$\Delta \epsilon_o$	$\Delta \epsilon_o$	$\Delta \epsilon_o$	ϵ_o
0,B	133 +67 21	107 +47 38	-99 +43 44	-21 +78 17	17 +28 120
A	-19 45 47	33 33 102	14 24 186	0 36 78	12 16 413
F	-10 75 24	-8 35 85	10 29 132	45 49 61	11 20 302
G	24 58 23	10 57 44	-12 30 129	-16 38 68	-7 20 264
K,M	27 35 76	-66 27 144	-30 26 174	153 58 24	-17 16 418
$\Delta \alpha_m$	22 23 191	-6 16 413	-11 13 665	20 21 248	

kov et al. 1975) showed also that the blue stars (spectral classes O, B, A) are more shifted towards the zenith than the yellow and red ones.

The values obtained by applying the Abbe's criterion reveal the existence of the systematic influence of this type in the declination but not in the right ascension.

CONCLUSION

1. The existence is confirmed of the systematic differences $\Delta \alpha_\alpha$, $\Delta \alpha_\delta$, $\Delta \delta_\alpha$, $\Delta \delta_\delta$ between the Belgrade and the Bordeaux catalogues, as well as their periodicity. The likelihood is suggested of the greater part of the systematic differences $\Delta \delta_\alpha$, $\Delta \delta_\delta$ to have to be ascribed to the Belgrade instrument system. As far as the systematic differences $\Delta \alpha_\alpha$ and $\Delta \alpha_\delta$ are concerned, at least part of them may originate from the difference in the seasonal effects at the two stations, while a minor part from the epoch difference of the two catalogues.

2. The systematic errors of the $\Delta \delta_{sp}$ type are the largest for the blue stars, and their variation is distinctly systematic. The parameters obtained

by applying the Abbe's criterion furnish confirmation of the systematic character of this influence, dependent on the spectral class, as well as that there is not any convincing dependence on the apparent magnitude.

We take the opportunity to thank Dr. D. Teleki for useful suggestions he made in the course of preparation of the manuscript.

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ORBITAL ELEMENTS OF VISUAL BINARY SYSTEMS
 Σ 248 = ADS 1786 and Σ 535 = ADS 3174

G. M. Popović

(Received May 21, 1981)

SUMMARY: Presented are the orbital elements, ephemerides and some astrophysical data of the visual binary systems ADS 1786 and ADS 3174.

Binary system Σ 248 = ADS 1786 = IDS 02148N4219
 α, δ (1950): $2^{\text{h}}17^{\text{m}}9\text{s}$, $+42^{\circ}33'$, 9.6–9.6, Sp. –

From 1832 to 1975 the position angle of this W. Struve's system changed by 66° , with considerable changes in distance. These changes afforded the possibility of determining a provisional orbit of the system (Fig. 1). The orbital elements were derived through the Innes' geometrical elements A, B, F and G. The effect of the precession on the position angle has been neglected. The following data on this pair are also found in the ADS catalogue: proper motion $0.^{\text{d}}230$ in $177^\circ.3$ (Bgd EB), the dynamical parallax is $0.^{\text{d}}013$ (J. – F) or $0.^{\text{d}}027$ (R. – M). The masses of the system's components are computed by means of the empirical mass-luminosity relation for the main HR sequence (Popović–Angelov, 1970). The computed values of the absolute brightness indicate that the spectrum of both components might well be K3. The pair's orbit (Fig. 1) is computer-drawn (WANG 2200) by intermediary of the calculated ephemerides from 1800 to 2040 with the one year step. The observations are denoted in this Fig. by crosslets. The Tables I, II and III present: the observations of this pair with the corresponding residuals O – C, orbital elements, the astrophysical constants as well as the ephemerides of the system from 1975 to 2004. Combell's elements are published in „Circulaire d'Information“ No 81, Comm. 26, UAI, 1980.

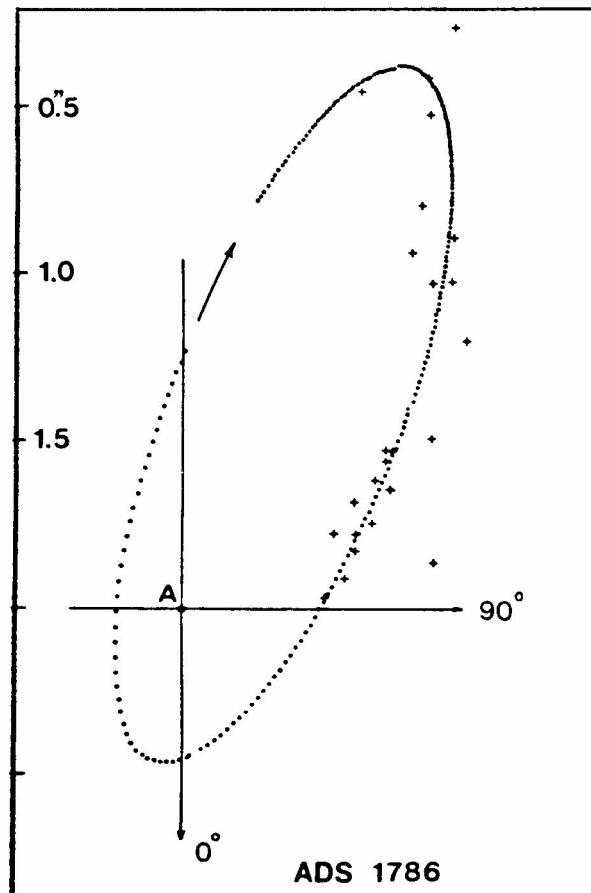


Fig. 1

Table I
Measurements of the pair ADS 1786 and 0-C

t	θ_t	ρ	n	Observer	0 - C
1832.13	161.0	1.64	4		-0.6 - 0.03
1869.99	155.3	1.76	4		+0.3 + .00
1879.61	155.1	1.93	1	Cin	+1.7 + .21
1894.50	153.8	1.65	3	HΣ	+3.2 + .02
1902.74	149.7	1.40	1	Hu	+0.9 - .16
1904.91	147.6	1.26	1	β	-0.8 - .28
1911.12	143.9	1.38	5	Dob2, Fox3	-3.0 - .09
1923.32	142.9	1.22	9	A2, Gr05, Gcb2	-0.7 - .10
1927.788	140.8	1.27	4	GΣ	-1.4 + .01
1944.64	133.6	1.16	3	Voûte	-1.6 - .17
1954.806	127.6	0.77	3	Djk	-0.9 - .03
1956.69	128.8	0.76	3	Cou	+1.9 - .01
1958.93	126.6	0.74	3	Cou	+1.8 + .01
1959.79	125.1	0.90	2	VBS	+1.1 + .18
1961.81	120.3	0.70	3	Bz	-1.5 + .02
1962.79	120.9	0.70	5	hz	+0.3 + .04
1962.874	122.8	0.59	4	Wor	+2.3 - .07
1963.00	124.7	0.69	3	Cou	+4.3 + .03
1967.445	114.5	0.61	3	DZ	+0.3 + .04
1967.728	117.1	0.49	1	Djk	+3.5 - .08
1969.10	113.2	0.56	4	hz	+1.9 + .02
1970.837	108.2	0.54	4	Wor	+0.2 + .03
1973.231	100.5	0.48	3	Wor	-2.2 + .01
1973.749	100.6	0.76	1	OLE	-0.9 + .30
1975.851	94.6	0.41	3	Wor	-1.3 - 0.02

Table II
Elements of the orbital motion and astrophysical quantities of the ADS 1786

P = 258 ^j .97	A = 0."936	$M_A = +6^m.47$
n = 1 ^o .39012	B = -0.479	$M_B = +6.47$
T = 2011.57	F = -0.601	
e = 0.626	G = -0.212	$m_A = 0.87 M_{\odot}$
a = 1."155	C = +0.478	$m_B = 0.87 M_{\odot}$
i = 111 ^o .40	H = +0.963	
$\omega = 206.41$	$t(\Omega) = 1824.20$	a = 48.1 A.U.
$\Omega = 163.16$	$t(\varpi) = 2008.07$	dp = 0."024

Table III
Ephemeris of the binary ADS 1786

1975.0	98 ^o .2	0."44	1990.0	38 ^o .5	0.35
1976.0	95.4	0.43	1991.0	34.2	0.35
1977.0	92.4	0.41	1992.0	30.1	0.36
1978.0	89.2	0.40	1993.0	26.2	0.37
1979.0	85.8	0.39	1994.0	22.4	0.38
1980.0	82.2	0.38	1995.0	18.8	0.39
1981.0	78.3	0.36	1996.0	15.4	0.40
1982.0	74.3	0.36	1997.0	12.2	0.41
1983.0	70.0	0.35	1998.0	9.2	0.42
1984.0	65.6	0.34	1999.0	6.2	0.43
1985.0	61.2	0.34	2000.0	3.5	0.44
1986.0	56.6	0.34	2001.0	0.8	0.45
1987.0	52.0	0.34	2002.0	358.2	0.46
1988.0	47.4	0.34	2003.2	355.7	0.46
1989.0	42.9	0.34	2004.0	353.2	0.46

Binary system Σ 535 = ADS 3174 = IDS 04178N1109
 α, δ (1950): 4^h20^m5^s, +11°16', 7^m1-8^m6, Sp_A = A2

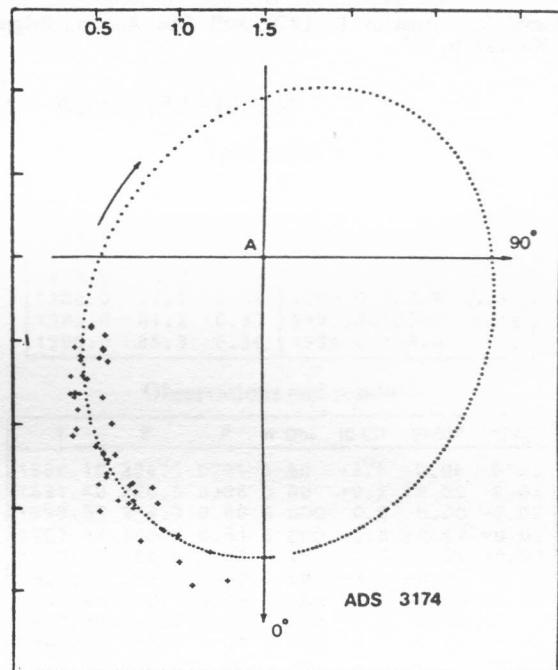
The position angle of this pair changed by 62° while ρ declined by approximately 0".9 in the period from 1931 to 1980. Its orbital motion is strongly pronounced (Fig. 2), affording the derivation of a preliminary orbit. Currently ρ exhibits a tendency of a slight declining but, according to the calculated preliminary elements, is not expected to descend lower than 0".84. Campbell's elements are derived from the Innes' geometrical elements A, B, F, G. The position angle is not corrected for the effect of precession. The components' masses, given in Table V are calculated by using the em-

pirical relation, characteristic of the stars lying below the main sequence of the HR diagram (Popović-Angelov, 1970). A better accordance of the observed spectrum and the calculated absolute brightness of the A component is thereby achieved than would be the case if A component were considered as belonging to the main sequence. The computed absolute magnitude $M_A = +3^m.1$ would entail the A7 spectrum for the sub-main sequence stars, which is not quite in accordance with the observed A2 spectrum.

Table IV
Measurements of the pair ADS 3174 and 0-C

t	θ_t	ρ	n	Observer	0 - C
1831.34	353.9	1.95	5	Σ	+3.3 + .16
1841.96	347.4	2.02	2	0Σ	-0.1 + .24
1843.10	349.8	1.79	2	Gsh	+2.6 + .01
1847.40	344.2	1.90	3	Da	-1.6 + .13
1862.99	342.4	1.74	5	Δ	+1.0 + .00
1870.01	339.6	1.74	2	Δ	+0.3 + .01
1878.76	336.4	1.60	1	Δ	-0.2 - .10
1887.19	336.4	1.86	1	Cel	+2.5 + .18
1888.14	334.5	1.52	3	Hl	+0.9 - .15
1891.01	332.4	1.70	19	Nis4, Spl5	-0.2 + .04
1891.42	331.9	1.65	3	$H\Sigma$	+2.4 - .01
1893.14	329.4	1.56	3	Maw	-2.5 - .10
1895.92	331.2	1.69	2	Collins	+0.2 + .04
1896.11	331.5	1.59	1	Cel	+0.6 - .06
1898.08	329.3	1.58	2	A	-1.0 - .06
1901.59	329.0	1.71	5	Lohl, Hul, Frederickl, Kqs02	-0.1 + .08
1902.41	323.6	1.68	7	Bowyer	-5.2 + .06
1902.9	324.5	1.61	15	Gr0	-4.1 - .01
1905.70	327.1	1.82	9	Frml,Doo8	-0.5 + .21
1910.74	324.0	1.59	28	Dob3,Has1, Lau2,Fox3, Wz3,Vou4, Gr012	-1.8 + .00
1914.17	322.7	1.52	18	Stol,VBs3, Dob2,Ph16, Rabe4,FBn2	-1.9 - .05
1918.15	323.1	1.47	9	Ptt3,Dob2, Ph14	+0.0 - .08
1921.41	320.6	1.52	19	Lv6,Chan5, Gr03,Guil, Nv11,Dob3	-1.2 - .01
1923.16	320.1	1.51	16	Lv4,Gcb2, B4,Kpz2, Gau2,Berm2	-1.1 - .01
1926.17	317.6	1.33	2	Fatou	-2.4 - .18
1929.86	318.3	1.52	4	Bz	-0.2 + .03
1934.90	312.7	1.57	4	Bz	-3.7 + .11
1936.20	311.0	1.26	2	Giacobini	-4.8 - .19
1942.70	312.5	1.53	3	Voûte	-0.4 + .12
1949.45	309.2	1.43	3	Prêtre	-0.6 + .06
1954.166	306.19	1.408	6	Rabe	-1.2 + .07
1956.178	305.36	1.390	4	Rabe	-1.0 + .06
1957.85	304.9	1.42	3	Wor	-0.6 + .10
1958.10	306.6	1.36	3	Gou	+1.2 + .05
1961.03	304.4	1.27	5	VBs	+0.6 - 0.02
1961.878	305.5	1.39	4	Bos	+2.2 + 0.10
1962.15	302.9	1.28	3	hz	-0.3 + .00
1964.12	297.9	1.24	3	Bz	-4.2 - .03

t	θ_t	ρ	n	Observer	0 - C
1964.948	299.5	1.25	5	Walker Jr.	-2.1 - .01
1965.976	304.1	1.11	1	Walker Jr.	+3.1 - .15
1965.998	301.3	1.15	2	DZ	+0.3 - .11
1966.058	299.9	1.10	1	Walker Jr.	-1.1 - .16
1966.117	299.1	1.10	2	Djk	-1.8 - .16
1966.153	296.0	1.04	1	GP	-4.9 - .22
1970.019	301.5	1.36	4	Wor	+2.8 + .13
1974.81	294.4	1.18	2	OLE	-1.3 + .01
1975.06	295.3	1.25	4	Muller	-0.2 + .06
1980.888	291.9	1.10	3	GP	+0.2 - .05
1980.892	291.5	1.09	1	DZ	-0.2 - 0.06

**Fig. 2**

The orbit (Fig. 2) is computer-drawn using ephemeris from 1790 to 2459 with a three years step. The observations from Table IV and here are marked by crosslets. The Tables IV, V and VI give the observations and the corresponding residuals O - C, orbital elements, astrophysical constants and ephemeris from 1981 to 2000. Cambell's elements are published in „Circulaire d'Information“ No 84, Comm. 26, UAI 1981.

The following data on this pair can also be found in the ADS catalogue: the dynamical parallax is 0.024 (J. - F.) or 0.021 (R. - M.); the proper motion is given as 0.06 in 180°.

Table V
**Elements of the orbital motion and astrophysical
 quantities of the ADS 3174**

P = 681 ^o .95	A = -1 ^m .2071	M _A = +3 ^m .10
n = 0 ^o .5279	B = -0.4785	M _B = +4.60
T = 2064.60	F = -0.7828	
e = 0.319	G = +1.1971	M _A = 1.02 M _⊕
a = 1 ^m .510	C = +0.7201	M _B = 0.84 M _⊕
i = 142 ^o .97	H = +0.5554	
ω = 307.64	tΩ = 2116.56	a = 94.4 A.U
Ω = 155.32	tΩ = 1886.16	dp = 0 ^m .016

Table VI
Ephemeris of the pair ADS 3174

1981.0	291 ^o .6	1 ^m .15	1991.0	284 ^o .2	1 ^m .08
1982.0	290.9	1.14	1992.0	283.4	1.07
1983.0	290.2	1.14	1993.0	282.6	1.06
1984.0	289.5	1.13	1994.0	281.8	1.06
1985.0	288.8	1.12	1995.0	281.0	1.05
1986.0	288.0	1.11	1996.0	280.2	1.04
1987.0	287.3	1.11	1997.0	279.3	1.04
1988.0	286.5	1.10	1998.0	278.5	1.03
1989.0	285.8	1.09	1999.0	277.6	1.02
1990.0	285.0	1.08	2000.0	276.7	1.01

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 No. 124, p. 147.

ORBITS OF FOUR VISUAL BINARIES

D. J. Zulević

(Received 20 June 1981)

SUMMARY: Orbit and dynamical parallaxes are presented for four visual binary systems: ADS 1371, ADS 7758, ADS 11989 and ADS 16877. Calculated positions are compared with observations and ephemerides are given for each system.

Orbits for four visual binaries have been computed using the methods of Thiele-Innes. Dynamical parallaxes were computed by the method of Baize and Ro-

mani (1946) with magnitudes spectral types taken from the Lick Index Catalogue of Visual Double Stars, 1961. o (1963). The relevant information is in Tables I, II, III. i IV.

Table I
ADS 1371 = IDS 01384N5637 = BU 453,10.1–10.4,65

P = 180.44		A = + 0"5333
n = 1.9951	$\Omega = 33^{\circ}9$	B = + 0.4733
T = 1791.53	$\omega = 13^{\circ}5$	F = - 0.3633
a = 0"73	$\pi_{\text{dyn}} = 0"021$	G = + 0.2360
e = 0.31	$\Sigma m_{AB} = 1.290$	C = + 0.1408
i = 55 ^o 7	a = 34.8 U.A.	H = + 0.5864

Ephemerides

T	P	ρ	T	P	ρ
1981.0	66.3	0"41	1987.0	89 ^o 7	0"36
1982.0	69.8	0.40	1988.0	94.3	0.35
1983.0	73.4	0.39	1989.0	99.0	0.34
1984.0	77.2	0.38	1990.0	103.8	0.34
1985.0	81.2	0.37	1991.0	108.7	0.34
1986.0	85.3	0.36	1992.0	113.6	0.34

Observations and residuals

T	P	ρ	n	Obs.	(O-C)P	(O-C)P	(O-C)P
1880.10	224.1	0"91	3	BU	+3.6	+0"06	-0"03
1891.60	228.5	0.86	3	BU	+0.2	+0.00	-0.03
1899.95	234.7	0.80	3	D00	0.0	0.00	-0.02
1903.73	233.2	0.81	3	D00	-4.8	-0.07	+0.03
1923.70	260.5	0.63	2	VBS	-3.9	-0.04	+0.09
1929.73	274.3	0.53	3	VBS	-3.2	-0.03	+0.06
1935.45	290.3	0.42	3	VBS	-4.5	-0.03	+0.00
1937.50	294.8	0.39	3	VBS	-7.0	-0.05	-0.01
1944.90	321.2	0.41	2	VBS	-7.7	-0.05	+0.02
1948.79	335.2	0.36	3	VBS	-7.5	-0.05	-0.05
1950.95	345.9	0.28	3	BAZ	-3.8	-0.03	-0.14
1951.8	(306.1)	(0.12)	2	VBS (352.7)	—	(0.42)	
1954.58	(312.1)	(0.11)	3	VBS (0.6)	—	(0.45)	
1958.03	322.3	0.20	1	B	+54.5	+0.47	-0.29
1960.74	17.9	0.40	4	WOR	+1.6	+0.01	-0.09
1966.873	35.4	0.43	4	WALK	+5.3	+0.05	-0.08
1974.0	52.9	0.44	2	ZUL	+6.5	+0.05	-0.04
1976.70	60.3	0.45	1	ZUL	+6.9	+0.05	-0.00
1978.29	58.4	0.48	2	ZUL	+0.5	+0.00	+0.04
1980.90	65.9	0.40	1	ZUL	-0.1	-0.00	+0.01
1980.90	68.2	0.42	1	GP	+2.2	+0.02	+0.01

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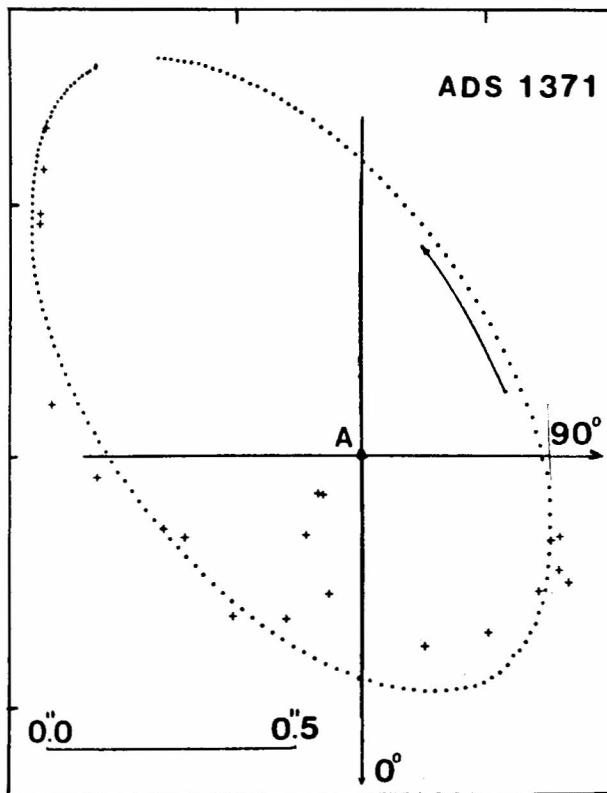


Fig. 1

Table II
ADS 7758 = IDS 10195N2468 = STF 1429, 9.0–9.0, G5

P = 1280.68	$\Omega = 113^{\circ}9$	A = -0"855
n = 0.2811	$\omega = 0^{\circ}0$	B = +1.920
T = 2170.15		F = +0.583
a = 2"10	$\pi_{\text{dyn}} = 0.014$	G = +0.265
e = 0.28	$\Sigma \Delta AB = 2.10$	C = +2.000
i = 107°7	a = 150.0 U.A.	H = +0.000

Observations and residuals

T	P	ρ	mag.	n	Obs.	(O-C)P	(O-C)P	(O-C) ρ
1827.29	272.2	1"48	8.3-8.3.2	STF	+1.1	+0.03	+0.01	
1833.26	267.4	1.58	8.5-8.8.1	STF	-2.7	-0.07	+0.15	
1857.33	267.9	1.11	8.1-8.5.2	SE	+2.7	+0.06	-0.13	
1869.17	262.4	1.19		4 D	+0.3	+0.01	+0.04	
1879.33	262.0	0.88		2 HL	+2.9	+0.05	-0.18	
1881.32	257.1	1.09		1 BIG	-1.3	-0.02	+0.04	
1885.34	257.3	0.99		3 HL	+0.2	0.00	-0.03	
1889.52	256.2	1.03		4 SP	+0.6	+0.1	+0.04	
1891.34	254.0	1.00		2 STH	-0.9	-0.02	+0.03	
1891.78	254.6	0.96		4 SP	-0.1	0.00	-0.01	
1893.27	250.6	1.19		2 TUR	-3.5	-0.06	+0.23	
1897.29	252.5	0.87		2 BRS	0.0	0.0	-0.06	
1897.31	246.4	0.80		1 BOW	-6.1	-0.10	-0.12	
1898.19	251.1	0.83		3 BOW	-1.0	-0.02	-0.09	
1899.16	250.1	0.95		3 BOW	-1.6	+0.03	+0.04	
1899.29	256.0	0.77		1 L	+4.4	+0.07	-0.14	
1900.24	252.2	0.75		1 L	+1.0	+0.02	-0.15	
1900.30	257.2	0.96		2 BOW	+6.0	+0.09	+0.06	
1900.31	249.0	0.97		1 BRY	-2.2	-0.03	+0.07	
1901.30	249.7	0.99		4 BOW	-1.0	-0.02	+0.09	
1901.32	248.2	0.65		1 L	-2.5	-0.04	-0.25	
1902.34	254.6	0.92		2 BOW	+4.3	+0.07	+0.03	
1903.20	259.8	0.82		3 BOW	+10.0	+0.15	-0.06	
1903.39	249.3	0.88	8.5-8.6.1	VBS	-0.5	-0.01	-0.11	
1904.28	256.0	1.06		3 BOW	+6.6	+0.10	+0.18	
1905.21	250.2	1.02		2 BOW	+1.3	+0.02	+0.15	
1906.23	245.6	0.98		1 DRM	-2.8	-0.04	+0.12	
1908.16	252.3	0.91		3 HZ	+4.8	+0.07	+0.06	
1908.26	246.6	0.74		2 BOW	-0.8	-0.01	-0.11	
1909.24	244.1	0.78		1 A	-2.8	-0.04	-0.06	
1909.28	248.6	0.81		4 BOW	+1.7	+0.02	-0.03	
1910.05	246.6	0.77		3 DOO	+0.1	0.00	-0.06	
1910.27	249.0	0.79		1 LAU	+2.6	+0.04	-0.04	
1912.13	244.0	0.82		3 FOX	-1.4	-0.02	0.00	
1912.28	246.3	0.62		2 BOW	+1.0	+0.01	-0.20	
1912.31	243.2	0.60		2 DOB	-2.1	-0.03	-0.22	
1914.23	250.6	0.76		2 DOB	+6.4	+0.09	-0.04	
1914.82	242.4	0.59		2 RAB	-1.5	-0.02	-0.21	
1915.08	288.6	1.57		3 PHL	+44.9	+0.63	+0.77	
1915.24	244.8	0.81		3 VBS	+1.2	+0.02	+0.01	
1918.24	237.6	0.71	0.2	1 A	-4.3	-0.06	-0.07	
1921.13	237.4	0.72		2 VBS	-2.7	-0.04	-0.04	
1922.37	234.8	0.66		3 FUR	-4.5	-0.06	-0.09	
1924.26	238.6	0.74	0.1	4 B	+0.5	+0.01	0.00	
1925.37	242.3	0.37		2 DOB	+5.0	+0.06	-0.36	
1926.27	236.1	0.65		5 FAT	-0.6	-0.01	-0.08	
1927.16	233.0	0.69		1 FUR	-3.1	-0.04	-0.03	
1927.20	231.7	0.58		1 FUR	-4.3	-0.05	-0.14	
1927.21	238.6	0.59		1 FUR	+2.6	+0.3	-0.13	
1927.22	237.4	0.79		1 FAT	+1-4	+0.02	+0.07	
1927.23	235.2	0.76		1 FUR	-0.8	-0.01	+0.04	
1927.24	239.3	0.82		1 FUR	+3.3	+0.04	+0.10	
1927.31	232.8	0.68		1 FUR	-3.2	-0.04	-0.04	
1927.32	230.7	0.59		4 RAB	-5.3	-0.07	-0.13	
1927.36	234.7	0.65		2 STG	-1.3	-0.02	-0.07	
1928.21	234.5	0.70		2 STG	-0.8	-0.01	-0.01	
1928.30	239.3	0.69		1 FAT	+4.0	+0.05	-0.03	
1930.44	228.9	0.61		3 BOW	-4.9	-0.06	-0.09	
1933.31	232.5	0.70	8.7-8.9.4	B	+0.9	+0.01	+0.01	

T	P	ρ	mag.	n	Obs.	(O-C)P	(O-C)P	(O-C) ρ
1934.21	228.1	0"64	8.9-8.9	3	VBS	-2.8	-0"03	-0"04
1934.32	232.8	0.64		4	BAZ	+2.0	+0.02	-0.04
1936.33	228.2	0.69		5	RAB	-1.0	-0.01	+0.02
1938.02	229.0	0.63		5	DUR	+1.2	+0.01	-0.03
1938.35	228.5	0.56		5	RAB	+0.9	+0.01	-0.10
1939.30	222.4	0.63		5	RAB	-4.4	-0.05	-0.02
1939.58	226.2	0.65		4	BAZ	-0.4	0.00	-0.01
1940.33	233.2	0.57		5	RAB	+7.3	+0.08	-0.8
1942.20	229.4	0.70		3	VOU	+5.2	+0.06	+0.06
1942.23	225.2	0.75	8.9-8.9	1	ARD	+0.9	+0.01	+0.10
1942.34	219.9	0.58		3	DUR	-4.3	-0.05	-0.06
1942.36	223.3	0.76		3	RAB	-0.9	-0.01	+0.12
1944.02	215.7	0.61	8.9-8.9	2	VBS	-7.0	-0.07	-0.03
1944.33	223.8	0.83		3	RAB	+1-4	+0.02	+0.19
1946.25	217.5	0.73		2	ARD	-3.2	-0.04	+0.10
1947.54	225.5	0.66		3	NEW	+6.0	+0.06	+0.04
1949.28	215.0	0.56		7	RAB	-2.9	-0.03	-0.06
1949.33	216.5	0.66	0.0	3	BAZ	-1.3	-0.01	+0.04
1950.18	213.0	0.68		1	DOM	-4.0	-0.04	+0.06
1950.18	214.5	0.52	8.7-8.7	2	MRZ	-2.5	-0.03	-0.09
1951.02	211.6	0.61		4	VBS	-4.6	-0.05	-0.00
1951.1	213.9	0.53	8.7-8.7	2	MRZ	-2.2	-0.02	-0.08
1951.22	215.2	0.58		2	DOM	-0.8	-0.01	-0.03
1951.31	215.4	0.52		6	RAB	-0.5	-0.01	-0.09
1952.30	217.5	0.57		1	DOM	+2.6	+0.03	-0.04
1953.16	212.5	0.60		2	DOM	-1.6	-0.02	-0.01
1953.31	211.6	0.54		6	RAB	-2.3	-0.02	-0.07
1954.33	211.6	0.50		5	RAB	-1.3	-0.01	-0.10
1955.19	209.2	0.65		4	MOR	-2.9	-0.03	+0.05
1955.31	210.4	0.63	8.2-8.4	4	BAZ	-1.5	-0.02	+0.03
1955.31	208.4	0.53		6	RAB	-3.5	-0.04	-0.07
1956.22	208.7	0.54	8.7-8.7	3	COU	-2.3	-0.02	-0.06
1956.33	209.8	0.50	8.9-8.9	6	RAB	-1.1	-0.01	-0.10
1958.04	209.6	0.67	9.0-9.1	1	B	+0.4	+0.00	+0.07
1958.26	208.1	0.66	8.7-8.7	3	COU	-0.9	-0.01	+0.06
1958.31	206.4	0.45	9.0-9.2	1	B	-2.5	-0.03	-0.15
1958.52	204.6	0.66	0.0	2	WOR	-4.1	-0.04	+0.06
1959.20	207.5	0.75		7	HEI	-0.5	0.00	+0.16
1960.26	206.2	0.58		6	HEI	-0.7	-0.01	-0.01
1961.14	202.9	0.62		3	VBS	-3.1	-0.03	+0.03
1961.25	204.7	0.56	8.8-8.8	3	DUJ	-1.2	-0.01	-0.03
1962.14	204.7	0.52	9.1-9.2	4	B	-0.2	0.00	-0.07
1962.33	204.8	0.50	0.0	4	HLN	0.0	0.00	-0.09
1963.14	204.1	0.51	0.0	4	WOR	+0.2	+0.00	-0.08
1964.3	197.9	0.66		5	HEI	-4.8	-0.05	+0.07
1964.57	201.5	0.54	8.6-8.6	3	COU	-1.0	-0.01	-0.05
1966.31	197.9	0.54	0.2	3	WOR	-2.7	-0.03	-0.05
1966.32	199.9	0.63	0.3	3	WOR	-0.7	-0.01	+0.04
1966.33	199.5	0.64	8.9-9.0	3	BAZ	-1.1	-0.01	+0.05
1969.20	191.1	0.58		4	HEI	-6.4	-0.07	-0.01
1970.94	194.1	0.58	8.6-8.6	3	COU	-1.6	-0.02	-0.01
1971.17	192.7	0.46		2	VBS	-2.8	-0.03	-0.13
1973.75	189.2	0.63	0.1	4	WOR	-3.6	-0.04	+0.04
1973.87	190.7	0.48	0.2	3	BEH	-2.0	-0.02	-0.11
1976.14	191.4	0.67	9.0-9.0	1	ZUL	+1.1	+0.01	+0.08
1976.20	187.8	0.64		1	MUL	-2.4	-0.02	+0.05
1977.19	187.9	0.63		3	HEI	-1.4	-0.01	+0.04
1981.29	182.3	0.63	8.8-8.8	3	ZUL	-2.7	-0.03	+0.03
1981.30	184.4	0.64		3	GP	-0.6	-0.01	+0.04

Ephemerides

T	P	ρ	T	P	ρ
1980.0	186.3	0"60	1992.0	174.8	0"64
1982.0	184.3	0.60	1994.0	173.1	0.64
1984.0	182.4	0.61	1996.0	171.3	0.65
1986.0	180.4	0.61	1998.0	169.6	0.66
1988.0	178.5	0.62	2000.0	168.0	0.67
1990.0	176.7	0.62	2002.0	166.4	0.68

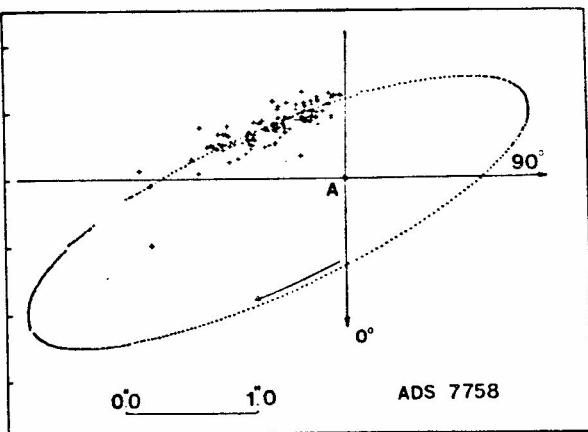


Fig. 2

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Table III
 ADS 11989 = IDS 18584S2141 = H 126, 7.5–7.8, GO

$P = 356.44$	$\Omega = 55.510$	$A = +0.7200$
$n = 1.010$	$\omega = 89.5$	$D = -0.4750$
$T = 1896.51\pi$	$\text{dyn} = 0.024$	$F = -0.7150$
$a = 1.37$	$\sum m_{AB} = 1.5$	$G = -1.0600$
$e = 0.44$	$a = 57.1$ U.A.	$C = +1.0116$
$i = 132.4$		$H = +0.0088$

Observations and residuals

T	P	ρ	mag.	n	Obs.	(O-C)P	(O-C)P	(O-C) ρ
1801.67	very close							
1873.00	$40^{\circ} \pm 1^{\circ} 0^{\prime}$			1	BU	+4"9	+0"07	+0"21
1879.53	22.4 0.83			2	CinO	-06	-0.01	+0.15
1890.48	353.2 0.58			3	BU	+2.0	+0.02	+0.06
1891.51	349.3 0.48			3	BU	+2.0	+0.02	-0.03
1891.51	331.6 0.35			6	BAR	-15.7	-0.14	-0.16
1896.73	317.5 0.44			2	A	-8.1	-0.07	-0.04
1897.75	310.7 0.56			IN SEE		-10.6	-0.09	+0.08
1898.46	312.8 0.51			3	A	-5.5	-0.05	+0.02
1899.55	310.0 0.51			2	A	-3.7	-0.03	+0.02
1900.75	303.2 0.40	7.8	3	HU		-5.6	-0.05	-0.10
1900.77	289.7 0.50			IN SOLA		-18.9	-0.16	+0.02
1901.55	300.9 0.43			2	A	-4.7	-0.04	-0.07
1903.41	295.4 0.48			2	A	-3.1	-0.03	-0.04
1908.72	279.0 0.64			6	MCO	-2.2	-0.02	+0.05
1909.62	276.4 0.70			3	Al,OL2	-2.4	-0.03	+0.09

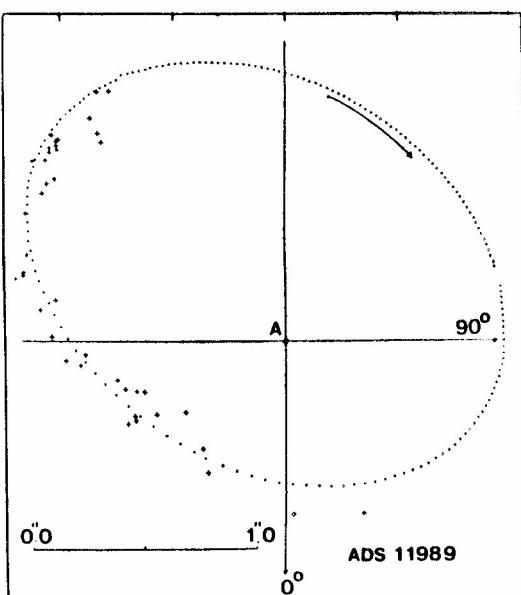


Fig. 3

T	P	ρ	mag.	n	Obs.	(O-C)P	(O-C)P	(O-C) ρ
1911.62	275.0	0.60		4	A	+1-4	+0.02	-0.04
1914.60	267.8	0.76		4	A2, I2	+0.9	+0.01	+0.07
1919.62	259.4	0.82		2	A	+1.7	+0.02	+0.04
1921.90	255.5	0.75		6	VBS	+1-4	+0.02	-0.07
1924.42	252.6	0.96		5	VOU4, 81	+2.1	+0.03	+0.10
1925.33	250.8	0.92		4	VOU	+1.5	+0.02	+0.05
1926.40	249.8	0.92		4	B2, A2	+1.9	+0.03	+0.03
1931.62	244.8	0.94		3	VBS	+3.0	+0.05	-0.02
1939.99	235.5	1.057.5-7.8	3	B		+1.9	+0.04	-0.01
1947.33	228.9	1.04		3	B	+1.4	+0.03	-0.09
1948.66	225.0	1.18		5	RAB	-1.4	-0.03	+0.04
1949.71	226.6	1.057.8-8.0	4	B		+1.0	+0.02	-0.09
1949.73	222.7	1.14		5	RAB	-2.9	-0.06	-0.01
1950.72	220.6	1.15		5	RAB	-4.2	-0.09	0.00
1951.70	219.0	1.14		4	RAB	-5.1	-0.10	-0.02
1952.79	218.1	1.21		3	RAB	-5.2	-0.11	+0.04
1953.68	224.6	1.047.5-7.5	2	CHURMS		+1.9	+0.04	-0.13
1953.77	217.4	1.17		5	RAB	-5.2	-0.11	0.00
1954.74	220.5	1.17		6	RAB	-1.4	-0.03	-0.01
1955.71	218.2	1.17		5	RAB	-3.1	-0.06	-0.02
1956.72	218.8	1.15		5	RAB	-1.7	-0.04	-0.04
1975.17	209.4	1.05		3	ZUL	+1.0	+0.02	-0.21
1975.72	209.4	1.10		4	WOR	+1.4	+0.03	-0.17
1976.54	209.4	1.18		2	HLD	+1-9	+0.04	-0.09
1981.30	205.0	1.28		2	ZUL	+0.4	+0.01	+0.00
1981.30	204.1	1.27		2	GP	-0.5	-0.01	-0.01

Ephemerides

T	P	ρ	T	P	ρ
1980.0	205.4	1.28	1992.0	198.1	1.29
1982.0	204.1	1.28	1994.0	196.9	1.29
1984.0	202.9	1.28	1996.0	195.7	1.29
1986.0	201.7	1.28	1998.0	194.5	1.29
1988.0	200.5	1.29	2000.0	193.3	1.29
1990.0	199.3	1.29	2002.0	192.1	1.29

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Table IV
ADS 16877 = IDS 23526N4453 = STT 500 AB, 6.3–7.2, B9

$P = 351.22$	$\Omega = 157.1$	$A = -0.3790$
$n = 1.0250$	$\omega = 0.0$	$B = +0.1600$
$T = 2103.29$		$F = -0.1222$
$a = 0.41$	$\pi_{\text{dyn}} = 0.004$	$G = -0.2890$
$e = 0.39$	$\Sigma \Delta AB = 8.7^\circ$	$C = 0.0000$
$i = 40.3$	$a = 102.5 \text{ U.A.}$	$H = +0.2652$

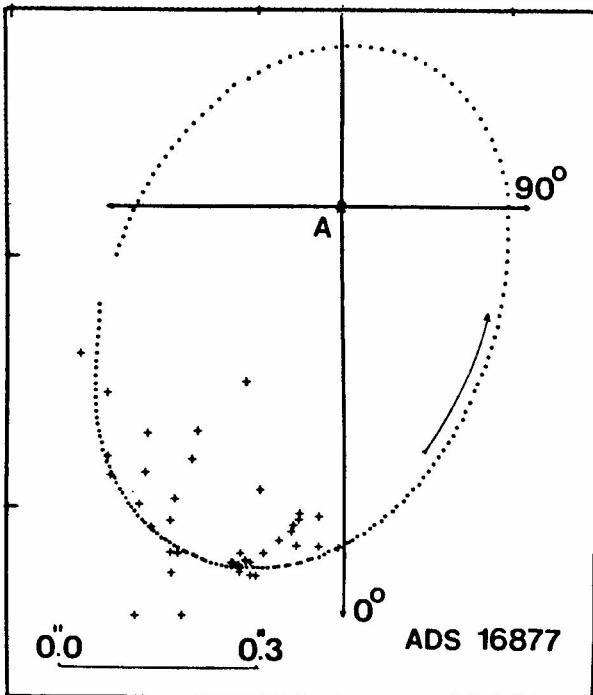


Fig. 4

Observations and residuals

T	P	ρ	mag	n	Obs.	(O-C)P	(O-C)P	(O-C) ρ
1845.24	299.4	0.45	2	STT		-0.1	-0.00	+0.02
1852.82	308.5	0.45	2	STT		+4.2	+0.03	+0.00
1873.52	319.5	0.45	7	D		+4.3	+0.04	-0.06
1879.80	316.5	0.52	3	HL		-1.6	-0.01	0.00
1883.2	326.7	0.40	6	EN		+7.1	+0.07	-0.13
1888.91	322.6	0.50	3	HL		+0.5	0.00	-0.04
1888.95	317.6	0.53	6	SP		-4.5	-0.04	-0.01
1889.80	319.3	0.53	2	H		-3.2	-0.03	-0.01
1898.62	325.8	0.54	3	HU		-0.2	-0.00	-0.01
1902.09	333.0	0.69	3	DOO		+5.6	+0.05	+0.13
1905.61	330.4	0.30	1	FRH		+1.6	+0.01	-0.26
1908.35	329.5	0.51	12PRZ4, FUR2,					
			DOB3, VBS			-0.3	0.00	-0.05
1910.15	328.6	0.44	4	DOB		-1.9	-0.02	-0.12
1913.07	330.6	0.54	10.NeuJ3,					
			VOU5, FOX2			-1.0	-0.01	-0.03
1914.97	343.3	0.44	1	RAB		+10.9	+0.11	-0.13
1916.53	334.0	0.58	6FOX3, DOO3			+1.0	+0.01	+0.01
1923.30	334.6	0.61	5 A2, MAG3			-0.9	-0.01	+0.04
1924.50	328.4	0.56	4 Gr0			-7.5	-0.07	-0.01
1925.58	333.4	0.58	7 KPZ1,					
			DOB2, RAB4			-2.9	-0.03	+0.01
1932.71	338.6	0.66	1 HINTze			-0.4	0.00	+0.09
1943.76	344.5	0.55	3 VBS			+1.4	+0.01	-0.02
1948.82	342.0	0.56	7 RAB			-3.1	-0.03	0.00
1949.79	342.7	0.56	7 RAB			-2.7	-0.03	0.00
1950.88	342.2	0.56	6 RAB			-3.6	-0.04	+0.00
1951.87	342.7	0.56	8 RAB			-3.6	-0.03	+0.00
1952.92	343.5	0.56	6 RAB			-3.2	-0.03	+0.00
1953.88	343.4	0.54	9 RAB			-3.6	-0.04	-0.02
1954.89	344.3	0.55	9 RAB			-3.1	-0.03	+0.01
1955.86	343.8	0.57	7 RAB			-4.0	-0.04	+0.02
1955.86	345.1	0.55	3 MUL			-2.7	-0.03	+0.00
1957.88	345.4	0.57	8 RAB			-3.2	-0.03	+0.02
1958.43	346.8	0.53	4 B			-2.1	-0.02	-0.02
1959-84	351.5	0.46	9 Gr0			+2.1	+0.02	-0.09
1961.76	348.3	0.51	4 B			-1.9	-0.02	-0.04
1961.81	350.7	0.48	1 HLD			+0.5	+0.00	-0.07

ORBITS OF FOUR VISUAL BINARIES

T	P	ρ	mag.	n	Obs.	(O-C)P	(O-C) ρ	(O-C)mag.
1962.80	345.9	0.57	3	HEI		-4.7	-0.04	+0.03
1963.85	351.8	0.51	4	WOR		+0.7	+0.01	-0.03
1069.85	350.2	0.49	2	COU		-3.4	-0.03	-0.04
1971.10	351.3	0.47	3	BAZ		-2.9	-0.03	-0.06
1971.83	355.1	0.46	3	WOR		+0.7	+0.01	-0.07
1975.72	355.2	0.51	1	ZUL		-1.0	-0.01	-0.01
1980.90	358.9	0.51	1	ZUL		+0.4	+0.00	+0.00

Ephemerides

T	P	ρ	T	P	ρ
1981-0	358.6	0.51	1993.0	4.5	0.48
1983.0	359.5	0.51	1995.0	5.6	0.47
1985.0	0.5	0.50	1997.0	6.7	0.47
1987.0	1.5	0.50	1999.0	7.8	0.46
1989.0	2.5	0.49	2001.0	9.0	0.46
1991.0	3.5	0.49	2003.0	10.2	0.45

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ORBITES DE TROIS ETOILES DOUBLES VISUELLES
(ADS 674 = A 921AB, ADS 1345 = A 1 et ADS 1393 = HU804)

V. Erceg

(Reçu le 28.7.81.)

RÉSUMÉ: On donne les éléments des orbites, les masses, les magnitudes absolues et les parallaxes dynamiques orbitales de trois étoiles doubles visuelles, les éléments étant déterminés en utilisant la méthode de Thiele-Innes Van den Bos.

ORBITE DE ADS 674 = A 921AB
Pos. (1950): 00^h 46^m 2; +57° 03' Mgn.: 9.3 – 9.8; Type sp. –

Tableau I
Eléments orbitaux, les quantités astrophysiques et les constantes

P = 123.18 ans	
n = 29226	
T = 1869.99	A = -0"0400 $\pi_{\text{dyn.orb}}$ = 0"004
e = 0.48	B = -0"1300 M_A = 2.6
a = 0"167	F = +0"1550 M_A = 3.1
i = 36.02	G = -0"0600 M_A = 1.76 0
Ω = 151.01	C = +0"0968 M_B = 1.56 0
ω = 99.66	H = ±0"0165 a = 36.8 U.A.
$T_{\Omega, \psi}$ = 1978.0; 2003.8	

Tableau II
Ephémérides

t	θ°	ρ''	t	θ°	ρ''
1981.0	160.8	0.12	1986.0	184.8	0.10
1982.0	164.7	0.12	1987.0	191.5	0.09
1983.0	169.0	0.11	1988.0	199.1	0.08
1984.0	173.7	0.11	1989.0	207.7	0.08
1985.0	178.9	0.10	1990.0	217.3	0.08

Tableau III
Observations et les résidus

t	θ°	ρ''	Obs.	n	Références	$(O-C) \frac{\circ}{\theta}$	$(O-C) \frac{''}{\rho}$
1905.620	38.9	0.17	A	3	Lick Obs.Bul.V.4,N.4, 1906.	+6.5	-0.01
1915.700	47.2	0.21	A	2	Lick Obs.Bul.V.14, N.62,1929	-2.0	+0.02
1925.820	60.4	0.19	A	1	Lick Obs.Bul.V.14, N.62,1929	-4.2	-0.01
1929.710	71.8	0.17	A	2	Lick Obs.Bul.V.16, N.96	+1.6	-0.03
1947.010	98.3	0.20	VBS	6	Pub.Yerkes Obs.V.9, Pt.1	+3.6	0.00
1958.030	110.6	0.21	B	1	-"-	-0.3	+0.02
1958.090	108.2	0.21	B	1	-"-	-2.8	+0.02
1962.320	120.3	0.18	B	4	Astr.J.V68,582	+2.5	-0.01
1962.960	111.	0.21	COU	2	J.Obs.V.46,155	-7.	+0.03
1964.775	128.3	0.17	WOR	4	Pub.US Naval Obs. V.22, Pt. 2	+6.3	-0.01

ORBITE DE ADS 1345 = A 1
 α, δ (1950): 01^h 39^m9; -07° 00' Mgns.: 8.5 - 9.0; Type sp. F2

Tableau I
Eléments orbitaux, les quantités astrophysiques et les constantes

$P = 298.83$ ans	
$n = 1^{\circ}2047$	
$T = 1879.36$	
$e = 0.51$	
$a = 0^{\circ}525$	
$i = 25^{\circ}8$	
$\Omega = 24^{\circ}0$	
$\omega = 58^{\circ}9$	
$T_{\Omega, \nu} = 2231.6, 1930.2$	

$A = +0.^{\circ}0825 \quad \pi_{dyn} = 0.^{\circ}008$

$B = +0.^{\circ}5100 \quad M_A = 3.1$

$F = -0.^{\circ}4800 \quad M_B = 3.6$

$G = +0.^{\circ}0395 \quad m_A = 1.570$

$C = +0.^{\circ}1959 \quad m_B = 1.400$

$H = +0.1180 \quad a = 64.0$ U.A

Tableau II
Ephémérides

t	θ°	ρ''	t	θ°	ρ''
1981.0	240.0	0.73	1986.0	242.3	0.74
1982.0	240.4	0.74	1987.0	242.8	0.75
1983.0	240.9	0.74	1988.0	243.2	0.75
1984.0	241.4	0.74	1989.0	243.7	0.75
1985.0	241.9	0.74	1990.0	244.2	0.75

Tableau III
Observations et les résidus

t	θ°	ρ''	Obs.	n	References	$(O-C)_{\theta}^{\circ}$	$(O-C)_{\rho}''$
1. 1899.78	165.2	0.31	A	3	Astron.Nachr.V.152,161	+8.9	+0.01
2. 1902.75	172.5	0.38	A	4	Lick Obs.Bul.V.14,62,1929	+8.4	+ 6
3. 1915.70	189.6	0.46	A	2	Lick Obs.Bul.V.14,62,1929	+0.4	+ 5
4. 1920.14	193.7	0.41	A	3	Lick Obs.Bul.V.14,62,1929	-1.7	- 3
5. 1926.03	199.2	0.44	A	2	Lick Obs.Bul.V.14,62,1929	-3.3	- 4
6. 1928.97	204.2	0.44	A	2	Lick Obs.Bul.V.14,62,1929	-1.4	- 6
7. 1929.62	207.2	0.56	VOU	3	Ann.Bosscha Obs.Lemb.V.6,Pt.1	+1.0	+ 5
8. 1929.78	201.8	0.51	VBS	2	Pub.Yerkes Obs.V.8,47	-4.6	0
9. 1939.91	208.7	0.52	A	2	Lick Obs.Bul.V.18,109	-6.7	- 5
10. 1933.69	208.0	0.53	KUI	1	Astrophys.J.Supp.V.6,1.	-2.1	0
11. 1934.96	210.0	0.53	B	4	Union Obs.Circ.N.100,481	-0.8	- 1
12. 1936.71	211.5	0.61	SMW	4	Ann.Bosscha Obs.V.9,Pt.1	-1.3	+ 6
13. 1938.59	212.4	0.54	VOU	4	Ann.Bosscha Obs.Lemb.V.6,Pt.4	-1.9	- 2
14. 1942.61	217.9	0.55	VOU	3	Manuscript, See J.Obs.V.38,109	+0.4	- 3
15. 1943.30	216.4	0.58	BAZ	3	J.Obs.V.28,1	-1.6	- 1
16. 1947.70	217.1	0.51	JEF	1	Unpublished	-4.1	- 10
17. 1948.17	220.0	0.54	B	4	Union Obst.Circ.N.111,13	-1.5	- 7
18. 1948.71	218.1	0.75	VBS	2	Pub.Yerkes Obs.V.8,15y	-3.8	+ 14
19. 1952.26	222.3	0.57	B	3	Union Obs.Circ.N.113,185	-1.9	- 6
20. 1957.61	226.1	0.69	VBS	3	Pub.Yerkes Obs.V.9,Pt.2	-1.5	+ 4
21. 1957.73	226.7	0.60	B	3	Astrophys.J.Supp.V.4,N36,45	-0.9	- 5
22. 1958.94	229.5	0.59	COU	3	J.Obs.V.42,17	+1.2	- 7
23. 1959.87	226.3	0.56	KNP	3	Union Obs.Circ.N.119,331	-2.6	- 10
24. 1959.92	225.8	0.78	HEI	4	Astron.Nachr.V.285,249	-3.1	+ 12
25. 1961.67	227.3	0.74	B	4	Astron.J.V.67,555	-2.6	+ 7
26. 1961.72	228.6	0.62	WOR	3	Pub.US Naval Obs.V.18,Pt.6	-1.4	- 5
27. 1962.72	239.0	0.72	VBS	1	Comm.Lunar Plan.Lab.V.3,N.51	+8.5	+ 5
28. 1962.94	234.4	0.70	VBS	1	Comm.Lunar Plan.Lab.V.3,N.51	+3.8	+ 2
29. 1962.96	232.7	0.68	VBS	1	Comm.Lunar Plan.Lab.V.3,N.51	+2.0	0
30. 1963.65	235.3	0.76	VBS	5	Contr.N.180,Kitt Peak N.O.	+4.3	+ 8
31. 1964.83	235.3	0.85	VBS	2	Comm.Lunar Plan.Lab.V.3,N.51	+3.6	+ 17

ORBITES DE TROIS ETOILES DOUBLES VISUELLES

32.	1965.94	235.8°	0.69	NBG	5	Republic Obs.Circ.N.125,105	+3.5	0	
33.	1969.20	232.1	0.63	WAK	3	Pub.US Naval Obs.V.22,Pt.5,1972	-2.0	-	7
34.	1971.21	234.2	0.87	WAK	3	Pub.US Naval Obs.V.22,Pt.5,1972	-0.9	+	17
35.	1971.67	237.7	0.67	WOR	4	Pub.US Naval Obs.V.24,Pt.6,1978	+2.4	-	4
36.	1974.81	236.1	0.74	HEI	3	Astrophys.J.Supp.V.29 315,1975	-0.8	+	2
37.	1977.92	237.4	0.72	HLN	1	Pub.Astr.Soc.Pacific,V.90,463, 1978	-1.1	0	
38.	1978.85	239.3	0.67	HEI	3	Unpublished	+0.4	-	6
39.	1980.90	238.4	0.64	DZ	1	Bul.Astr.Obs.Belgrade,N.131,1981	-1.5	-0.09	

* Quadrant change

ORBITE DE ADS 1393 = HU804
Pos. (1950): 01^h 43^m 3; + 33° 28' Mgn.: 8.5 – 10.3; Type sp. F2

Tableau I
Eléments orbitaux, les quantités astrophysiques et les constantes

P =	610.17 ans
n =	0°.5900 A = +0°.2380 $\pi_{dyn.orb.}$ = 0°.003
T =	1874.81 B = -0°.2600 M_A = 1.1
e =	0.21 F = +0°.2297 M_B = 2.9
a =	0°.380 G = +0°.2770 Ω_A = 2.670
i =	29.5 C = ±0°.1420 Ω_B = 1.640
Ω =	87°.14 H = ±0°.1221 a = 126.7 U.A.
ω =	229°.30
$T_{\Omega, \Omega}$ =	2062.1; 1819.1

Tableau II
Ephémérides

t	θ°	ρ''	t	θ°	ρ''
1981.0	46.6	0.34	1986.0	49.3	0.34
1982.0	47.3	0.34	1987.0	50.4	0.34
1983.0	47.9	0.34	1988.0	51.0	0.35
1984.0	48.6	0.34	1989.0	51.5	0.35
1985.0	49.2	0.34	1990.0	52.2	0.35

Tableau III
Observations et les résidus

t	θ°	ρ''	Obs.	n	References	$(O-C) \frac{\circ}{\theta}$	$(O-C) \frac{''}{\rho}$
1903.620	337.6	0.27	HU	2	Lick Obs.Bul.V.3,95	-3.1	0.000
1921.670	0.2	0.30	A	2	Lick Obs.Bul.V.11,58	+1.6	+0.03
1923.696	2.4	0.28	VBS	2	Pub.Yerkes Obs.V.5,Pt1	+1.8	+0.01
1928.830	2.1	0.28	VBS	1	Pub.Yerkes Obs.V.8,159	-3.3	0.00
1930.690	3.2	0.24	VBS	1	Pub.Yerkes Obs.V.8,159	-4.0	-0.04
1937.670	13.2	0.30	VBS	1	Pub.Yerkes Obs.V.8,159	-0.4	+0.02
1946.00	20.9	0.26	VBS	1	Pub.Yerkes Obs.V.8,159	+0.1	-0.03
1946.810	20.4	0.30	VBS	1	Pub.Yerkes Obs.V.8,159	-1.1	+0.01
1951.810	24.1	0.23	VBS	1	Pub.Yerkes Obs.V.9,Pt.2	-1.5	-0.07
1958.660	34.2	0.32	B	3	Pub.Yerkes Obs.V.9,Pt.1	+3.2	+0.01
1961.910	36.4	0.29	B	3	AstronJ.V.67,555	+3.0	-0.02
1964.549	35.0	0.31	WOR	3	Pub.US Naval Obs.V.22,Pt.2.	-0.4	0.00
1967.850	35.3	0.30	COU	2	J.Obs.V.51,337	-2.5	-0.02
1971.830	37.7	0.36	MUL	1	Astron Astrophys.Supp.V.23,205, 1976	-2.9	+0.04
1976.900	51.6	0.34	HEI	3	Astrophys.J.Supp.V.37,343,1978	+7.6	+0.01

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MICROMETER MEASURES OF DOUBLE STARS

(Series 33)

D. J. Zulević

(Received January 10, 1981)

SUMMARY: Presented here are 289 measures of 134 Double Stars made with 65/1055 cm refractor of Belgrade Observatory.

The present list is a continuation of the published series of measures of double stars in Belgrade, made in the period 1980. 103 – 1980. 902. In the Table I of measures the columns give: ADS number, double star designation, multiple, position for 1900 (IDS), epoch omitting the century, position angle, separation, esti-

mated magnitudes, number of nights and notes. An asterisk in the secind column indicates that there is a note at the end of Table I.

In the present work the distribution of 289 measures of distances is as folows:

$\rho \leq 0^{\circ}50$	$0^{\circ}50 < \rho \leq 1^{\circ}00$	$1^{\circ}00 < \rho \leq 1^{\circ}50$	$1^{\circ}50 < \rho \leq 2^{\circ}00$	$\rho > 2^{\circ}00$	Σ
28	146	57	33	25	289
9.6%	50.5%	19.7%	11.1%	8.8%	100%

Table I
Micrometer measures of double stars

ADS	DISC IDC	Mult.	Epoch 1900+	P	ρ	m	n	Notes
61	STT 00010N5753	3062	80.902	289°1	1°34	6.4-7.5	1	Baize, 1958: -0°2, - 0"09
102	STF 00038N7910	2	80.678 80.881 80.779	26.2 21.7 23.9	0.62 0.62 0.62	6.8-7.1 1 2	1 1 Heintz, 1952: -4°9, +0"01	
147	BU 00067N2752	255	80.897	71.6	0.50		1	Changed 27° since 1875

D. J. ZULEVIĆ

ADS	DISC IDC	Mult.	Epoch 1900+	P	ρ	m	n	Notes
202	ES 00106N5200	865 AB	80.810 <u>80.881</u>	80°2 77.5	1".67 <u>1.60</u>	9.5-9.6	1 <u>1</u>	
			80.845	78.9	1.63		2	Changed 133° since 1910
207	STF 00106N7624	13	80.678 80.881 <u>80.902</u>	60.3 59.9 <u>59.8</u>	0.91 0.88 <u>0.85</u>	6.7-7.2	1 <u>1</u> <u>1</u>	
			80.820	60.0	0.88		3	Heintz, 1960: +3°1, +0".02
220	STF 00115N3604	19	80.678 80.790 80.793 80.815 <u>80.829</u>	141.6 139.3 139.6 140.4 <u>140.2</u>	2.07 2.24 2.17 2.24 <u>2.27</u>	7.0-9.2	1 <u>1</u> <u>1</u> <u>1</u> <u>1</u>	
			80.781	140.2	2.20		5	Changed 7° since 1836
221	STT 00115N3556	4	80.678 <u>80.711</u>	172.1 <u>173.0</u>	0.57 <u>0.57</u>	8.5-8.9	1 <u>1</u>	
			80.695	172.6	0.57		2	Muller, 1957: +1°6, +0".03
256	HJ 00135N2512	1015	80.728	302.5	5.73	9.1-9.6	1	Changed 25° since 1896
283	HJ 00154N6707	1018	80.825 <u>80.881</u>	90.5 <u>87.7</u>	1.44 <u>1.40</u>	8.6-9.2	1 <u>1</u>	
			80.853	89.1	1.42		2	Muller, 1957: +2°8, -0".03
293	STT 00158N6627	6	80.892	151.3	0.48	7.5-8.5	1	Muller, 1954: 0°0, -0".14
434	STT 00262N5358	12	80.892	182.9	0.52	5.5-5.8	1	Heintz, 1963: -0°2, -0".05
684	BU 00448N5005	232	80.678 <u>80.711</u>	239.5 <u>237.5</u>	0.81 <u>0.72</u>	8.5-9.0	1 <u>1</u>	
			80.695	238.5	0.76		2	Baize, 1964: +1°9, -0".07
755	STF 00496N2305	73	80.678 <u>80.891</u>	258.7 <u>260.8</u>	0.64 <u>0.59</u>	6.1-6.7	1 <u>1</u>	
			80.784	259.8	0.62		2	Muller, 1957: -2°5, 0".0
862	STT 00573N4650	21	80.881 80.892 <u>80.902</u>	177.2 175.8 <u>172.9</u>	0.76 0.64 <u>0.63</u>	6.7-8.0	1 <u>1</u> <u>1</u>	
			80.892	175.3	0.68		3	Heintz, 1966: +0°2, -0".21
999	BU 01084N6025	1100	80.892	207.7	0.41	8.3-8.3	1	Muller, 1954: +0°5, -0".08 Zulević, 1972: -0°9, -0".17
1345	A 01374S0676	1	80.897	238.4	0.64	8.5-9.0	1	Changed 73° since 1899
1371	BU 01384N5637	453	80.897	65.9	0.40	10.1-10.4	1	Baize, 1973: +7°1, -0".18
1548	A 01513N3032	819 AB	80.728 <u>80.810</u>	189.0 <u>189.4</u>	0.58 <u>0.59</u>	8.2-8.7	1 <u>1</u>	
			80.769	189.2	0.59		2	Zulević, 1980: -3°8, -0".08
1709	STF 02076N4701	228	80.815 80.881 <u>80.892</u>	268.2 267.2 <u>268.0</u>	0.97 0.96 <u>0.96</u>	6.4-7.3	1 <u>1</u> <u>1</u>	
			80.863	267.8	0.96		3	Heintz, 1954: +1°1, -0".09
1737	STF 02100N6053	234 AB	80.815 80.825 <u>80.892</u>	235.9 235.8 <u>235.5</u>	0.87 0.85 <u>0.82</u>	8.5-9.4	1 <u>1</u> <u>1</u>	
			80.844	235.7	0.85		3	Baize, 1955: -4°6, -0".04

MICROMETER MEASURES OF DOUBLE STARS

ADS	DISC IDC	Mult.	Epoch 1900+	P	ρ	m	n	Notes
2034	STT 43 02349N2612		80.900	9°.0	0".75	7.2-8.8	1	Heintz, 1962: +1°.2, -0".27
2122	STF 305 AB 02418N1857		80.810	310.1	3.47	7.4-8.3	1	Rabe, 1961: +1°.0, -0".18
2377	STT 50 AB 03027N7110		80.825 80.889 80.857	346.4 348.1 347.2	1.09 1.09 1.09	8.5-8.5	1 1 2	Popovic, 1972: +0°.2, +0".02
2416	STF 367 03089N0022		80.147	143.1	0.97	8.9-8.9	1	Heintz, 1963: -0°.3, -0".03
2436	STT 52 AB 03088N6517		80.892 80.895 80.893	71.7 72.8 72.2	0.49 0.53 0.51	6.9-7.5	1 1 2	Heintz, 1963: +4°.2, +0".07
2446	STT 53 03113N3816		80.711 80.810 80.815 80.900 80.809	262.0 264.7 265.0 265.1 264.2	0.70 0.82 0.83 0.82 0.79	7.7-8.5	1 1 1 1 4	Rabe, 1948: -1°.1, -0".06
2612	STF 400 AB 03268N5942		80.825 80.889 80.895 80.870	258.5 258.0 258.4 258.3	1.03 1.04 1.02 1.03	6.9-7.9	1 1 1 3	Baize, 1952: -0°.6, -0".24
2630	A 1535 03294N4201		80.895	318.6	0.45	9.0-9.4	1	Morel, 1969: +3°.6, -0".11 Heintz, 1969: +10°.6, -0".16
2799	STT 65 03443N2557		80.900	211.0	0.80	6.3-6.0	1	Wierzbinski, 1957: +2°.2, +0".19
2959	STF 483 03574N3914		80.895	78.4	0.81	7.4-8.9	1	Couteau, 1958: +2°.7, -0".10
2980	A 1710 03595N4309		80.889	335.7	0.39	8.3-8.3	1	Heintz, 1969: +14°.1, -0".07
2995	STT 531 04009N3749		80.139 80.711 80.810 80.815 80.619	3.3 11.2 5.1 9.9 7.4	1.35 1.32 1.37 1.46 1.37	7.3-9.0	1 1 1 1 4	Rabe, 1961: +5°.0, -0".09
3082	STT 77 04096N3127		80.139 80.147 80.143	271.9 269.1 270.5	0.78 0.71 0.75	8.2-8.2	1 1 2	Muller, 1956: +0°.3, -0".02
3098	STF 511 04095N5832		80.889	115.6	0.37	7.4-7.9	1	Heintz, 1969: +10°.3, -0".03
3174	STF 535 04178N1109		80.889 80.895 80.892	291.4 291.6 291.5	1.08 1.09 1.09	7.0-8.5	1 1 2	Changed 63° since 1831
3264	STF 544 04244N1525		80.108 80.810 80.815 80.578	23.1 18.6 18.0 19.9	1.65 1.68 1.70 1.68	5.8-8.3	1 1 1 3	Kuiper, 1937: +3°.2, -0".11
3390	STF 577 04355N3719		80.103 80.124 80.139 80.122	20.0 22.2 17.2 19.8	1.07 1.17 1.10 1.11	8.5-8.5	.1 .1 .1	
3614	HZ 445 04558N2041		80.889	266.5	0.36	8.6-8.9	1	Baize, 1957: -4°.1, -0".05

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ADS	DISC IDC	Mult.	Epoch 1900+	P	ρ	m	n	Notes
3956	STF 677 05153N6317		80.141 153.5 80.147 155.1 80.171 153.3 80.810 157.1 <u>80.815</u> <u>159.1</u> 80.419 155.6	0."93 0.87 0.90 0.97 0.97 0.93		7.9-8.2	1 1 1 1 1 5	Heintz, 1962: -2°.5, -0."10
4076	A 1034 05208N7044		80.895 150.4	0.48		8.5-9.0	1	Heintz, 1963: +5°.4, -0."10
5400	STF 948 AB 06374N5933		80.139 81.9 80.810 80.1 <u>80.815</u> <u>79.9</u> 80.588 80.6	1.63 1.64 1.67 1.61		5.3-6.2	1 1 1 3	Brosche, 1957: +2°.1, -0."08
5400	STF 948 06374N5933		80.810 310.3	8.91		5.3-8.3	1	No change since 1831
5447	STT 156 06416N1818		80.141 242.8 <u>80.147</u> <u>243.7</u> 80.144 243.2	0.53 0.58 0.56		6.7-7.2	1 1	
5469	A 2731 06432N0744		80.141 55.6 <u>80.889</u> <u>56.1</u> 80.515 55.8	0.98 0.97 0.98		8.7-9.5	1 1 2	Muller, 1957: -4°.9, +0."03
5535	A 513 06466N2465		80.890 235.7	0.34		9.6-9.7	1	Heintz, 1963: +1°.8, -0."17
5871	STF 1037 AB 07066N2724		80.139 320.7 <u>80.147</u> <u>320.1</u> 80.143 320.4	1.20 1.25 1.23		7.2-7.2	1 1 2	Karmel, 1939: -0°.3, -0."08
6117	STF 1093 07227N4971		80.141 182.2 80.171 184.3 <u>80.237</u> <u>185.8</u> 80.183 184.1	0.70 0.73 0.72 0.72		8.8-8.8	1 1 1 3	Baize, 1958: -5°.5, -0."03
6175	STF 1110 AB 07282N3166		80.139 97.3 80.147 100.1 <u>80.810</u> <u>97.9</u> 80.365 98.4	1.94 1.93 2.40 2.09		2.0-2.9	1 1 1 3	Rabe, 1958: +4°.2, -0."15
6276	STT 177 07350N3740		80.141 169.2 80.890 167.4 <u>80.895</u> <u>166.8</u> 80.645 167.8	0.38 0.38 0.45 0.40		8.0-9.0	1 1 1 3	Muller, 1977: -10°.4, 0."0
6347	HO 247 07402N2122		80.141 220.2 80.890 221.2 <u>80.895</u> <u>220.5</u> 80.645 220.6	0.35 0.34 0.38 0.36		8.0-8.5	1 1 1 3	Heintz, 1962: -14°.9, +0."03
6650	STF 1196 AB 08065N1757		80.147 276.8 80.890 275.5 <u>80.895</u> <u>273.0</u> 80.644 275.1	0.83 0.82 0.76 0.80		5.6-6.3	1 1 1 3	Gasteyer, 1954: -0°.8, 0."0
7054	A 1584 08456N5480		80.141 176.2	0.30		8.2-8.2	1	Heintz, 1966: +17°.6, +0."02 Heintz, 1980: -5°.1, +0."10 Zulević, 1973: +17°.2, -0."03
7067	STF 1280 08460N7071		80.128 114.6 80.139 115.9 80.141 116.1 <u>80.248</u> <u>117.0</u> 80.164 115.9	1.28 1.25 1.17 1.26 1.24		9.3-9.4	1 1 1 1 4	Heintz, 1973: -2°.1, -0."01

MICROMETER MEASURES OF DOUBLE STARS

ADS	DISC IDC	Mult.	Epoch 1900+	P	ρ	m	n	Notes
7203	STF 09016N732	1306 AB	80.128 80.248 <u>80.890</u> 80.422	0°0 3.0 0.2 1.1	2°95 2.90 3.07 2.97	5.0-8.2 1 1 3	1 1 1 Baize, 1948: -3°5, -0".24	
7284	STF 09120N2860	3121	80.895	189.0	0.40	8.1-8.1	1	Bos, 1938: -1°0, +0".12
7307	STF 09147N3837	1338 AB	80.248 80.890 <u>80.895</u> 80.678	255.2 255.7 258.6 256.5	1.12 0.93 0.93 0.99	6.6-6.8 1 1 3	1 1 1 Starikova, 1966: -1°1, +0".8	
7685	STF 10075N2755	213	80.895	127.6	0.66	8.4-10.4	1	Heintz, 1962: +0°6, -0".18
8032	A 10576N5464	1590	80.128	1.16	9.2-9.7		1	Heintz, 1963: +2°7, -0".06
8680	HU 12458N2065	640	80.388	147.1	0.64	10.1-10.1	1	Baize, 1973: +7°3, -0".11
9229	STF 14166N4858	1834	80.388 <u>80.574</u> 80.481	101.1 101.9 101.5	1.26 1.16 1.21	7.9-8.0 1 2	1 1 Bos, 1936: -2°5, -0".02	
9324	A 14334N4819	347	80.388 <u>80.574</u> 80.481	276.8 273.6 275.2	0.58 0.56 0.56	8.5-8.7 1 2	1 1 Guntzel-Lingner, 1956: -1°9, -0".07	
9343	STF 14364N1369	1865 AB	80.577	306.1	1.12	4.4-4.8	1	Wierzbinski, 1956: +1°4, -0".06
9380	STF 14414N0965	1879 AB	80.484 80.577 80.530	90.4 91.7 91.0	1.39 1.39 1.39	7.6-8.6 1 2	1 1 Wierzbinski, 1956: +1°2, -0".12	
9425	STT 14487N1567	288	80.498 80.577 80.538	170.6 173.8 172.2	1.35 1.26 1.30	6.9-7.6 1 2	1 1 Heintz, 1956: +1°3, +0".06	
9578	STF 15140N2672	1932	80.498 80.568 80.533	249.9 250.5 250.2	1.27 1.27 1.27	7.1-7.6 1 2	1 1 Heintz, 1965: -0°4, -0".13	
9617	STF 15191N3039	1937	80.549 80.574 80.562	328.1 329.1 328.6	0.56 0.55 0.55	5.6-6.1 1 2	1 1 Danjon, 1938: -2°0, +0".12	
9626	STF 15207N3742	1938 BC	80.549 80.552 80.551	13.2 14.6 13.9	2.04 2.00 2.02	7.2-7.8 1 2	1 1 Baize, 1952: -1°7, -0".16	
9716	STT 15325N3968	298	80.549 80.574 <u>80.724</u> 80.616	217.3 218.9 220.0 218.7	0.58 0.58 0.72 0.63	7.4-7.7 1 1 3	1 1 1 Couteau, 1966: +4°0, -0".06	
9982	STF 16111N0737	2026	80.582	24.7	2.74	9.1-9.6	1	Heintz, 1963: +1°2, -0".17
10075	STF 16245N1837	2052	80.582 80.708 80.645	136.2 132.9 134.5	1.19 1.26 1.23	7.8-7.8 1 2	1 1 Siegrist, 1952: +0°2, -0".21	
10188	D 16408N4340	15	80.568 80.582 80.575	140.2 144.7 142.5	1.16 1.10 1.13	9.1-9.1 1 2	1 1 Wierzbinski, 1955: +0°8, -0".03	

ADS	DISC IDC	Mult.	Epoch 1900+	P	ρ	m	n	Notes
10235	STF 2107 16479N2850		80.552 80.569 80.585 <u>80.713</u> 80.605	89.0 88.0 84.2 <u>85.2</u> 86.6	1".15 1.17 1.20 <u>1.18</u> 1.18	6.7-8.2	1 1 1 1 4	Rabe, 1927: -1°5, -0."19
10279	STF 2118 16559N6511		80.552 80.577 80.713 <u>80.719</u> 80.640	69.2 70.1 70.9 <u>70.2</u> 70.1	1.15 1.02 0.99 <u>1.01</u> 1.04	6.7-7.4	1 1 1 1 4	Giannuzzi, 1956: -1°5, -0."24
10341	BU 823 AB 17015N0047		80.549 80.574 <u>80.656</u> 80.593	125.2 128.5 <u>119.4</u> 124.4	0.82 0.73 <u>0.74</u> 0.76	8.7-9.7	1 1 1 3	Arend, 1956: +5°7, -0."24
10345	STF 2130 AB 17033N5436		80.552 <u>80.577</u> 80.565	44.3 <u>47.4</u> 45.8	2.10 <u>1.94</u> 2.02	5.8-5.8	1 1 2	Heintz, 1966: +4°6, +0."11
+45°2505	Kui 79 17092N4551		80.724	271.4	0.87	10.1-10.6	1	Baize, 1952: -1°9, +0."07
10598	STF 912 17252S0059		80.549 <u>80.574</u> 80.562	344.1 <u>347.8</u> 345.9	0.60 0.72 0.66	5.9-6.2	1 1 2	Wilson, 1976: -3°6, +0."09
11010	BU 1127 17596N4414		80.724	76.8	0.83	7.5-9.3	1	Popović, 1970: -2°6, -0."24
11046	STF 2272 AB 18004N0232		80.498 <u>80.577</u> 80.538	321.2 <u>320.2</u> 320.7	2.14 <u>2.14</u> 2.14	4.1-6.3	1 1 2	Heintz, 1973: +0°3, -0."10
11324	AC 11 AB 18198S0138		80.550 <u>80.574</u> 80.562	357.7 <u>353.1</u> 355.4	0.82 0.73 0.78	6.8-7.0	1 1 2	Heintz, 1958: -0°8, -0."0
11334	STF 2315 AB 18210N2720		80.719 <u>80.722</u> 80.720	131.0 <u>130.6</u> 130.8	0.59 0.57 0.58	6.6-7.6	1 1 2	Heintz, 1960: +1°7, -0."11
11483	STT 358 18314N1654		80.574 80.569 80.582 <u>80.656</u> 80.588	165.2 163.0 163.5 <u>162.8</u> 163.6	1.57 1.56 1.64 <u>1.77</u> 1.64	6.8-7.2	1 1 1 1 4	Starikova, 1966: +3°0, +0."09 Hopmann, 1970: +1°9, +0."08
11479	STT 359 18314N2331		80.719	12.7	0.59	6.4-6.7	1	Symms, 1964: +3°1, -0."02
11568	STF 2384 AB 18385N6702		80.713 <u>80.722</u> 80.717	311.5 <u>311.8</u> 311.7	0.61 0.57 0.59	8.6-9.1	1 1 2	Heintz, 1975: +0°1, -0."01
11623	A 253 18400N3135		80.514 80.577 <u>80.719</u> 80.603	117.1 128.1 <u>120.4</u> 121.9	0.64 0.64 <u>0.63</u> 0.64	9.1-10.1	1 1 1 3	Muller, 1956: -0°2, -0."12
11635	STF 2382 AB 18410N3934		80.569	353.1	5.1-6.1		1	Guntzel-Lingner, 1956: -1°9, -0."07
11635	STF 2383 CD 18410N3934		80.547	90.4	2.41	5.7-6.0	1	Guntzel-Lingner, 1956: +7°1, +0."07
11722	STF 2402 18450N1034		80.730	208.6	1.18	8.0-8.4	1	Changed 10° since 1830

MICROMETER MEASURES OF DOUBLE STARS

ADS	DISC TDC	Mult.	Epoch 1900+	P	ρ	m	n	Notes
11/91	A 93 18484S0540		80.550 <u>80.577</u> 80.563	141.8 <u>141.8</u> 141.8	0.59 <u>0.64</u> 0.62	9.4-10.0	1 <u>1</u> 2	Heintz, 1973: -2°.0, -0".07
11811	BU 137 AB 18505N3715		80.724	156.2	1.35	8.3-8.7	1	Changed 32° since 1875
11897	STF 2438 18558N5805		80.574 80.577 <u>80.713</u> 80.621	5.4 0.9 2.7 3.0	0.74 0.78 <u>0.79</u> 0.77	6.8-7.4	1 <u>1</u> <u>1</u> 3	Jastrzebski, 1959: +0°.8, -0".11
11956	STF 2437 18575N1902		80.722 <u>80.730</u> 80.726	28.8 <u>28.8</u> 28.8	0.58 <u>0.56</u> 0.57		1 <u>1</u> 2	Changed 52° since 1830
12447	STF 2525 19225N2707		80.547 80.569 80.675 80.675 80.713 <u>80.790</u> 80.659	292.5 295.4 294.0 294.0 294.3 <u>295.1</u> 294.3	1.75 1.68 1.70 1.70 1.68 <u>1.70</u> 1.70	8.5-8.7	1 <u>1</u> <u>1</u> <u>1</u> <u>1</u> <u>1</u> 5	Tamburini, 1969: +1°.5, -0".14
12667	STT 377 AB 19326N3526		80.577 <u>80.708</u> 80.642	37.3 <u>33.8</u> 35.5	0.84 <u>0.85</u> 0.85	8.4-8.5	1 <u>1</u> 2	Changed 16° since 1842
12889	STF 2576 AB 19418N3322		80.498 80.547 80.574 80.675 80.708 <u>80.790</u> 80.632	359.0 357.3 357.1 357.8 356.7 <u>358.2</u> 357.7	2.04 1.94 1.93 1.95 1.95 <u>1.97</u> 1.96	9.3-9.3	1 <u>1</u> <u>1</u> <u>1</u> <u>1</u> <u>1</u> 6	Rabe, 1948: +1°.7, -0".09
12930	HU 758 19432N3307		80.675	144.5	0.92	9.5-9.6	1	No changed since 1904
12972	STT 387 19450N3504		80.547 80.550 80.656 80.722 <u>80.725</u> 80.646	167.5 165.4 164.2 167.5 <u>167.0</u> 166.3	0.75 0.58 0.69 0.57 <u>0.61</u> 0.65	6.9-7.9	1 <u>1</u> <u>1</u> <u>1</u> <u>1</u> 5	Baize, 1961: -0°.3, +0".05
13665	A 1205 20141N2854		80.719 <u>80.828</u> 80.774	102.4 <u>100.1</u> 101.2	0.48 <u>0.53</u> .0.51	9.0-9.6	1 <u>1</u> 2	Heintz, 1978: -0°.4, -0".06
13723	STT 406 20166N4503		80.719	118.2	0.52	7.4-8.3	1	Heintz, 1975: +1°.3, -0".07
13750	STF 2672 20173N2327		80.725 <u>80.825</u> 80.775	327.6 <u>327.1</u> 327.3	0.72 <u>0.72</u> 0.72	8.7-8.8	1 <u>1</u> 2	Changed 49° since 1831
14238	BU 64 AB 20403N1222		80.713	160.9	0.53	9.1-9.3	1	Baize, 1957: -2°.7, -0".02
14296	STT 413 AB 20435N3607		80.547 80.569 <u>80.574</u> 80.563	16.2 17.5 <u>16.1</u> 16.6	0.80 0.83 <u>0.78</u> 0.80	4.8-6.1	1 <u>1</u> <u>1</u> 3	Rabe, 1948: +1°.3, -0".04
14424	BU 367 AB 20508N2743		80.719 <u>80.828</u> 80.773	115.5 <u>118.2</u> 116.8	0.44 <u>0.45</u> 0.45	8.6-9.0	1 <u>1</u> 2	Heintz, 1962: -2°.4, -0".05

ADS	DISC IDC	Mult.	Epoch 1900+	P	ρ	m	n	Notes
14499	STF 2737 AB 20541N0355	80.498 <u>80.547</u> 80.522	286.0 <u>288.0</u> 287.0	0.96 <u>0.97</u> 0.97	6.0-6.3	1 <u>1</u>		
						2		Zeller, 1965: +1°.6, -0".07
14558	STF 2746 20580N3852	80.577 80.708 80.725 <u>80.828</u> 80.709	315.9 316.5 314.9 <u>315.5</u> 315.7	0.97 0.97 0.94 <u>1.02</u> 0.97	8.0-8.6	1 1 1 <u>1</u>		
						4		Changed 40° since 1830
14562	LV 9 20582N3850	80.708 80.790 <u>80.828</u> 80.775	192.0 188.7 <u>193.5</u> 191.4	2.24 2.40 <u>2.16</u> 2.27	9.8-10.8	1 1 <u>1</u>		
						3		No changed since 1896
14573	STF 2744 20580N0108	80.498 80.569 <u>80.713</u> 80.660	129.6 127.2 <u>130.2</u> 129.0	1.17 1.22 <u>1.21</u> 1.20	7.0-7.5	1 1 <u>1</u>		
						3		Popović, 1964: +3°.9, -0".07
14766	A 884 21098N4630	80.722	140.5	0.38	6.6-8.7	1		Changed 45° since 1904
14783	H 48 21117N6400	80.550 80.569 80.656 <u>80.660</u> 80.608	255.7 252.0 253.5 <u>253.7</u> 253.7	0.56 0.71 0.72 <u>0.68</u> 0.67	7.1-7.3	1 1 1 <u>1</u>		
						4		Baize, 1950: +2°.6, 0".00
14957	A 1220 21209N3103	80.676 <u>80.891</u> 80.783	155.6 <u>155.0</u> 155.3	1.34 <u>1.38</u> 1.36	8.5-9.0	1 <u>1</u>		
						2		No changed since 1905
15267	HO 166 21394N2723	80.719 <u>80.828</u> 80.773	91.6 <u>91.6</u> 91.6	0.39 <u>0.41</u> 0.40	8.8-8.8	1 <u>1</u>		
						2		Couteau 1958: -3°.0, +0".02
15270	STF 2822 AB 21397N2817	80.569 80.574 80.660 80.665 <u>80.675</u> 80.626	294.2 296.3 297.1 296.5 <u>296.9</u> 296.2	1.84 1.82 1.81 1.81 <u>1.86</u> 1.83	4.7-6.1	1 1 1 1 <u>1</u>		
						5		Heintz, 1966: -2°.1, +0".05
15447	BU 75 AB 21506N1025	80.902	350.1	0.44	8.4-8.9	1		Baize, 1973: +14°.2, +0".14
15843	BU 1216 22156N2901	80.722 <u>80.725</u> 80.723	284.6 <u>285.5</u> 285.0	0.67 <u>0.66</u> 0.66	8.4-8.7	1 <u>1</u>		
						2		Changed 33° since 1890
15972	KRU 60 22244N5712	80.708 80.790 <u>80.828</u> 80.775	175.5 174.6 <u>173.5</u> 174.6	2.44 2.39 <u>2.51</u> 2.45	9.4-10.9	1 1 <u>1</u>		
						3		Lippincott, 1953: +0°.9, -0".25 Wielen, 1962: +2°.6, -0".27
15988	STF 2912 AB 22249N0355	80.902	121.4	0.74	5.8-7.2	1		Knipe, 1960: +3°.9, -0".28
16057	STF 2924 AB 22301N6923	80.881 <u>80.897</u> 80.889	91.4 <u>85.4</u> 88.4	0.64 <u>0.58</u> 0.61	6.5-7.0	1 <u>1</u>		
						2		Heintz, 1956: +2°.2, +0".08
16185	STF 2934 22370N2054	80.547 80.574 80.713 <u>80.875</u> 80.677	72.8 74.7 72.4 <u>73.1</u> 73.3	0.95 0.87 0.91 <u>0.96</u> 0.92	8.7-9.7	1 1 1 <u>1</u>		
						4		Heintz, 1962: +3°.7, -0".03

MICROMETER MEASURES OF DOUBLE STARS

ADS	DISC IDC	Mult.	Epoch 1900+	P	ρ	m	n	Notes
16326	A 632 22480N5712		80.569 80.574 80.722 80.725 <u>80.875</u> 80.693	169.1 170.7 170.3 161.2 <u>169.3</u> 168.1	0.87 0.83 0.92 0.94 0.90 0.89	8.2-9.0	1 1 1 1 1 5	Heintz, 1962: -0°.4, +0".07
16345	BU 382 AB 22492N4413		80.711 80.829 <u>80.881</u> 80.807	208.2 204.1 <u>205.5</u> 205.9	0.75 0.87 0.90 0.84	6.5-8.5	1 1 1 1	Muller, 1954: -0°.2, -0".11
16373	HU 987 22508N1515		80.678 <u>80.829</u> 80.753	95.0 <u>91.0</u> 93.0	0.71 0.74 0.72	9.1-9.3	1 1 2	Heintz, 1966: +5°.7, +0".06
16428	STF 483 22542N1117		80.547 <u>80.574</u> 80.562	292.7 <u>296.8</u> 295.7	0.72 0.63 0.67	6.0-7.5	1 1 2	Guntzel-Lingner, 1956: -6°.0, -0".01
16626	ES 1039 23105N5248		80.897	211.0	1.15	10.0-10.5	1	No changed since 1910
16665	BU 80 23138N0452		80.569 80.574 80.678 <u>80.902</u> 80.681	312.0 309.0 312.2 <u>311.7</u> 311.3	0.82 0.83 0.89 0.85 0.85	9.0-10.0	1 1 1 1 4	Couteau, 1960: -6°.9, +0".01
16785	A 1487 23244N4009		80.719 <u>80.728</u> 80.723	166.3 <u>165.1</u> 165.7	0.89 0.88 0.89	8.7-9.2	1 1 2	Changed 23° since 1906
16877	STT 500 AB 23326N4353		80.897	358.9	0.57	6.3-7.2	1	Changed 60° since 1845
16928	BU 858 AB 23363N3201		80.719 <u>80.728</u> 80.723	232.4 <u>231.1</u> 231.7	0.72 0.77 0.75	7.4-8.9	1 1 2	Changed 55° since 1881
16951	A 1242 23380N1117		80.719 <u>80.829</u> 80.773	323.7 <u>326.3</u> 325.0	0.77 0.77 0.77	9.0-9.0	1 1 2	Zulević, 1977: -0°.6, +0".02
17020	STT 507 AB 23438N6420		80.719 <u>80.829</u> 80.774	308.6 <u>304.1</u> 306.3	0.71 0.73 0.72	6.8-7.5	1 1 2	Zulević, 1977: +2°.1, +0".002
17149	STF 3050 23544N3310		80.569 80.574 <u>80.711</u> 80.618	308.3 309.6 <u>309.3</u> 309.1	1.54 1.52 1.56 1.54	6.6-6.6	1 1 1 3	Heintz, 1973: -0°.8, -0".01
17178	HLD 60 23563N3905		80.711 <u>80.728</u> 80.720	180.6 <u>180.5</u> 180.6	1.00 0.92 0.96	9.2-9.6	1 1 2	Heintz, 1963: -1°.6, -0".09
34	BU 862 23596N3737		80.891	8.1	0.43	10.1-10.2	1	Morel, 1969: -8°.8, 0".00
	GP 5 GP 39		80.492 80.678	138.1 97.8	0.68 0.48	10.8-10.9	1 1	

MIKROMETER MEASURES OF DOUBLE STARS

(Series 34)

G. M. Popović

(Received 30 March 1981)

SUMMARY: Presented are 280 measures of 156 double and multiple systems of stars.

Presented are 280 measures of 156 systems respectively 170 pairs made with Zeiss 65/1055 cm refractor of the Belgrade Observatory. Among them there are 66 systems with the designation GP. This is the 34 th Belgrade series of the double stars measures and, at the sa-

me time, a continuation of my measures published as Series 30 (G. M. Popović, 1979). The structure of the measured pairs according to ρ is given in Table I.

Table II lists the measurements in the form identical to that of my previous series (G. M. Popović, 1979).

Table I

$\rho < 0.^{\circ}50$	$0.^{\circ}50 < \rho < 1.^{\circ}00$	$1.^{\circ}00 < \rho < 2.^{\circ}00$	$\rho > 2.^{\circ}00$	Σ
9m 3.2%	68m 24.3%	77m 27.5%	126m 45.0%	280m 100%

Table II
Micrometer measures of double stars

ADS m	Disc 1900-2000 Mult.	Epoch 1900+	θ	ρ	m	w	Notes
202 00106-159N5200-33 10.9-11.1	Es 865 AB	80.709 80.810 80.760	79°0 75.6 77.3	1°68 1.59 1.84	0.3 0.3. 0.3	1+2 1+2 2n	If Espin's measure of the position angle in 1910 is 306° as reported in IDS, and not 206° as reported in ADS, then the pair is undoubtedly optical one.

MICROMETER MEASURES OF DOUBLE STARS

ADS m	Disc Mult.	1900-2000	Epoch 1900+	θ	ρ	m	w	Notes
220	Σ 19	80.791	135.8	2.00	9.0-11.0	1+1		
00115-167N3604-37		80.793	139.5	2.01	8.0-10.0	1+1		
7.1-9.6.		80.815	138.3	2.06	8.0-10.0	1+2		
		80.828	137.9	2.20	8.5-11.0	2+2		
		80.811	137.9	2.09	8.4-10.6	4n		
221	0Σ 4	80.829	172.4	0.97	9.0-9.5	3+2	Muller, 1957: + 1°.7, + 0".43	
00115-167N3556-89								
8.2-8.9								
256	h 1015	80.708	304.4	5.12	9.0-9.5	1+2	Bos (rectilinear trajectory), 1959:	
00135-187N2512-45		80.741	303.0	5.53	0.8	2+2	+0°.5, -0".16	
9.7-10.4	AB	80.889	302.0	5.41	1.0	2+2		
		80.786	303.0	5.37	0.8	3n		
	AC	80.708	255.0	26.8	9.0-11.0	1+2		
-	GP 35	80.651	295.8	0.62	9.5-10.5	3+3	Up to now measurements indicate retrograde motion.	
00143-195N3511-44		80.889	295.4	0.56	1.0	2+2	GP 35 = BD + 34°33 (9 ^m)	
9.5-10.7 (6n)		80.746	295.6	0.60	1.0	2n		
-	GP 36	80.651	111.7	0.54	10.0-10.1	2+2	Direct motion.	
00145-196N3428-61							GP 36 = BD + 34°34 (9 ^m)	
9.8-10.1 (3n)								
399	A 909	79.690	38.6	0.84	0.7	2+2		
00242-297N5822-55		79.701	39.7	0.96	6.8-9.2	1+1		
9.5-9.8		79.694	39.0	0.88	0.8	2n		
-	GP 38	80.657	97.3	2.39	9.5-14.0	2+2	GP 38 = BD + 34°72 (9 ^m)	
00267-320N3504-38		80.725	96.9	2.47	9.3-13.0	1+1		
9.2-13.2 (6n)		80.680	97.2	2.42	9.4-13.7	2n		
-	GP 52	80.657	80.5	98.6	10.0-10.5	1+2		
00268-340N3512-46								
10.0-10.2 (1/5n)AB								
10.2-11.4 (5n)	BC	80.657	320.2	8.00	10.5-12.0	1+2		
		80.725	323.6	7.54	10.5-12.0	1+2		
		80.758	321.5	7.69	10.0-11.0	1+1		
		80.708	321.8	7.75	10.3-11.7	3n		
	AD	80.758	129	63.9		1+1		
473	h 1036	79.673	84.8	6.83	9.3-9.8	1+2		
00292-346N4220-53		80.900	86.4	7.30	9.0-9.5	2+1	Quadrant certain.	
10.7-11.3		80.286	85.8	7.06	9.1-9.6	2n	My measurement does not respond to expected position. It would be useful to establish the cause of this.	
673	β 495	79.690	207.1	0.37	0.0	2+1	The angle has decreased by 24° since 1878 with decrease in distance.	
00435-488N1809-42								
8.4-8.4								
692	β 781	79.690	23.4	0.92	0.6	2+2	About 9° retrograde motion since 1881.	
00453-516N6827-59		79.701	20.7	0.78	8.5-9.5	1+2		
8.7-9.2		79.695	22.2	0.88	0.8	2n		
710	Σ 65	79.701	38.3	2.86	0.0	1+2		
00464-528N6819-51		80.725	40.1	3.02	0.0	1+1		
8.0-8.0		80.111	39.0	2.92	0.0	2n	Change in both coordinates uncertain.	
-	GP 46	80.714	222.3	28.2	9.0-12.5	1+1	GP 46 = BD + 34°189	
01027-082N3502.34								
9.0-11.4 (1/4n) AB								
11.4-12.8 (4n)	BC	80.714	151.8	3.74	12.5-13.0	1+1		
-	GP 47	80.714	217.6	3.86	12.8-14.0	1+1		
01029-085N3503-35								
11.9-13.5 (3n)								
-	Stein 1562	80.709	89.1	3.17	9.0-11.0	1+1	Change in θ uncertain since 1913. Decrease in distance.	
01132-194N5746-78								
9.8-11.6								

ADS m	Disc Mult.	1900-2000 Epoch 1900+	ϵ	ρ	m	w	Notes	
1133	Ho 310 01200-256N2801-32 10.0-10.2	80.725 80.900 80.830	360.1 359.7 359.9	1.54 1.72 1.65	0.2 0.2 0.2	1+1 2+1 2n	Very slow direct motion.	
1370	D 3 01384-450N5641-71 9.7-11.2	80.897	333.1	2.97	1.5	3+2	Unchanged in 103 years.	
1371	β 453 01384-450N5637-67 10.1-10.4	80.897	68.2	0.42	10.0-10.3	3+2	Baize, 1973: +9°.8, -0".16	
1438	Σ 162 01430-492N4724-54 6.5-7.0 AB	80.957	205.5	1.69	7.0-7.5	1+2	The angle has decreased by 19° since 1836.	
	6.1-8.4	AC	80.957	178.8	20.6	7.0-9.5	1+2	The component D is brighter 0.2 than C.
1560	S 404 01520-580N4054-83 7.6-9.6	80.709 80.741 80.900 80.798	81.8 81.2 82.2 81.7	27.69 27.65 27.77 27.71	- - - 3n	1+1 2+2 2+2 3n	An optical pair. Herschel's measure in 1783 does not agree with the rectilinear trajectory.	
1709	Σ 228 02076-141N4701-29 6.4-7.3	79.668 80.714 80.741 80.881 80.562	266.4 267.8 268.7 267.0 267.8	1.04 1.01 0.97 1.01 1.00	1.0 0.4 0.7 9.0-9.3 0.6	1+1 2+2 1+2 1+1 4n	Heintz, 1952: +1°.2, -0".05	
1801	β 8 02160-213N0825-52 8.2-9.1	80.741 80.900 80.809	220.6 220.3 220.5	1.23 1.35 1.28	0.7 9.0-10.5 1.0	2+2 1+2 2n	The angle has increased by 20° since 1875	
1821	Ho 313 02184-233S0777-50 9.3-9.7	80.741 80.791 80.768	74.3 73.9 74.1	2.02 2.03 2.02	0.5 8.7-9.2 0.5	1+1 1+1 2n		
1859	A 658 02216-279N4102-29 9.3-10.9	80.889	212.1	2.55	-	3+2		
2034	$\Omega\Sigma$ 43 02349-407N2612-38 8.3-9.9	80.900	10.5	0.88	8.2-9.0	2+2	Heintz, 1961: +2°.7, -0".14	
2436	$\Omega\Sigma$ 52 03088-175N6517-39 6.9-7.5 AB	80.892	74.2	0.38	0.1	3+3	Heintz, 1962: +6°.2, -0".06	
	GP 83 03291-354N3508-28 8.0-8.8 (7n)	80.076 80.111 80.127 80.105	261.5 268.0 266.3 265.3	0.65 0.66 0.74 0.68	0.7 8.5-9.5 8.0-9.5 1.1	1+1 1+1 1+1 3n	GP 83 = COU 1080 = BD + 34°685 (8.0)	
2656	A 2420 03328-385N1716-36 8.7-9.7	80.834	264.0	1.95	1.0	1+1		
2800	Es 1137 03438-511N4922-51 10.2-10.3	80.714	119.2	0.49	0.2	3+3	Since 1912 the angle increased by 42° with decrease in distance. My measure indicate the orbital motion.	
2801	Σ 457 03444-504N2222-41 9.5-9.5 AB	80.714	92.4	0.93	0.0	3+3	The angle has decreased by 13° since 1831	
3082	$\Omega\Sigma$ 77 04096-159N3127-42 8.2-8.2 AB	80.892 80.903 80.898	268.8 270.7 269.8	0.66 0.73 0.70	0.0 8.0-8.0 0.0	2+2 2+2 2n	Muller, 1955: -0°.7, -0".06	
3174	Σ 535 04178-233N1109-23 7.1-8.6	80.881 80.889 80.892 80.886	291.0 292.6 291.9 291.9	1.19 1.03 1.10 1.10	7.0-9.0 2.0 7.0-8.5 2.4	1+2 2+2 3+3 3n	Popović, 1981: +0°.2, -0".05	

MICROMETER MEASURES OF DOUBLE STARS

ADS m	Disc 1900-2000 Mult.	Epoch 1900+	θ	ρ	m	w	Notes
3329	0Σ 86	80.892	15.7	0.88	0.0	3+2	The angle has decreased by 62° since 1845.
04307-366N1933-45		80.897	17.8	0.54	0.0	2+2	
8.2-8.2		80.894	16.6	0.73	0.0	2n	
3580	Es 14	79.022	156.0	33.2	8.0-9.5	1+2	
04527-598N4310-19							
9.5-9.5-9.0	ABxC						
9.0-11.8	CD	79.022	279.2	3.18	9.5-11.0	1+1	
		79.934	282.7	4.47	10.0-13.0	2+2	
		79.630	281.5	4.04	9.8-12.3	2n	
-	GP 70	80.221	7.8	0.98	9.5-9.8	2+2	All measures of the position angles discordant. GP 70 = BD + $35^\circ 1056$ ($9^m 5$)
05136-202N3555-63							
9.7-10.0 (4n)							
-	GP 69	80.076	298.4	3.19	1.0	1+1	
05148-215N3509-15		80.221	300.5	3.73	10.0-11.5	1+2	
10.4-11.6 (9n)		80.163	299.7	3.51	1.2	2n	
-	GP 107	79.937	141.5	2.90	13.0-13.4	1+2	
05180-248N3610-17							
12.2-13.1 (3n)							
4032	Ho 226	80.892	259.0	0.74	0.1	2+2	The angle has increased by 29° since 1887.
05208-271N2731-36		80.897	258.2	0.74	0.1	2+2	
8.6-8.6		80.894	258.6	0.74	0.1	2n	
-	GP 14	80.100	117.3	4.48	0.0	1+1	
05583-050N3504-05		80.128	124.6	5.94	14.0-14.2	1+1	
12.5-12.8		80.114	121.0	5.21	0.1	2n	At the limit of visibility.
4950	Σ 881	80.898	135.4	0.75	1.2	1+2	The angle has increased by 46° since 1830.
06132-221N5925-23							
6.2-7.7 AB							
5026	Es 1534	79.186	92.0	1.53	9.5-10.2	2+2	
06184-255N4326.23		79.934	92.5	1.74	0.8	2+2	
10.8-11.1		79.560	92.2	1.64	0.8	2n	
5054	1191	80.898	309.2	2.07	7.0-14.0	3+2	This measure does not offer a convincing argument the pair has to be optical, though there are the elements for this assertion. My measure of the distance in this case would have to be about $0.^m 7$ wider.
06203-262N1849-46							
6.9-13.7							
5081	J 910	79.186	331.9	1.45	8.5-9.5	3+3	Retrograde motion of about 6° since 1912
06213-285N4308-05							
10.0-10.4 AB							
5261	A 2451	79.202	140.6	0.49	9.0-9.0	1+2	Unchanged in 68 years.
06312-383N4207-02		79.934	144.6	0.58	0.0	2+2	
9.4-9.4		79.620	142.9	0.54	0.0	2n	
5269	Σ 941	79.202	82.3	1.66	7.5-8.7	1+2	
06316-387N4140-35		79.260	82.9	1.71	7.5-9.0	1+2	
7.2-8.2 AB		79.934	81.7	1.75	8.0-9.0	1+2	
		79.465	82.3	1.71	7.7-8.9	3n	
5278	Mlb 63	79.202	259.5	5.28	9.7-10.2	1+2	Unchanged in 61 years.
06320-391N4225-20							
9.2-9.8							
5290	H Σ -	80.892	282.9	0.68	0.0	1+2	
06337-392N0944-39		80.898	281.4	0.73	0.0	2+2	
8.6-8.6		80.895	282.0	0.71	0.0	2n	
5831	Σ 1024	80.898	315.0	1.26	0.3	3+3	
07034-103N3817-06							
9.0-9.5 AB							
-	GP 84	80.128	188.7	1.79	0.0	1+2	GP 84=BD + $35^\circ 1573$ ($9^m 5$)
07060-127N3522-13							
9.5-9.8 (1n)							

ADS m	Disc 1900-2000 Mult.	Epoch 1900+	ϵ	ρ	m	w	Notes
-	GP 74	80.084	148.7	8.92	10.0-12.0	1+1	
U7174-240N3439-28		80.111	148.2	8.92	10.0-11.0	1+1	
9.6-11.2 (10 n)		80.975	148.4	8.92	10.0-11.5	2n	
-	GP 105	80.240	49.1	0.51	10.0-10.2	3+3	GP 105 = BD + 36°1643 (9 ^m .2)
07259-325N3555-42							
9.8-10.0 (3n)							
6194	Ku 30	79.934	108.9	2.90	1.0	1+2	
07295-360N3433-20							
9.4-9.8							
6875	0Σ 189	79.180	291.5	3.41	8.0-14.0	1+1	
08080-149N4320-02		79.186	293.1	4.44	7.0-12.0	2+2	
		79.184	292.6	4.10	7.3-12.7	2n	
-	GP 111	80.240	51.3	0.76	10.0-10.1	3+2	GP 111 = BD + 36°1771 (9 ^m .1)
08084-149N3553-35							
10.0-10.2 (3n)							
6871	β 1244	80.898	17.2	1.11	8.0-9.0	2+2	The angle has decreased by 33° since 1891.
08086-138N0177-59							
8.3-8.5							
-	GP 97	79.142	276.5	0.72	0.0	1+1	GP 97 = BD + 34°1888 (9 ^m .3)
08372-435N3363-41		70.175	280.3	0.70	10.0-10.5	3+2	
10.0-10.5 (1n)		79.279	276.2	0.71	0.0	2+2	
		79.207	278.1	0.71	0.2	3n	
6963	A 2472	80.898	81.5	0.67	0.0	2+2	The angle has decreased by 15° since 1912.
08388-444N1577-55							
9.2-9.2							
7067	Σ 1280	79.142	111.5	1.36	0.1	2+2	Heintz, 1973: -2°.7, +0".08
08460-557N7071-48		80.221	116.7	1.44	0.1	2+2	Heintz, 1973: -1°.5, +0".20
9.3-9.4	AB						The brighter component is red and fainter blue.
-	GP 106	80.240	66.8	1.21	0.3	1+2	
09158-219N3338-13		80.898	70.3	1.61	1.0	2+2	
9.6-10.3 (3n)		80.616	68.8	1.44	0.8	2n	
-	GP 115	80.221	265.5	1.79	10.5-11.5	2+2	There is the component C: θ ~ 16° m ~ 10 ^m 0 GP 115 = BD + 44°1943 (9 ^m .4)
09556-017N4356-27							
10.0-10.8 (2n)	AB						
-	GP 85	79.142	38.5	5.00	11.0-12.0	1+1	There are still two components: at ρ 15" and ρ 30" with magnitudes 14 ^m respectively 11 ^m
10054-113N3416-47							
10.9-12.3 (4n)	AB						
-	GP 116	79.279	243.8	0.42	0.0	3+2	GP 116 = BD + 43°1996 (9 ^m .2)
10057-120N4252-25		80.221	249.6	0.46	10.0-10.5	2+2	
9.4-9.6 (2n)		79.698	248.4	0.44	0.2	2n	
7704	0Σ 215	79.339	185.6	1.48	-	2+2	Wierzbinski, 1953: +2°.2, +0".10 Zaera, 1957: +4°.3, +0".10
10108-163N1774-44							
7.3-7.5							
-	GP 117	79.175	260.5	0.61	7.5-8.5	3+2	GP 117 = BD + 44°1972 (7 ^m .7)
10123-154N4416-01		79.271	262.8	0.48	7.8-9.5	3+2	
7.9-9.1 (4n)		79.279	260.4	0.53	8.0-9.0	2+2	
		79.239	261.3	0.54	7.7-9.0	3n	
-	GP 73	79.271	207.8	0.67	0.0	1+2	GP 73 = BD + 34°2186 (9 ^m .4)
10511-560N3351-18		79.279	206.1	0.68	0.1	2+2	
9.7-9.7 (3n)		79.276	206.8	0.68	0.1	2n	
-	GP 113	79.279	47.3	0.95	0.1	3+3	GP 113 = BD + 33°2119 (9 ^m .1)
11245-299N3336-03		80.221	48.0	0.94	9.2-9.2	3+3	
9.1-9.2 (3n)		79.750	47.6	0.94	0.1	2n	
8820	A 1606	80.249	21.4	1.18	0.0	1+2	Very slow retrograde motion.
13083-128N4062-30							
9.0-9.0							

MICROMETER MEASURES OF DOUBLE STARS

ADS m	Disc Mult.	1900-2000	Epoch 1900+	θ	ρ	m	w	Notes
-	GP 72	79.271	321.3	1.50	0.7	1+1		
13119-166N3464-32 9.7-10.8 (4n)		80.238 79.754	327.1 324.2	1.46 1.48	1.0 0.8	1+2 2n		
9031	Σ 1785	79.424	158.7	3.27	0.2	2+2		
13445-492N2689-59 7.9-8.2		80.238 79.831	159.7 159.2	3.35 3.31	0.2 0.2	2+2 2n		Stand, 1953: $-0^{\circ}1$, $-0.^{\prime\prime}08$
9182	Σ 1819	79.424	243.6	0.77	0.0	1+2	Baize, 1973: $+3^{\circ}6$, $-0.^{\prime\prime}05$	
14103-153N0336-08 7.7-7.8								
9229	Σ 1834	80.249	102.8	1.14	0.0	1+2	Van den Bos: $-1^{\circ}0$, $-0.^{\prime\prime}10$	
14166-203N4858-31 7.9-8.0								
9254	Σ 1837	79.424	283.1	1.09	1-2	1+2	About 45° retrograde motion since	
14193-247S1113-40 6.7-8.3		80.238 79.831	280.2 281.6	1.09 1.09	8.0-9.0 1.1	1+2 2n	1829 and decrease in distance.	
9265	Σ 1842	80.238	17.5	2.56	9.0-9.2	2+1		
14220- . N0368-41 9.3-9.3								
9343	Σ 1865	80.249	305.9	0.95	-0.1	1+2	Wierzbinski, 1953: $+1^{\circ}1$, $-0.^{\prime\prime}12$	
14364-411N1369-44 4.4-4.8	AB							
9982	Σ 2026	80.492	23.9	2.70	9.0-9.5	1+2	Heintz, 1962: $0^{\circ}3$, $-0.^{\prime\prime}10$	
16111-160N0737-22 9.1-9.6								
9988	A 2180	79.424	309.7	1.06	11.0-12.0	1+1		
16118-168N0132-17 14.0-14.5	BC							
-	GP 2	80.593	134.4	3.32	11.0-13.0	1+2		
16334-371N3363-50 11.0-12.9 (5n)								
10188	D 15	80.492	142.6	1.13	0.0	1+2		
16408-439N4340-29 9.1-9.1		80.596 80.561	141.4 141.9	1.07 1.10	9.0-9.0 0.0	2+2 2n		Wierzbinski, 1955: $+0^{\circ}2$, $-0.^{\prime\prime}06$
-	GP 5	80.492	139.6	0.54	10.0-10.3	1+2	GP 5 = BD + $34^{\circ}2834$ (9 ^m .3)	
16413-450N3366-53 9.7-10.0 (5n)								
10235	Σ 2107	80.492	87.7	1.15	6.0-8.0	1+2		
16479-519N2850-40 6.7-8.2	AB	80.596 80.561	87.2 87.4	1.24 1.20	8.0-9.0 7.3-8.7	2+2 2n		Rabe, 1926: $-0^{\circ}7$, $-0.^{\prime\prime}17$
10429	A 2984	79.692	357.2	0.86	3.0	2+2		
17114-165N0020-27 4.9-7.9								
-	GP 10	80.653	215.0	1.97	0.3	1+1		
17537-573N3542-41 10.5-10.7 (5n)	AB							The position of the system GP 10 related to BD + $35^{\circ}3112$ (7.0) = ADS 10932: $=-9^{\circ}$
-	GP 123	79.665	232.2	2.97	12.0-12.2	1+2		
17556-586N4231-31 11.2-11.4 (3n)		79.670 79.668	230.7 231.4	2.95 2.98	11.0-11.2 11.2-11.4	3+2 2n		
-	GP 78	80.596	249.5	13.74	9.7-10.0	1+2		
18078-113N3505-06 9.6-10.1 (4n)								Perhaps identical with ALI 140=IDS 180079N3505. According to IDS ALI 140 has not BD number and the magnitudes are: 13.0-13.8 GP 78 = BD + $35^{\circ}3173$ (9 ^m .5)

ADS m	Disc 1900-2000 Mult.	Epoch 1900+	θ	ρ	m	w	Notes
11280	Σ 2312	80.495	341.2	1.42	1.0	1+2	
18172-211N2817-20							
9.0-10.0							
-	GP 15	80.640	312.7	8.44	10.0-12.0	1+1	
18308-343N3458-63		80.651	312.0	7.42	10.0-13.0	2+1	
10.7-12.9 (5n)		80.647	312.3	7.83	10.0-12.7	2n	
11483	$\Omega\Sigma$ 358	80.495	164.3	1.76	-0.1	1+2	
18314-359N1654-59		80.653	165.7	1.85	0.0	1+2	Starikova, 1966: $+4^{\circ}4$, $+0.^{\circ}26$
6.8-7.2 AB		80.574	165.0	1.80	0.0	2n	Hopmann, 1970: $+3.^{\circ}3$, $+0.^{\circ}25$
-	GP 16	80.500	308.1	8.61	10.5-12.0	1+2	
18343-379N3500-05		80.593	308.4	7.94	10.0-13.0	1+2	
9.9-12.7 (7n)		80.596	308.3	8.20	10.0-13.0	1+2	
		80.563	308.3	8.25	10.0-12.6	3n	According to "Carte du Ciel" the position from 1939: $\theta \sim 291^{\circ}$, $\rho \sim 6"$ Direct motion.
-	GP 17	80.593	15.1	8.86	14.0-14.0	1+2	
18354-390N3456-62							
13.4-13.4 (4n)							
-	GP 45	80.593	276.2	6.28	12.0-12.0	1+2	
18389-425N3426-32		80.596	276.9	6.21	12.5-12.5	1+1	
12.1-12.1 (5n)		80.594	278.5	6.25	12.2-12.2	2n	
-	GP 43	80.500	273.2	2.90	14.5-15.0	1+1	At the limit of visibility.
18447-484N3425-32		80.593	274.5	2.89	12.5-14.0	1+2	
12.8-13.8 (6n)		80.547	273.8	2.90	13.5-14.5	2n	
-	GP 44	80.594	230.6	9.70	9.0-12.0	1+2	GP 44 = BD + $35^{\circ}3368$ (9^m5)
18459-495N3505-11		80.596	230.2	8.97	9.0-11.0	1+2	
8.9-10.6 (4n)		80.595	230.4	9.34	9.0-11.5	2n	
11863	Σ 3130	80.681	258.4	2.39	8.0-11.0	2+2	
18529-559N4406-13		80.711	258.9	2.23		1+1	
7.2-8.3-10.6 AB-C		80.691	258.6	2.34	8.0-11.0	2n	
-	GP 41	80.632	353.1	5.34	14.0-15.0	1+1	
18532-567N3438-45							
13.5-14.3 (2n)							
-	GP 130	80.632	125.4	55.5	9.5-12.0	1+2	GP 130 = BD + $34^{\circ}3369$ (9^m5)
18538-574N3459-65							
9.5-11.2 (1/4n)AB							
11.2-12.2 (4n) BC		80.632	271.9	8.61	12.0-13.0	1+2	
		80.727	270.9	8.28	11.0-12.0	3+2	
		80.691	271.3	8.40	11.2-12.2	2n	
-	GP 29	80.501	171.4	2.37	10.0-11.0	1+2	
18539-575N3446-54		80.632	174.3	2.28	10.0-11.0	1+2	
10.5-11.6 (10n)		80.727	172.6	2.63	12.0-13.0	3+2	
		80.639	172.7	2.46	11.2-11.9	3n	
11894	A 587	80.681	190.4	1.75	10.0-10.7	3+2	
18550-580N4358-66		80.711	190.5	1.59	10.0-10.8	1+2	
10.4-11.2 BC		80.692	190.4	1.69	10.0-10.7	2n	
-	GP 40	80.501	79.6	5.79	12.0-13.0	1+2	
18583-620N3427-36		80.632	75.0	6.76	11.0-12.5	1+1	
10.8-11.9 (4n) AB		80.563	77.8	6.18	11.5-12.8	2n	
12.3-12.3 CD		80.501	352.7	3.89	13.0-13.0	1+2	
		80.632	349.2	3.39	12.3-12.5	1+1	
		80.553	361.3	3.09	12.6-12.8	2n	
12040	Σ 2454	80.716	276.2	0.98	1.5	1+1	Baize, 1975: $-2^{\circ}6$, $-0.^{\circ}24$ (elipt. trajec.)
19023-062N3017-26							Olević, 1976: $-3^{\circ}4$, $-0.^{\circ}13$ (rectilin. trajec.)
8.5-9.7 AB							
-	GP 30	80.634	317.1	2.16	12.0-12.7	1+1	
19051-087N3412-21							
10.4-10.9 (4n)							

MICROMETER MEASURES OF DOUBLE STARS

ADS m	Disc Mult.	1900-2000 Epoch 1900+	θ	ρ	m	w	Notes
-	GP 34	79.660	52.2	2.43	-	1+1	
19252-289N3503-25		79.665	51.2	-	9.5-14.0	1+2	According to "Carte du Ciel" the
9.5-13.1 (11n) AB		79.670	52.0	3.08	9.5-13.0	3+3	position angle in 1925 is $\sim 135^\circ$.
		79.692	54.6	3.22	3.0	2+1	Retrograde motion.
		79.698	52.1	2.88	9.5-14.0	2+2	GP 34 = BD + $34^\circ 3568$ ($9^m 5$)
		79.678	52.3	3.07	3.8	5n	
-	GP 135	80.727	36.1	1.86	1.5	3+2	
19253-290N3456-78							
10.8-11.7 (5n)							
12521	Es 2240	79.698	151.7	2.35	12.0-12.1	1+1	About 8° direct motion since 1926.
19259-296N3441-54							
12.4-12.9 BC							
-	GP 96	80.727	111.1	1.19	0.1	3+2	
19327-363N3419-33							
11.0-11.4 (2n)							
-	GP 121	79.611	323.3	1.39	-0.3	1+2	GP 121 = BD + $36^\circ 3687$ ($9^m 2$)
19417-454N3641-56		79.660	324.1	1.34	0.3	1+1	
10.0-10.0 (1n)		79.631	323.4	1.37	0.0	2n	
12965	$\Omega \Sigma$ 386	80.640	71.7	0.99	0.2	2+2	
19446-482N3655-70		80.653	70.8	0.99	-0.1	1+2	
8.2-8.5		80.744	70.7	0.96	0.0	1+2	
		80.760	72.8	0.87	0.0	2+2	
		80.699	71.8	0.95	0.0	4n	Very slow retrograde motion.
-	SEI 695	80.793	347.2	8.31	10.0-10.3	1+2	
19469-507N3414-29		80.809	347.8	8.12	0.2	1+2	
10.0-10.7 AB		80.801	347.5	8.22	0.2	2n	The angle has increased by 10° since 1894
	AC	80.793	291.0	99.7	10.0-11.5	1+2	This is probably first measure of AC.
-	SEI 696	80.793	129.1	19.4	0.3	1+2	
19470-508N3415-30		80.809	129.3	19.5	0.1	2+2	
11.2-10.1 BA		80.802	129.2	19.4	0.2	2n	
13030	β 978	80.640	237.7	0.93	0.0	1+2	
19472-514N2316-31							
9.5-9.6							
-	GP 37	80.711	334.5	2.49	-	1+1	GP 37 = BD + $34^\circ 3771$ ($9^m 5$)
19494-532N3501-16		80.793	334.3	2.63	9.7-9.8	1+1	
9.7-9.9 (4n)		80.809	330.1	2.27	10.0-10.2	1+2	
		80.776	332.6	2.43	9.8-10.0	3n	
-	GP 138	80.760	35.2	0.45	9.6-9.6	3+2	GP 138 = BD + $45^\circ 3102$ ($9^m 1$)
20107-139N4521-39							
9.4-9.5 (2n)							
-	GP 128	80.815	307.8	4.27	12.0-12.5	1+2	
20152-190N3531-50							
12.0-12.4 (2n)							
-	GP 129	80.815	289.9	55.6	9.5-12.0	1+2	
20172-210N3506-25							
9.5-12.5 (1/2n) AB							
12.5-13.0 (2n) BC		80.815	33.8	5.50	12.0-12.5	1+2	
	AD	80.815	214.9	80.5	9.5-11.0	1+2	
	DE	80.815	259.9	13.10	11.0-13.0	1+2	
-	GP 144	80.651	122.8	5.63	13.0-13.5	1+1	
20190-223N4527-46							
13.0-13.5 (3n)							
-	GP 94	80.760	154.6	2.54	10.5-11.5	1+2	
20203-242N3437-56							
11.8-12.3 (3n)							

ADS m	Disc 1900-2000 Mult.	Epoch 1900+	θ	ρ	m	w	Notes
-	GP 80	80.760	334.1	2.59	0.5	2+2	
20204-243N3447-66 12.4-12.9 (3n)							
-	BRT 2263	79.785	30.7	3.49	11.0-11.3	1+2	
20236-268N4551-71 10.-10.6		79.791 79.788	30.7 30.7	3.65 3.58	11.0-11.2 11.0-11.2	2+2 2n	
14194	ARG 39	79.780	168.3	11.60	-	1+2	An optical pair.
20393-424N4854-76 8.7-9.3	AB						
14499	Σ 2737	79.668	286.1	0.85	8.1-8.3	2+2	
20541-591N0355-78 5.8-6.3	AB	80.716 80.117	284.9 285.6	1.01 0.92	0.1 0.2	1+2 2n	Van den Bos, 1932: +0°.1, -0".13
5.8-7.1	AC	79.668	66.9	10.34	8.1-10.0	2+2	
-	GP 82	80.634	23.6	4.44	9.0-11.0	1+2	
20544-583N3549-71 10.4-12.5 (4n)		80.651 80.657 80.645	26.1 26.2 25.1	4.55 4.40 4.46	10.0-13.0 12.0-14.0 10.3-12.7	1+1 1+1 3n	
-	GP 27	79.698	158.0	1.26	0.0	1+1	GP 27 = BD + 34°4228 (9m)
20552-593N3412-35 10.0-10.7 (2n)		80.651 80.174	161.0 159.5	1.36 1.31	11.0-11.5 0.2	1+1 2n	
-	GP 28	80.634	0.8	2.68	12.5-12.7	1+2	The position of the pair related to
20553-593N3407-30 12.6-13.1 (4n)		80.651 80.642	6.7 3.8	2.92 2.80	13.0-14.0 12.8-13.4	1+2 2n	BD + 33°4112 (8m): $\Delta\alpha=+13^{\circ}$, $\Delta\delta=+2.$
-	GP 24	80.755	246.3	6.55	9.5-11.5	1+1	GP 24 = BD + 34°4278 (9m)
21019-059N3435-59 9.5-11.5 (1n)							
-	GP 25	80.755	299.7	4.33	11.0-13.0*	1+1	
21021-061N3435-59 11.0-12.5 (2n)							
-	GP 22	80.755	96.2	6.05	9.5-10.7	1+1	GP 22 = BD + 34°4283 (9m)
21025-066N3412-56 9.5-10.7 (3n)	AB						
9.5-10.5	AC	80.755	16.9	31.8	9.5-10.5	1+1	
-	GP 26	80.755	243.9	4.55	9.5-10.2	1+1	
21051-092N3507-31 9.8-11.4 (6n)							
14784	Σ 2783	79.660	7.9	0.71	0.0	1+2	
21114-141N5753-78 7.8-7.8		79.665 79.670 79.665	9.5 8.4 8.7	0.73 0.77 0.74	0.0 0.0 0.0	2+2 2+2 3n	The angle has decreased by 34° since 1830 with decrease in distance.
14783	HI 48	80.656	252.2	0.58	0.2	2+2	Baize, 1949: +1°.1, -0".09
21117-137N6400-25 7.1-7.3							
14894	0Σ 435	79.682	227.4	0.58	0.2	1+2	The angle has increased by 24° since 1842
21163-214N0228-54 8.1-8.6							
14889	0Σ 437	79.681	25.3	2.04	0.5	1+2	
21166-208N3202-28 6.9-7.6	AB	80.711 80.196	27.6 26.4	2.24 2.14	0.3 0.4	2+1 2n	
14919	J 851	80.711	139.9	1.46	0.5	1+2	
21179-226N1508-34 10.2-10.5	AB						
14992	A 619	79.704	61.6	0.70	8.0-8.7	2+2	Very slow angular increase.
21230-268N4202-28 8.5-9.3							

MICROMETER MEASURES OF DOUBLE STARS

ADS m	Disc 1900-2000 Mult.	Epoch 1900+	θ	ρ	m	w	Notes
15139	Hn 46	80.653	199.3	1.70	0.0	1+1	
21325-366N3557-84							
11.3-11.6							
15270	Σ 2822	79.673	296.2	2.04	1.5	1+1	
21397-441N2817-45		80.809	297.8	1.87	-	1+2	
4.7-6.1	AB	80.355	297.2	1.94	1.5	2n	Heintz, 1965: $-0^{\circ}9$, $+0.^{\circ}16$
15417	Σ 2845	80.640	172.9	1.96	0.1	1+1	
21496-524N6237-65		80.889	172.8	1.75	0.1	1+2	
8.4-8.5		80.533	172.8	1.83	0.1	2n	
-	GP 93	80.632	284.2	5.86	11.0-11.7	1+1	
22227-271N3446-76		80.653	283.5	5.11	11.0-11.5	2+1	
11-1-11.6 (4n)		80.645	283.8	5.41	11.0-11.6	2n	
15971	Σ 2909	79.673	227.8	1.78	0.1	1+1	
22237-288S0032-02		80.809	224.6	1.57	0.2	1+1	
4.4-4.6		80.889	222.4	1.56	0.2	1+2	
		80.519	224.6	1.62	0.2	3n	Harrington, 1967: $+1^{\circ}6$, $-0.^{\circ}15$
-	GP 90	80.632	284.4	6.67	9.5-9.3	1+2	
22239-284N3503-33		80.654	285.4	6.31	0.0	2+2	
9.5-9.4 (2n)		80.645	285.0	6.46	-0.1	2n	
-	GP 99	80.632	290.9	1.47	10.5-11.0	2+2	
22242-286N3443-73		80.654	289.4	1.59	10.0-10.6	1+2	
10.0-10.5 (4n)		80.641	290.2	1.52	10.3-10.9	2n	
15988	Σ 2912	80.681	117.2	0.84	1.0	2+2	
22249-299N0355-85		80.727	116.0	0.62	1.2	1+2	
5.8-7.2		80.700	116.7	0.75	1.1	2n	Knipe, 1960: $-0^{\circ}8$, $-0.^{\circ}27$
-	GP 39	79.690	93.5	0.45	0.2	3+2	
22280-325N3429-60		80.654	92.4	0.62	10.0-10.5	1+2	
10.0-10.2 (4n)		80.052	93.1	0.51	0.3	2n	There is the discordance in the measures of the position angles put it is visible the tendency of direct motion. GP 39 = BD + $34^{\circ}4710$, (9 ^m .3)
16185	Σ 2934	79.690	77.0	1.08	1.2	3+3	
22370-418N2054-85		79.692	73.6	1.02		1+1	
8.7-9.7	AB	79.701	72.4	1.01	0.8	1+1	
		80.714	74.1	1.00	9.5-10.0	2+2	
		80.727	73.3	1.00	1.0	1+1	
		80.077	74.8	1.03	1.1	5n	Heintz, 1960: $+4^{\circ}8$, $+0.^{\circ}08$
16317	Σ 2950	80.894	288.8	1.49	1.0	1+2	About 30° retrograde motion and decrease in distance since 1832.
22474-513N6109-41							
6.1-7.4	AB						
16326	A 632	80.725	170.6	0.94	1.0	1+1	Heintz, 1961: $+2^{\circ}2$, $+0.^{\circ}12$
22480-521N5712-44							
8.2-9.0	AB						
16345	β 382	79.673	203.9	0.77	2.0	1+1	
22492-537N4413-45		80.711	207.4	0.90	2.5	1+1	
5.8-7.8	AB	80.192	205.6	0.84	2.2	2n	Muller, 1953: $0^{\circ}0$, $-0.^{\circ}11$
16435	Hn 56	79.668	96.9	1.08	0.0	3+2	
22551-597N4117-49		79.941	96.0	1.04	0.0	1+2	
9.3-9.4		79.770	96.6	1.06	0.0	2n	The angle has decreased by 28° since 1881
-	GP 68	80.727	327.2	0.98	0.0	1+1	GP 68 = BD + $35^{\circ}5010$ (9 ^m .4)
23158-205N3548-80							
9.5-9.5 (1n)							
-	GP 67	80.640	312.3	2.99	9.0-13.5	1+1	
23169-217N3550-83							
9.2-12.6 (5n)							
-	GP 21	80.714	209.6	0.61	-0.3	2+3	GP 21 = BD + $29^{\circ}4929$ (9 ^m .5)
23223-272N2953-86							
9.7-9.8 (1n)							

ADS m	Disc 1900-2000 Mult.	Epoch 1900+	Θ	ρ	m	w	Notes
-	GP 3	80.714	120.5	4.25	12.0-12.0	2+1	
23224-272N2941-74							
12.6-12.6 (8n)							
16951	A 1242	79.668	323.1	0.69	0.0	2+2	Zulević, 1977: -1°8, -0".06
23380-431N1117-50							
9.6-9.6							
17020	0Σ 507	79.668	305.4	0.75	7.0-8.0	2+2	
23438-486N6420-53		79.673	305.0	0.70	0.7	1+2	
6.9-7.6	AB	79.690	307.1	0.59	0.5	2+2	
		79.701	304.6	0.75	0.7	1+2	
		79.682	305.8	0.69	0.7	4n	Zulević, 1977: +1°8, -0".05
17053	Es 1476	80.889	65.0	1.46	12.0-12.0	1+2	
23468-518N4231-64							
10.9-11.0	AB						
17063	β 728	80.793	6.7	1.17	8.5-8.7	1+2	
23471-521N4257-90		80.809	6.0	1.16	0.1	3+2	
8.7-8.7		80.889	8.8	1.15	0.2	2+2	
		80.832	7.1	1.16	0.2	3n	The angle has increased by 14° since 1878.
10	A 800	80.894	291.3	1.53	-0.1	2+2	
23577-629N4642-75		80.900	292.3	1.54	-0.1	2+2	
8.8-8.8		80.897	291.8	1.54	-0.1	2n	The component A is fainter than B.

REFERENCE

G.M.Popović, 1979: Bull. Obs. Astron., Belgrade, No 130, pp. 28
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**RESULTS OF OBSERVATIONS OF THE SUN, MERCURY AND VENUS
WITH THE BELGRADE LARGE MERIDIAN INSTRUMENT**

S. Sadžakov, M. Dačić and D. Šaletić

(Received February, 12, 1980)

Visual observations of the Sun, Mercury and Venus in right ascension and declination are being carried out with the Belgrade Large Meridian Instrument AS-KANA ($d = 190$ mm, $f = 2578$) since 1973. The instrument, the pavilion, the method of observation and the data processing are described elsewhere (Sadžakov et al. 1976).

We present here the results obtained within the scheme until June 1979.

In cases where the number of reference stars during daytime observations was unsufficient, we were constrained to let the coronograph running until the evening, when 10 to 15 fundamental stars could be observed, with declinations close to that of the members of the solar system concerned, in addition to two or three polar stars. Thus, due corrections Δm , Δn , and Δc could be introduced into corresponding formulae on the basis of these subsequently observed FK4 stars.

The ephemeris positions of the Sun and the planets are calculated in the Computing Centre of the Pulkovo Observatory (USSR) under the programme of M. Chubey.

The apparent right ascensions and declinations of the Sun and the two planets, obtained from our obser-

vations, are compared with the ephemeris positions and the results are given in Tables 1, 2 and 3.

All three Tables consist of ten columns, namely:

Column:

- I – date of observation
- II – observer and his assistant
- III – mean temperature within the dome
- IV – number of FK4 reference stars
- V – right ascension α from the observation
- VI – $(O - C)_\alpha$
- VII – declination δ from the observation
- VIII – $(O - C)_\delta$
- IX – epoch of observation in fraction of year
- X – clamp position

Code numbers of the observers: 1 – S. Sadžakov, 2 – D. Šaletić, 3 – M. Dačić.

Code numbers of the assistants: 1 – V. Protić – Benišek, 2 – V. Trajkovska and 3 – D. Božičković.

A dependence of the sign of $(O - C)$ on the air temperature during observation is apparent.

Sun

Table 1

Date of observ.	ob-serv.	$t^o C$	n	α	$(O - C)_\alpha$	δ	$(O - C)_\delta$	E_p	clamp. posit.
				h m s	0.001	o ' "	0.01	197	
1974					- 80	- 18 11 05.06	- 14	4.87	KE
14.XI	1,2,3,	12.0	8	15 16 58.138	+ 85	- 18 26 37.96	- 32	4.87	KE
14.XI	1,4	13.0	8	15 21 04.396	-	- 18 41 49.58	- 24	4.87	KE
16.XI	1,3	10.1	21	15 25 11.501	-	- 19 11 15.15	- 8	4.88	KE
18.XI	1,3	12.9	19	15 33 27.947	-	- 19 39 17.89	+ 9	4.89	KE
20.XI	1,3	9.5	32	15 41 47.740	-	- 20 05 54.74	+ 1	4.89	KW
22.XI	1,2,3	9.0	7	15 50 10.720	-	- 22 54 01.25	- 36	4.94	KW
10.XII	1,3,4	5.9	8	17 07 44.497	-	-	-	-	-

S. SADŽAKOV, M. DAČIĆ, D. ŠALETIĆ

Sun

Table 1

Date of observ.	ob- serv.	t°C	n	ε	(0-C) _a	δ	(0-C) _b	E _p	clan pos
1975									
15.I	1,3	8.5	7	19 46 20.424	+ 29	- 21 11 46.31	- 10	5.04	K
16.I	1,3	8.0	12	19 50 38.185		- 21 00 41.25	+ 13	5.04	K
21.IV	1,3	15.4	13	01 54 21.528	-135	+ 11 43 14.48	- 32	5.30	K
29.IV	1,3	15.0	12	02 24 28.704	+ 45	+ 14 20 10.43	- 17	5.33	K
16.VII	1,3	27.2	8	07 40 34.053	+ 57	+ 21 25 46.74	+ 10	5.54	K
18.VII	1,3	26.7	11	07 48 37.631	-118	+ 21 05 42.73	+ 5	5.54	K
20.VII	1,3	23.4	19	07 56 39.301		+ 20 44 14.05	+ 8	5.56	K
21.VII	1,3	23.8	11	08 00 39.233		+ 20 32 57.27	- 4	5.56	K
10.IX	1,3	18.1	14	11 12 33.354	- 7	+ 05 05 28.42	- 13	5.69	K
18.IX	1,2,3	22.8	10	11 41 15.488	- 36	+ 02 01 38.37	+ 92	5.71	F
19.IX	1,3	23.0	11	11 44 50.441	-142	+ 01 38 23.45	- 26	5.72	F
20.IX	1,3	23.4	11	11 48 25.668	- 16	+ 01 15 08.28	- 8	5.72	F
1977									
10.III	1,5	12.1	18	23 22 19.491	- 38	- 04 03 36.29	- 61	7.19	I
17.III	1,5	12.6	17	23 47 59.559	-134	- 01 18 07.02	+ 36	7.21	I
18.III	1,5	12.5	6	23 51 38.970	+130	- 00 54 24.14	- 14	7.21	I
24.III	1,5	21.5	9	00 13 30.891	+ 40	+ 01 27 42.82	+ 13	7.22	I
25.III	1,5	18.5	7	00 17 09.208		+ 01 51 19.04	+ 7	7.31	I
7.IV	1,5	15.3	7	01 04 31.101	-164	+ 06 52 01.62	+ 35	7.32	I
18.IV	1,5	10.9	17	01 45 04.251		+ 10 51 38.99	+ 89	7.32	I
19.IV	1,5	14.1	17	01 48 47.509	- 74	+ 11 12 26.93	- 12	7.32	I
21.IV	1,5	10.4	17	01 56 15.490		+ 11 53 33.85	+ 39	7.33	I
22.IV	1,5	11.0	12	02 00 00.086		+ 12 13 48.56	+ 28	7.33	I
21.VI	1,5	12.0	10	05 59 43.668		+ 23 26 18.42	- 2	7.47	I
25.VI	1,2,5	21.6	10	06 16 21.836		+ 23 23 07.17	- 2	7.47	I
27.VII	1,3	23.1	14	06 24 40.149		+ 19 11 12.62	+ 34	7.58	I
29.IX	1,3,5	19.5	32	12 22 38.700		- 02 27 03.71	+ 25	7.75	I
30.IX	1,3,5	12.6	8	12 26 15.294	-118	- 02 50 21.62	- 22	7.75	I
1.X	1,3	18.0	8	12 29 52.594	-188	- 03 13 38.33	- 7	7.75	I
4.X	1,3,5	13.0	14	12 40 45.126	-175	- 04 23 15.42	+ 17	7.76	I
5.X	3,5	16.8	19	12 44 23.602	- 39	- 04 46 22.56	+ 41	7.76	I
6.X	1,3,5	18.6	12	12 48 02.356	- 12	- 05 09 24.85	- 38	7.77	I
7.X	1,3,5	19.5	12	12 51 41.489	- 14	- 05 32 24.98	- 64	7.77	I
17.X	1,3,5	13.9	15	13 28 39.182	+113	- 09 17 26.70	- 3	7.80	I
18.X	1,3,5	14.7	26	13 32 23.836	+ 55	- 09 39 19.23	+ 27	7.80	I
19.X	1,3,5	16.7	12	13 36 09.043	- 42	- 10 01 03.31	+ 6	7.80	I
20.X	1,3,5	15.4	12	13 39 55.072	+ 77	- 10 22 37.56	- 29	7.80	I
21.X	1,3,5	15.5	12	13 43 41.716	+186	- 10 44 02.74	- 32	7.81	I
24.X	3,5	16.0	12	13 55 05.050	- 8	- 11 47 21.25	+ 14	7.81	I
27.X	3,5	15.5	12	14 06 34.862		- 12 49 01.62	- 45	7.81	I
28.X	3,5	15.9	12	14 10 26.275		- 13 09 17.00	+ 75	7.81	I
11.XI	1,3,5	15.8	16	15 05 51.608	+ 63	- 17 27 01.58	+ 49	7.87	I
21.X	1,3,5	6.5	13	15 47 07.377		- 19 56 16.94	- 51	7.87	I
1978									
5.I	1,3,5	2.0	10	19 04 01.391	+ 11	- 22 37 33.47	- 4	8.07	
6.I	1,5	-2.0	10	19 08 24.730		- 22 30 37.73	+ 31	8.07	

RESULTS OF OBSERVATIONS OF THE SUN, MERCURY AND VENUS

Table 1

Sun

Date of observ.	ob-serv.	t°C	n	α	(0-C)α	δ	(0-C)δ	E _p	clamp. posit.
12.I	1,5	5.4	10	19 34 34.332	+ 38	- 21 39 48.34	+ 32	8.07	KW
2.II	1,5	5.9	10	21 02 49.880		- 16 50 46.69	- 25	8.08	KW
3.III	1,3,5	15.6	10	22 55 28.390		- 06 52 14.34	+ 1	8.17	KW
13.III	1,3	16.0	10	23 32 27.863	- 29	- 02 58 35.90	- 6	8.17	KW
15.III	1,3	13.2	10	23 39 47.759		- 02 11 16.96	+ 43	8.17	KW
23.III	1,3	7.8	20	00 08 58.765	- 4	+ 00 58 17.09	+ 67	8.24	KW
28.III	1,3,5	10.0	16	00 27 09.805	- 77	+ 02 56 00.48	- 10	8.25	KW
30.III	3,5,6	17.9	16	00 34 27.435	- 37	+ 03 42 45.04	+ 19	8.25	KW
31.III	1,3,5	15.2	16	00 38 04.925	+ 37	+ 04 06 01.27	+ 25	8.25	KW
3.IV	1,3,5	20.5	16	00 49 00.773	- 69	+ 05 15 21.40	+ 11	8.25	KW
7.IV	1,3	21.0	18	01 03 37.743	+ 16	+ 06 46 30.18	- 18	8.28	KW
4.V	3,5	17.0	18	02 44 35.543		+ 15 55 17.48	+ 16	8.29	KW
5.V	1,3,5	21.0	18	02 48 26.598	- 54	+ 16 12 34.20	+ 9	8.29	KW
17.V	3,5	17.1	14	03 35 24.046		+ 19 17 35.89	- 50	8.38	KW
18.V	1,3,5	17.1	4	03 39 22.430		+ 19 31 00.29	- 11	8.38	KW
20.VI	1,3,5	22.1	12	05 54 32.282	+159	+ 23 25 57.98	+ 72	8.47	KW
23.VI	1,3	22.7	11	06 07 01.091	+ 22	+ 23 25 43.62	+ 9	8.48	KW
24.VI	3,5	23.9	11	06 11 10.518		+ 23 24 49.56	+ 23	8.48	KW
25.VI	1,3	24.1	11	06 15 20.036	+148	+ 23 23 30.84	+ 48	8.48	KW
5.VII	3,5	26.3	12	06 56 44.358		+ 22 47 51.68	- 6	8.53	KW
11.VII	1,3,5	22.4	12	07 21 20.469	- 11	+ 22 07 20.31	- 19	8.53	KW
12.VII	1,3	25.2	14	07 25 24.890	-113	+ 21 59 14.69	- 21	8.53	KE
13.VII	1,3	24.4	11	07 29 29.047	+ 19	+ 21 50 47.13	+ 27	8.54	KE
14.VII	1,3,5	24.1	11	07 33 32.692	+124	+ 21 41 55.06	- 15	8.54	KE
18.VII	1,3	26.0	11	07 49 41.448	-100	+ 21 02 51.39	+ 15	8.55	KE
19.VII	1,3	27.8	21	07 53 42.447	*	+ 20 52 11.74	+ 11	8.55	KE
21.VII	1,3,5	23.1	14	08 01 42.742	+162	+ 20 29 48.58	- 37	8.56	KE
26.VII	1,3	22.2	13	08 21 32.959	- 76	+ 19 27 54.73	0	8.58	KW
27.VII	1,3	23.7	13	08 25 29.508	+ 96	+ 19 14 32.76	- 23	8.58	KW
31.VII	1,3	25.3	18	08 41 09.275	+177	+ 18 17 55.78	+ 5	8.62	KW
1.VIII	1,3	25.2	17	08 45 02.377	-128	+ 18 03 00.11	0	8.65	KW
2.VIII	1,3	25.8	11	08 48 55.378	+ 42	+ 17 47 47.26	- 42	8.66	KW
3.VIII	1,3	25.6	11	08 52 47.563	+ 1	+ 17 32 18.37	+ 3	8.68	KW
7.VIII	1,3	27.8	22	09 08 10.520	+119	+ 16 27 30.25	+ 36	8.78	KW
14.VIII	1,3	21.4	22	09 34 42.207	-100	+ 14 24 08.47	+ 65	8.78	KW
15.VIII	1,2,3	22.8	7	09 38 27.487	+ 73	+ 14 05 32.25	+ 21	8.79	KW
22.VIII	1,3	20.8	7	10 04 28.634	+ 30	+ 11 49 32.90	+ 19	8.80	KE
23.VIII	1,3	23.4	4	10 08 09.716	- 34	+ 11 29 18.74	+ 8	8.80	KE
4.IX	1,3,5	20.2	5	10 51 54.623		+ 07 13 54.80	- 31	8.81	KE
25.IX	1,3	21.8	6	12 07 20.941	- 8	- 00 47 55.04	+ 47	8.84	KE
26.IX	1,3	20.8	10	12 10 56.789	- 57	- 01 11 16.29	- 2	8.84	KE
9.X	1,3	18.9	13	12 56 08.009		- 06 12 40.01	+ 3	8.86	KW
10.X	1,3,5	18.6	13	13 01 48.272	- 11	- 06 35 27.41	+ 11	8.87	KW
15.X	1,3	15.0	13	13 20 16.763	+ 59	- 08 27 51.42	- 102	8.88	KW
18.X	1,3	16.9	13	13 31 28.257	+ 38	- 09 33 56.52	- 37	8.88	KW
19.X	2,3,5	15.9	12	13 35 13.270		- 09 55 51.48	- 32	8.89	KW
24.X	1,3,5	11.6	12	13 54 08.696	+245	- 11 42 11.99	+ 14	8.89	LW
25.X	1,3,5	11.5	12	13 57 57.584		- 12 02 00.62	+ 79	8.90	KW

Sun

Table 1

Date of observ.	ob-serv.	t°C	n	ε	(0-C)α	δ	(0-C)δ	E _p	clamp. posit.
1979									
21.II	1,2,3	2.4	14	22 16 46.975		- 10 41 21.72	- 2	9.14	KW
22.II	1,2	3.0	13	22 20 36.448	+ 7	- 10 19 37.98	+ 33	9.15	KW
23.II	1,2,3	-0.6	16	22 24 25.411	+121	- 09 57 44.55	+ 57	9.15	KW
27.II	1,3	2.2	16	22 39 34.753		- 08 28 42.89	+ 11	9.16	KW
2.III	1,3	8.3	19	22 50 50.963		- 07 20 37.25	+ 56	9.17	KW
7.III	1,2,3	7.3	13	23 09 28.052	+ 58	- 05 25 07.29	- 52	9.18	KW
11.III	1,3	6.3	13	23 24 14.137	+162	- 03 51 24.67	- 80	9.19	KW
20.III	1,3	15.3	12	23 57 10.399	+ 26	+ 00 18 28.07	- 35	9.21	KW
21.III	3,5	15.0	8	00 00 48.973	- 98	+ 00 05 13.84		9.22	KW
22.III	1,3	15.2	12	00 04 27.617	- 53	+ 00 28 55.97	+ 77	9.23	KW
25.III	1,2,3	10.0	6	00 15 23.115	+ 55	+ 01 39 51.19	- 21	9.23	KW
4.IV	1,3	13.5	12	00 51 48.300	- 70	+ 05 32 54.42	+ 26	9.25	KW
12.IV	1,3	11.6	18	01 21 06.035	+ 7	+ 08 32 34.87	+ 6	9.28	KW
16.IV	1,3	17.0	10	01 35 51.399		+ 09 59 09.54	+ 20	9.29	KW
19.IV	1,3	7.5	15	01 46 59.329		+ 11 02 22.86	- 87	9.30	KW
20.IV	1,2,3	7.2	15	01 50 42.803		+ 11 23 07.78	- 24	9.31	KW
23.IV	1,3	14.8	12	02 01 55.874	- 18	+ 12 24 09.82	+ 10	9.31	KW
24.IV	1,3,5	18.3	14	02 05 41.128	- 55	+ 12 44 06.34	- 8	9.32	KW
3.V	1,3	18.4	14	02 39 51.066	- 42	+ 15 33 34.20	- 32	9.34	KW
15.V	1,3	16.4	10	03 26 31.730	- 89	+ 18 46 24.61	- 4	9.36	KW
17.V	1,3	18.4	12	03 34 26.633		+ 19 14 18.85	+ 6	9.37	KW
18.V	3,5	18.0	10	03 38 24.912		+ 19 27 49.43	- 77	9.37	KW
20.V	1,3	22.9	12	03 46 23.192		+ 19 53 46.88	+ 61	9.38	KW
21.V	1,3	26.5	14	03 50 23.183		+ 20 06 14.16	- 40	9.39	KW
23.V	1,3	24.1	12	03 58 24.607	-213	+ 20 30 08.57	- 55	9.39	KW
24.V	1,3	23.4	12	04 02 26.332	-122	+ 20 41 35.34	+ 4	9.40	KW
25.V	1,3	24.3	12	04 06 28.459	-145	+ 20 52 40.53	- 29	9.40	KW
26.V	1,3	24.6	10	04 10 31.260		+ 21 03 23.43	- 20	9.40	KW
27.V	1,3	25.7	10	04 14 34.451	+ 46	+ 21 13 45.60	+ 2	9.40	KW
29.V	1,3,5	18.6	12	04 22 42.092		+ 21 33 21.94	- 15	9.42	KW
30.V	1,3	21.7	12	04 26 46.578	- 21	+ 21 42 38.15	+ 47	9.43	KW
31.V	1,3	23.2	12	04 30 51.580	+ 53	+ 21 51 29.86	- 20	9.43	KW
1.VI	1,3	25.0	12	04 34 56.844	- 14	+ 21 59 58.63	- 24	9.43	KW
4.VI	1,3	26.0	12	04 47 15.074	- 26	+ 22 23 08.03	- 14	9.44	KW
5.VI	1,3	24.7	12	04 51 21.799	- 76	+ 22 30 04.10	- 13	9.45	KW
6.VI	1,3	25.0	12	04 55 28.969		+ 22 36 36.23	- 55	9.45	KW
7.VI	1,3	24.4	12	04 59 36.310	- 55	+ 22 42 45.33	- 24	9.46	KW
8.VI	1,3	24.5	12	05 03 44.048		+ 22 48 30.75	- 20	9.46	KW
11.VI	1,3	24.9	12	05 16 08.719	+ 69	+ 23 03 19.81	- 85	9.47	KW

Mercury

Table 2

Date of observ.	ob-serv.	t°C	n	ε	(0-C)α	δ	(0-C)δ	E _p	clamp. posit.
1974									
14.XI	1,2,3	12.0	8	h m s	0.001	o ' "	0°01	197	
				14 07 14.228	+ 21	- 10 32 48.86	+ 21	4.87	KE
16.XI	1,4	13.0	8	14 17 20.982	+ 21	- 11 33 11.21	+ 32	4.87	KE
10.XII	1,3,4	5.9	8	16 44 55.755	+ 21	- 22 51 30.59	- 17	4.94	KW

RESULTS OF OBSERVATIONS OF THE SUN, MERCURY AND VENUS

Mercury

Table 2

Date of observ.	ob-serv.	t°C	n	a	(0-C)α	δ	(0-C)δ	E _p	clump. posit.
1975									
15.I	1,3	8.5	7	20 53 53.030	- 79	- 19 05 57.66	- 10	5.04	KW
21.IV	1,3	15.4	13	02 05 56.397	-109	+ 12 36 11.18	- 45	5.30	KE
29.IV	1,3	15.0	12	03 10 42.693	- 48	+ 19 05 14.54	+ 18	5.33	KE
16.VII	1,3	27.2	8	06 27 55.455	- 79	+ 22 41 37.09	+ 22	5.54	KW
18.VII	1,3	26.7	11	06 43 28.585	- 79	+ 22 53 35.64	- 21	5.54	KW
1977									
21.VI	1,5	12.0	10	05 14 50.656	+ 53	+ 22 55 25.05	+ 3	7.47	KE
29.IX	1,3,5	19.5	32	11 31 27.239	+ 53	+ 05 04 02.59	- 24	7.75	KE
30.IX	1,3,5	12.6	8	11 37 39.330	+101	+ 04 26 49.90	+ 25	7.75	KE
1.X	1,3	18.0	8	11 43 56.795	+103	+ 03 47 50.37	+ 6	7.75	KE
4.X	1,3,5	13.0	14	12 03 08.772	- 79	+ 01 42 46.47	- 25	7.76	KE
5.X	3,5	16.8	19	12 09 36.275	+165	+ 00 59.07.30	- 58	7.76	KE
6.X	1,3,5	18.6	12	12 16 04.711	+ 3	+ 00 14 47.34	+ 48	7.77	KE
7.X	1,3,5	19.5	12	12 22 31.129	+ 22	+ 00 30 02.15	+ 13	7.77	KW
1978									
5.I	1,3,5	2.0	10	17 27 55.451	- 1	- 20 43 56.62	- 13	8.07	KW
12.I	1,5	5.4	10	17 53 08.108	- 53	- 21 58 23.55	- 18	8.07	KW
2.II	1,5	5.9	10	19 55 21.031	-136	- 22 06 44.73	- 26	8.08	KE
23.III	1,3	7.8	20	01 04 39.166	+ 4	+ 10 12 20.00	- 43	8.24	KE
29.III	1,3	5.6	16	01 29 58.311	- 53	+ 12 43 33.13	- 38	8.25	KE
20.VI	1,3,5	22.1	12	06 26 35.395	-201	+ 24 56 30.49	- 17	8.47	KW
23.VI	1,3	22.7	11	06 54 07.042	+ 45	+ 24 42 11.49	- 9	8.48	KW
25.VI	1,3	24.1	11	07 11 40.948	-316	+ 24 20 20.25	- 10	8.48	KW
5.VII	3,5	26.3	12	08 28 00.754	-149	+ 20 42 12.57	+ 6	8.53	KW
12.VII	1,3	25.2	14	09 09 38.442	+ 16	+ 17 07 46.11	- 12	8.53	KE
13.VII	1,3	24.4	11	09 14 48.762	- 53	+ 16 35 30.88	- 6	8.54	KE
18.VII	1,3	26.0	11	09 37 46.289	+118	+ 13 53 55.05	+ 14	8.55	KE
19.VII	1,3	27.8	21	09 41 46.182	- 53	+ 13 22 09.35	- 11	8.55	KE
25.IX	1,3	21.8	6	11 54 09.402	- 53	+ 02 27 49.68	+ 4	8.84	KE
9.X	1,3	18.9	13	13 22 50.736	+ 90	- 08 15 56.94	- 7	8.86	KW
1979									
21.II	1,2,3	2.4	14	22 56 14.119	+016	- 07 58 34.88	+ 1	9.14	KW
23.II	1,2,3	-0.6	16	23 09 28.605	-205	- 06 16 31.66	- 16	9.15	KW
23.IV	1,3	14.8	12	00 24 01.220	- 33	- 00 15 32.38	- 18	9.31	KW
24.IV	1,3,5	18.3	14	00 28 07.812	+ 29	00 06 39.50	- 24	9.32	KW
3.V	1,3	18.4	14	01 11 07.362	+ 47	04 29 04.23	- 14	9.34	KW
20.V	1,3	22.9	12	03 08 57.187	+ 3	16 02 12.55	- 10	9.38	KW
21.V	1,3	26.5	14	03 09 54.519	- 51	16 45 08.15	- 52	9.39	KW
23.V	1,3	24.1	12	03 26 21.147	+ 29	18 09 17.75	- 12	9.39	KW
24.V	1,3	23.4	12	03 34 50.094	+138	18 50 09.74	- 8	9.40	KW
25.V	1,3	24.3	12	03 43 28.868	+133	19 29 57.38	- 51	9.40	KW
26.V	1,3	24.6	10	03 52 16.908	+ 35	20 08 28.62	+ 20	9.40	KW
4.VI	1,3	26.0	12	05 15 53.940	+ 26	24 25 45.69	- 1	9.44	KW
5.VI	1,3	24.7	12	05 25 16.405	+ 45	24 41 42.18	+ 36	9.45	KW
6.VI	1,3	25.0	12	05 34 34.511	+ 16	24 54 51.22	+ 12	9.45	KW

Venus

Table 3

Date of observ.	ob-serv.	t°C	n	α	(0-C)α	δ	(0-C)δ	E _p	clamp. posit.
1974									
14.XI	1,2,3	12.0	8	15 25 33.755	- 0001	o' "	001	197	
15.XI	1,4	13.0	8	15 30 39.634	- 3	- 18 33 09.71	+ 24	4.87	KE
16.XI	1,3	10.1	21	15 35 46.746	- 5	- 18 53 17.96	+ 1	4.87	KE
20.XI	1,3	9.5	32	15 56 27.377	- 6	- 20 08 32.69	0	4.89	KE
10.XII	1,3,4	5.9	8	17 43 57.951	- 4	- 23 55 17.94	+ 54	4.94	KE
1975									
15.I	1,3	8.5	7	20 57 49.125	- 17	- 18 45 42.39	+ 21	197	KW
21.IV	1,3	15.4	13	04 27 51.302	+ 243	+ 23 27 00.68	+ 32	5.30	KE
29.IV	1,3	15.0	12	05 08 06.995	+ 4	+ 24 56 10.53	- 40	5.33	KE
16.VII	1,3	27.2	8	10 25 05.817	- 57	+ 08 52 23.22	- 41	5.54	KW
18.VII	1,3	26.7	11	10 29 02.728	+ 118	+ 08 09 12.32	+ 15	5.54	KW
20.VII	1,3	23.4	19	10 32 37.355	+ 28	+ 07 26 56.95	+ 5	5.54	KE
21.VII	1,3	23.8	11	10 34 16.071	+ 28	+ 07 06 13.12	- 8	5.56	KE
10.IX	1,3	18.1	14	09 44 40.219	+ 7	+ 05 16 02.54	+ 30	5.69	KW
19.IX	1,3	23.0	11	09 42 11.773	- 93	+ 06 57 25.38	+ 27	5.72	KE
20.IX	1,3	23.4	11	09 42 41.505	+ 19	+ 07 05 59.32	+ 8	5.72	KW
1977									
17.III	1,5	12.6	17	01 19 42.121	+ 134	+ 16 16 28.15	- 35	197	KE
18.III	1,5	12.5	6	01 19 13.814	- 129	+ 16 20 54.60	+ 14	7.21	KW
24.III	1,5	21.4	9	01 13 24.017	- 25	+ 16 20 54.74	- 13	7.22	KW
25.III	1,5	18.4	7	01 11 57.034	- 26	+ 16 13 21.17	- 8	7.31	KW
7.IV	1,5	15.3	7	00 45 52.736	+ 233	+ 12 55 53.61	- 36	7.32	KW
19.IV	1,5	14.1	17	00 27 04.922	+ 74	+ 08 22 56.89	+ 13	7.32	KW
21.IV	2,5	10.4	17	00 25 45.136	- 25	07 43 53.14	- 39	7.33	KW
22.IV	1,5	11.0	12	00 25 18.574	- 25	07 25 42.20	- 28	7.33	KW
25.VI	1,2,5	21.6	10	00 24 51.876	- 25	13 21 14.77	+ 2	7.47	KE
30.IX	1,3,5	12.6	8	10 45 46.796	- 25	09 05 53.11	- 3	7.75	KE
1.X	1,3	18.0	8	10 50 24.866	- 66	08 39 47.51	+ 2	7.75	KE
4.X	1,3,5	13.0	14	11 04 16.704	+ 117	07 19 59.81	+ 5	7.76	KE
5.X	3,5	16.8	19	11 08 52.828	- 141	06 52 55.20	+ 17	7.76	KE
6.X	1,3,5	18.6	12	11 13 28.901	- 76	06 25 37.33	- 10	7.77	KE
7.X	1,3,5	19.5	11	11 18 04.628	- 8	05 58 08.60	+ 52	7.77	KW
17.X	1,3,5	13.9	15	12 03 48.375	- 160	01 14 28.14	+ 4	7.80	KW
18.X	1,3,5	14.7	26	12 08 22.331	- 15	00 45 30.59	- 27	7.80	KW
19.X	1,3,5	16.7	12	12 12 56.228	+ 42	- 00 16 28.04	- 6	7.80	KW
20.X	1,3,5	15.4	12	12 17 29.902	- 187	- 00 12 38.09	+ 28	7.80	KW
21.X	1,3,5	15.4	12	12 22 04.056	- 40	- 00 41 44.94	+ 32	7.81	KW
24.X	3,5	16.0	12	12 35 47.056	- 71	- 02 09 12.99	- 15	7.81	KW
11.XI	1,3,5	15.8	16	13 59 31.368	- 107	- 10 40 11.27	0	7.87	KW
21.XI	1,3,5	6.5	13	14 47 55.841	- 25	- 14 55 00.94	+ 50	7.87	KW
1978									
05.I	1,3,5	+2.0	10	18 46 45.412	- 10	- 23 28 26.42	+ 17	197	KW
06.I	1,5	-2.0	10	18 52 14.577	- 9	- 23 28 26.27	- 31	8.07	KW
12.I	1,5	+5.4	10	19 24 59.093	- 38	- 22 43 55.92	- 14	8.07	KW
02.II	1,5	5.9	10	21 15 12.272	- 129	- 17 17 23.56	+ 50	8.08	KW

RESULTS OF OBSERVATIONS OF THE SUN, MERCURY AND VENUS

Venus

Table 3

Date of observ.	ob-serv.	t°C	n	α	(0-C)α	δ	(0-C)δ	E_p	clamp. posit.
03.III	1,3,5	15.6	10	23 33 45.640	- 9	- 04 20 32.13	- 1	8.17	KW
13.III	1,3	16.0	10	00 19 15.604	+ 28	00 45 57.32	+ 6	8.17	KW
23.III	1,3	7.8	20	00 08 58.765	- 4	05 51 48.61	- 24	8.24	KW
28.III	1,3,5	10.0	16	01 27 31.697	+ 78	08 20 41.28	+ 9	8.25	KW
29.III	1,3	5.6	16	01 32 07.662	+ 22	08 49 56.67	+ 38	8.25	KW
30.III	3,5,6	17.8	16	01 36 44.300	+ 37	09 18 59.51	- 10	8.25	KW
31.III	1,3,5	15.8	16	01 41 21.488	- 37	09 47 49.79	- 25	8.25	KW
03.IV	1,3,5	20.5	16	01 55 17.562	+ 68	11 12 56.22	- 10	8.25	KW
07.IV	1,3	21.0	18	02 14 03.051	- 16	13 02 40.27	+ 18	8.28	KW
05.V	1,3,5	21.0	18	04 32 45.629	+ 54	22 55 54.80	- 9	8.29	KW
20.VI	1,3,5	22.1	12	08 31 35.521	+ 61	20 53 06.95	- 54	8.47	KW
23.VI	1,3	22.7	11	08 46 05.021	- 68	19 58 28.76	0	8.48	KW
25.VI	1,3	24.1	11	08 55 37.054	- 58	19" 19 29.53	- 38	8.48	KW
11.VII	1,3,5	22.4	12	10 08 12.454	+ 11	13 05 53.20	+ 20	8.53	KW
12.VII	1,3	25.2	14	10 12 32.175	+142	12 39 32.68	+ 32	8.53	KE
13.VII	1,3	24.4	11	10 16 50.070	- 20	12 12 54.72	- 21	8.54	KE
14.VII	1,3,5	24.1	11	10 21 06.606	-127	11 46 00.29	+ 15	8.54	KE
18.VII	1,3	26.0	11	10 37 59.548	+ 53	09 55 55.98	- 32	8.55	KE
21.VII	1,3,5	23.1	14	10 50 25.007	- 97	08 31 08.13	+ 37	8.56	KE
26.VII	1,3	22.2	13	11 10 43.159	+ 76	06 06 21.74	+ 1	8.58	KW
27.VII	1,3	23.7	13	11 14 43.116	- 95	19 37 00.10	+ 24	8.58	KW
31.VII	1,3	25.3	18	11 30 32.779	-178	03 38 28.24	- 5	8.62	KW
01.VIII	1,3	25.2	17	11 34 27.928	+128	03 08 37.39	+ 1	8.65	KW
02.VIII	1,3	25.8	11	11 38 21.643	- 9	02 38 44.12	+ 42	8.66	KW
03.VIII	1,3	25.6	11	11 42 14.499	- 1	02 08 47.56	- 2	8.68	KW
07.VIII	1,3	27.8	22	11 57 36.247	-119	00 08 41.24	- 36	8.78	KW
14.VIII	1,3	21.4	22	12 23 54.464	+ 99	- 03 02 58.34	- 64	8.78	KW
15.VIII	1,2,3	22.8	7	12 27 36.142	- 73	- 03 50 43.07	- 21	8.79	KW
22.VIII	1,3	20.8	7	12 53 04.141	- 31	- 07 15 56.89	- 19	8.80	KE
23.VIII	1,3	23.4	4	12 56 38.906	+ 35	- 07 44 43.26	- 9	8.80	KE
25.IX	1,3	21.8	6	14 41 32.435	+ 8	- 21 03 41.17	- 51	8.84	KE
26.IX	1,3	20.8	10	14 44 01.591	- 2	- 21 21 13.64	+ 2	8.84	KE
09.X	1,3	18.9	13	15 08 41.676	+101	- 24 15 19.93	+ 4	8.86	KW
10.X	1,3,5	18.6	13	15 09 50.876	+ 12	- 24 23 52.97	- 10	8.87	KW
15.X	1,3	15.0	13	15 13 35.294	- 59	- 24 53 40.92	+102	8.88	KW
18.X	1,3	16.8	13	15 14 04.283	- 38	- 24 59 58.02	- 36	8.88	KW
24.X	1,3,5	11.6	12	15 10 49.332	-145	- 24 42 08.79	- 15	8.89	KW
1979									
22.II	1,2	3.0	13	19 21 48.694	- 3	- 20 25 59.55	- 32	9.13	KW
23.II	1,2,3	-0.6	16	19 26 42.795	-121	- 20 20 14.45	- 40	9.15	KW
27.II	1,3	2.2	16	19 46 20.295	+ 25	- 19 51 39.68	- 11	9.16	KW
2.III	1,3	8.3	19	20 01 02.353	+ 25	- 19 24 22.51	- 57	9.17	KW
7.III	1,2,3	7.3	13	20 25 27.058	+ 30	- 18 15 12.01	+ 53	9.18	KW
9.III	1,3	5.3	13	20 35 10.310	+166	- 18 01 47.98	- 66	9.19	KW
15.III	1,3	14.4	15	21 04 07.094	+ 39	- 16 31 58.75	- 21	9.20	KW
20.III	1,3	15.3	12	21 27 57.946	+ 42	- 15 02 16.06	+ 34	9.21	KW
21.III	3,5	13.9	20	21 32 42.119	+ 19	- 14 43 11.54	- 7	9.22	KW
22.III	1,3	15.2	12	21 37 25.737	+122	- 14 23 41.40	- 77	9.23	KW

Venus

Table 3

Date of observ.	ob- serv.	$t^{\circ}\text{C}$	α	ϵ	$(0-\text{C})_{\alpha}$	δ	$(0-\text{C})_{\delta}$	E_p	clamp. posit.
25.III	1,2,3	10.0	6	21 51 31.988	+ 25	- 13 22 48.99	+ 22	9.23	KW
4.IV	1,3	13.5	12	22 37 48.187	+195	- 09 36 33.78	- 25	9.25	KW
12.IV	1,3	11.6	18	23 14 04.498	- 6	- 06 15 02.84	+ 14	9.28	KW
16.IV	1,3	17.0	10	23 32 01.003	+ 25	- 04 29 16.87	- 21	9.29	KW
19.IV	1,3	7.5	15	23 45 26.669	+ 25	- 03 08 21.94	+ 65	8.30	KW
20.IV	1,2,3	7.2	15	23 49 54.467	+ 25	- 02 41 08.33	+ 25	9.31	KW
23.IV	1,3	14.8	12	00 03 16.927	+ 52	- 01 18 50.27	+ 9	9.31	KW
24.IV	1,3,5	18.3	14	00 07 44.141	+ 26	- 00 51 15.49	+ 31	9.32	KW
3.V	1,3	18.4	14	00 47 50.356	- 4	+ 03 18 28.36	- 28	9.34	KW
15.V	1,3	16.4	10	01 41 53.857	+ 88	08 45 08.40	+ 5	9.36	KW
17.V	1,3	18.4	12	01 51 02.048	- 61	09 37 40.66	- 5	9.37	KW
18.V	3,5	18.0	10	01 55 37.379	+ 25	10 03 39.21	+ 77	9.37	KW
20.V	1,3	22.9	12	02 04 50.212	- 3	10 54 59.16	- 63	9.38	KW
21.V	1,3	26.5	14	02 09 27.900	+ 52	11 20 19.92	+ 92	9.39	KW
23.V	1,3	24.1	12	02 18 45.865	+190	12 10 13.40	+ 67	9.39,	KW
24.V	1,3	23.4	12	02 23 25.904	- 15	12 34 46.45	+ 7	9.40	KW
25.V	1,3	24.3	12	02 28 07.094	+ 13	12 59 04.47	+ 83	9.40	KW
26.V	1,3	24.6	10	02 32 49.185	+ 5	13 23 02.38	0	9.40	KW
27.V	1,3	25.6	10	02 37 32.189	- 47	13 46 43.82	- 2	9.40	KW
29.V	1,3,5	18.6	12	02 47 01.292	+ 25	14 33 08.32	+ 16	9.42	KW
30.V	1,3	21.7	12	02 51 47.347	+ 21	14 55 50.55	- 47	9.43	KW
31.V	1,3	23.2	12	02 56 34.334	- 53	15 18 12.45	+ 20	9.43	KW
1.VI	1,3	25.0	12	03 01 22.504	+ 15	15 40 11.25	+ 24	9.43	KW
4.VI	1,3	26.0	12	03 15 53.115	- 65	16 43 51.48	+ 16	9.44	KW
5.VI	1,3	24.7	12	03 20 45.608	+ 30	17 04 16.13	- 23	9.45	KW
6.VI	1,3	25.0	12	03 25 39.076	+ 25	17 24 16.75	+ 45	9.45	KW
7.VI	1,3	24.4	12	03 30 33.737	+ 55	17 43 50.67	+ 24	9.46	KW
8.VI	1,3	24.5	12	03 35 29.438	+ 36	18 02 58.55	+ 19	9.46	KW
11.VI	1,3	24.9	12	03 50 23.202	- 70	18 57 35.14	+ 86	9.47	KW

**OBSERVATIONS À LA LUNETTE ZENITHALE (DE 110 mm) DU
SERVICE DE LATITUDE DE L'OBSERVATOIRE DE BELGRADE
EN 1980**

R. Grujić et M. Djokić

(Reçu le 31 mai, 1981)

RÉSUMÉ: On présente les valeurs de latitude ainsi que quelques données météorologiques prises au cours d'observations.

Tableau I
Les valeurs de latitude ainsi que quelques données
météorologiques au cours d'observation

DATE	τ	Obs.	T_z	T_i	T_v	B_o	GR	φ_a	φ_b	φ_d
								44° 48' +		
1980										
I	31	1980.084 RG	8°.2C	5°.3C	9°.6C	735.8	II	—	10°.236	10°.236
II	7	.103 RG	7.1	6.4	6.4	737.6	II	10°.406	10.400	10.403
	16	.128 RG	-0.1	0.6	0.1	741.8	II	10.337	10.398	10.368
	20	.138 MD	-2.1	-0.3	-1.3	750.3	II	10.372	10.349	10.360
III	2	.169 RG	4.8	3.1	3.5	737.4	II	—	10.366	
		.169 RG	5.1	2.7	3.7	734.9	III	10.021	—	10.194
	5	.177 MD	0.1	0.8	0.1	746.1	II	—	10.226	10.226
	13	.199 RG	3.9	4.8	4.3	740.4	II	—	10.317	
		.199 RG	2.0	2.6	2.0	740.5	III	10.259	—	10.288
	18	.213 RG	3.4	3.9	3.2	737.3	III	10.227	10.145	10.186
	27	.237 RG	9.6	9.2	8.6	736.3	III	10.274	10.238	10.256
IV	13	.284 RG	7.5	8.0	7.3	746.7	III	—	10.208	10.208
	14	.286 MD	11.3	10.1	9.5	745.7	III	10.311	—	10.311
	17	.295 RG	15.6	14.4	13.6	738.2	III	10.321	10.286	10.304
V	21	.388 MD	11.3	12.1	11.6	736.6	IV	10.111	10.163	10.137
	25	.399 RG	9.6	10.5	10.0	739.0	IV	10.146	10.308	10.227
VI	10	.443 RG	21.0	20.1	19.8	735.1	IV	—	10.121	10.121
	29	.495 RG	16.4	18.6	17.2	733.2	V	10.332	—	10.332
VII	8	.521 RG	22.4	22.4	21.6	735.7	V	10.245	10.180	10.212
	14	.536 RG	18.7	20.3	19.2	739.2	V	10.077	10.178	10.128
	16	.541 RG	23.8	23.8	23.1	737.0	V	10.112	10.233	10.172
	28	.574 RG	19.4	21.8	20.8	737.4	V	10.291	10.334	10.312
VIII	3	.590 RG	22.8	23.2	22.5	742.5	V	—	10.188	10.188
	11	.612 MD	18.4	20.9	19.6	738.6	V	10.087	10.153	10.120
	26	.653 RG	14.9	16.5	15.5	743.5	V	—	10.227	
		.654 RG	14.2	14.9	14.6	743.3	VI	10.338	—	10.282
IX	4	.678 RG	13.9	14.5	14.0	744.5	VI	10.349	—	10.349
	9	.692 RG	17.8	19.3	18.4	741.6	V	—	10.227	
		.692 RG	16.2	17.5	16.5	741.4	VI	10.313	—	10.270
	21	.725 MD	17.3	16.9	16.2	741.3	VI	10.262	10.290	10.276
X	28	.826 RG	10.0	9.6	9.2	745.9	VI	10.344	10.400	10.372
	29	.829 MD	12.7	11.3	11.5	736.8	I	10.155	10.220	10.188
XI	17	.880 RG	11.8	8.8	9.3	744.9	VI	10.358	10.332	
		.881 RG	11.8	8.2	9.1	745.0	I	10.324	—	10.338
	20	.888 RG	3.2	4.4	3.5	748.9	VI	—	10.539	
		.889 RG	2.4	3.1	2.4	749.6	I	10.538	—	10.538
XII	23	.980 RG	0.4	1.3	0.8	744.2	II	10.428	10.299	10.364
	30	.998 RG	-1.4	-0.6	-1.0	748.0	I	—	10.182	10.182

OBSERVATIONS A LA LUNETTE ZENITHALE

Les valeurs observées de φ (Tableau I) sont réduites à la manière déjà signalée (Ševarlić, B., Teleki, G.) mais sans tenir compte des erreurs progressives et périodiques et du coefficient de température (Milovanović, V. et les autres). Les réductions ont été faites dans le système FK4 et on a appliqué les corrections des déclinaisons présentées dans le Tableau 2 (Grujić, R. et les autres).

La valeur du tour de la vis micrométrique adoptée était: $R = 40^{\circ}0481$.

Tv	: Température de l'air dans la salle d'observation (valeur moy. des lectures des thermomètres sud et nord).
Bo	: Lecture du baromètre en mm Hg (tenant compte de la température de baromètre).
GR.	: Numéro de la groupe.
φ_{apb}	: La latitude de la sous-groupe <u>a</u> , resp. <u>b</u> .
φ_d	: La valeur moy. de la latitude de la nuit.

LA LÉGENDE:

Date	: Année, mois et date d'observation.
τ	: Partie d'année tropique.
Obs.	: Observateurs R. Grujić (RG), M. Đokić (MD).
Tz	: Température à l'abri météorologique éloigné de 50 m de l'instrument.
Ti	: Température de l'instrument.

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TABLE DES MATIERES

Pero M.Djurković	1
M.Mijatov and Dj.Bozhichkovich	522.982
	3
S.Sadžakov, M. Dačić, D.Šaletić and B.Ševarlić	521.27
	7
G.M.Popović	521.31
	13
D.J.Zulević	521.31
	17
V.Erceg	521.31
	22
D.J. Zulević	523.84
	25
G.M.Popović	523.84
	34
S.Sadžakov, M.Dačić and D.Šaletić	523.4
	45
R.Grujić et M.Djokić	522.7
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